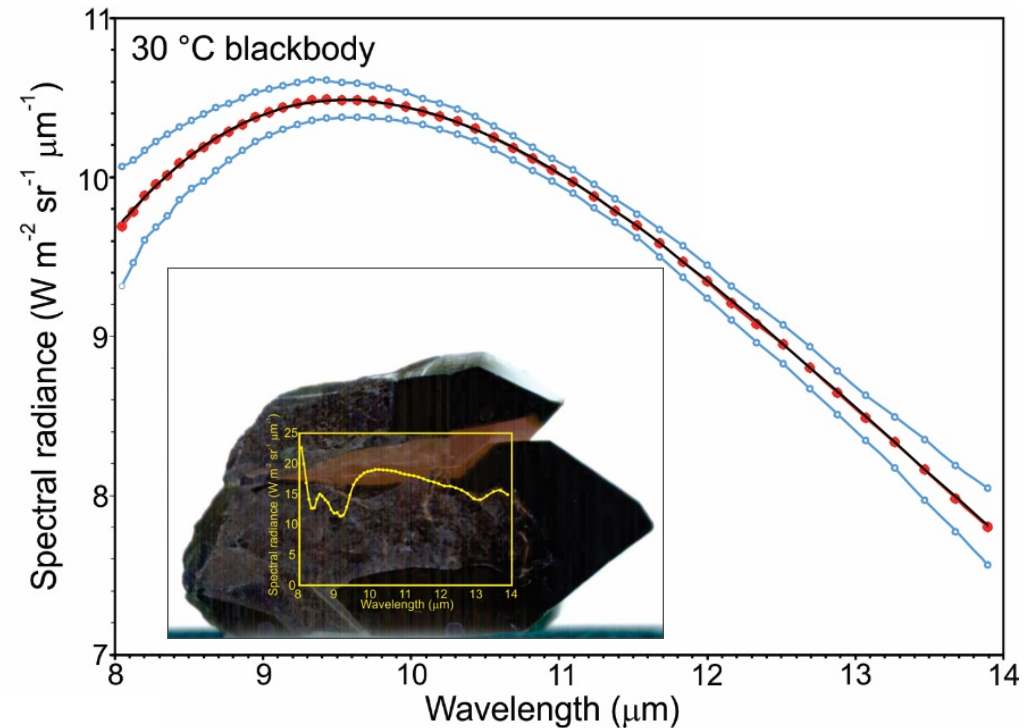
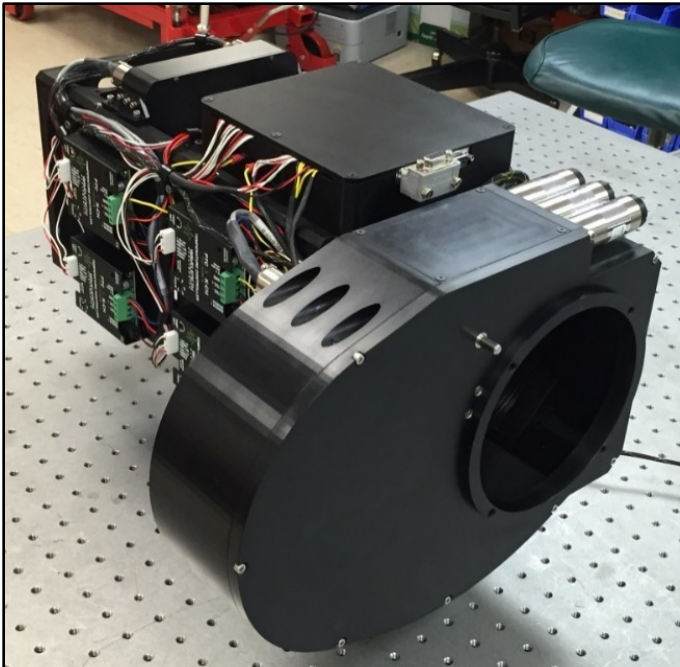




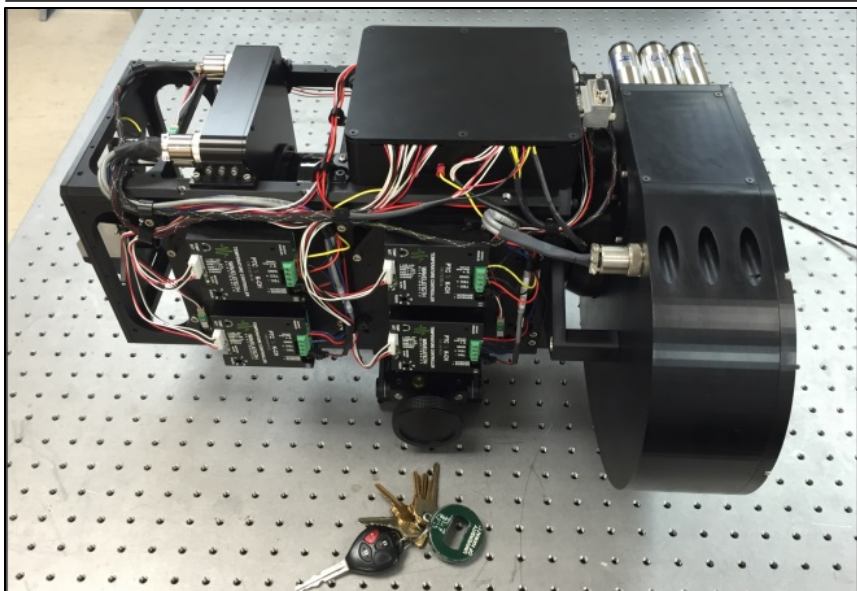
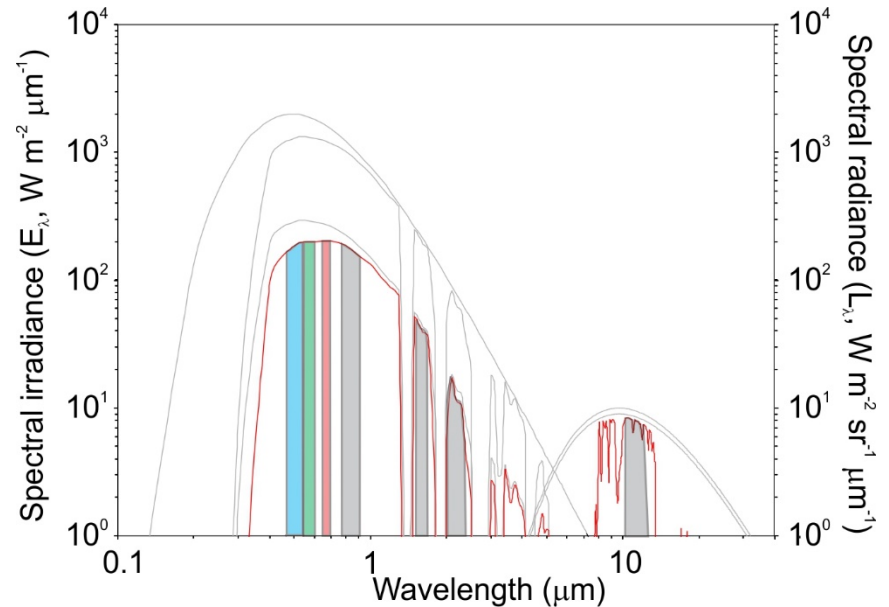
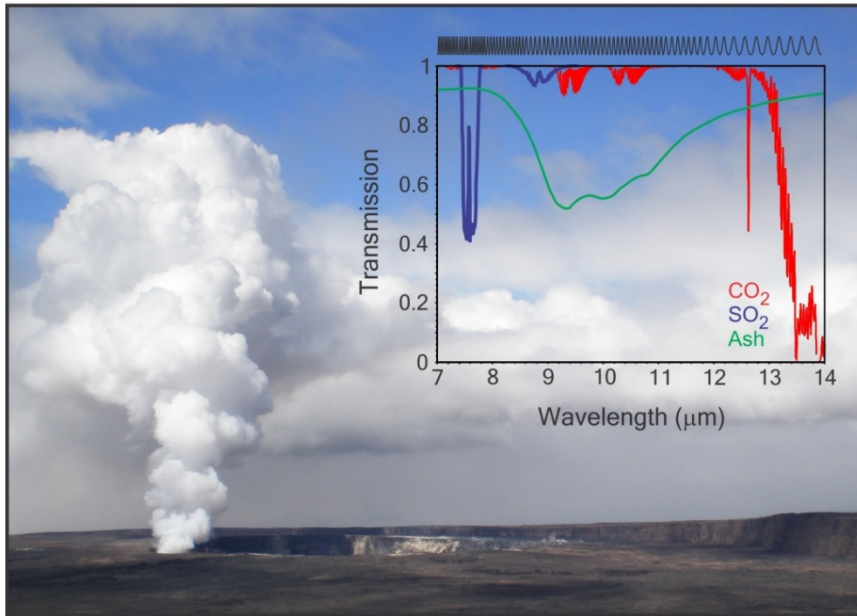
THERMAL INFRA-RED COMPACT IMAGING SPECTROMETER: LABORATORY CHARACTERIZATION AND FIRST FLIGHTS

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Andrea Gabrieli¹, and Casey Honniball¹

1. Hawai'i Institute of Geophysics and Planetology, University of Hawai'i at Mānoa, Honolulu, U.S.A.
2. JAXA, Japan

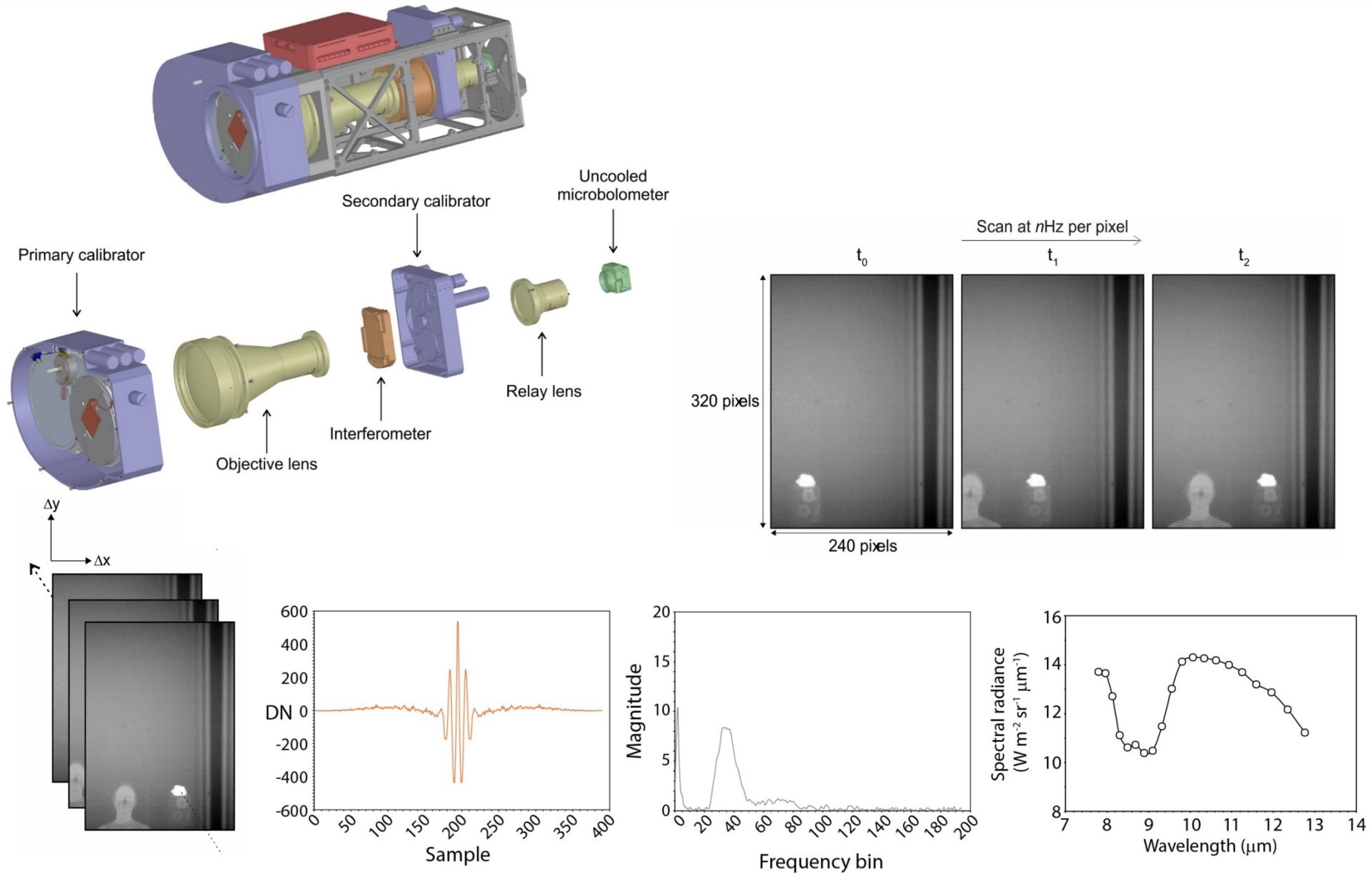


Why is hyperspectral LWIR imaging useful to Earth scientists?

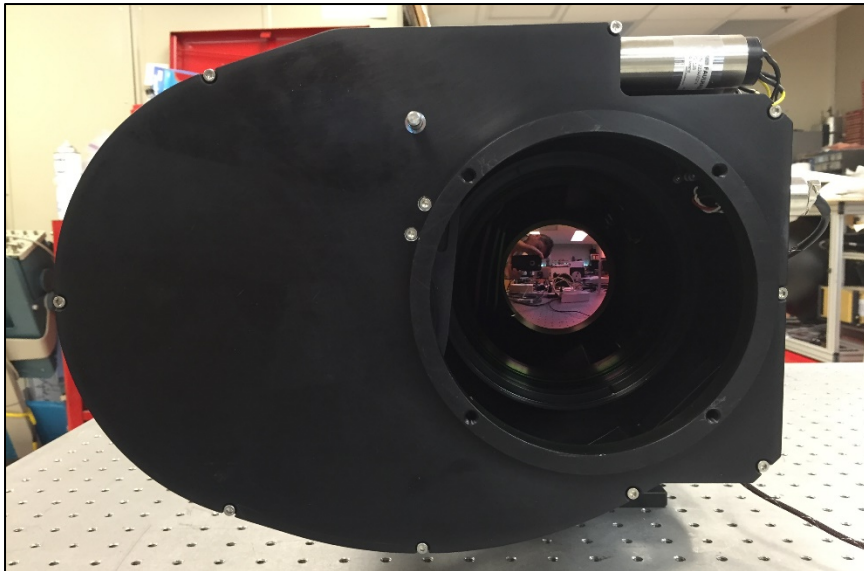
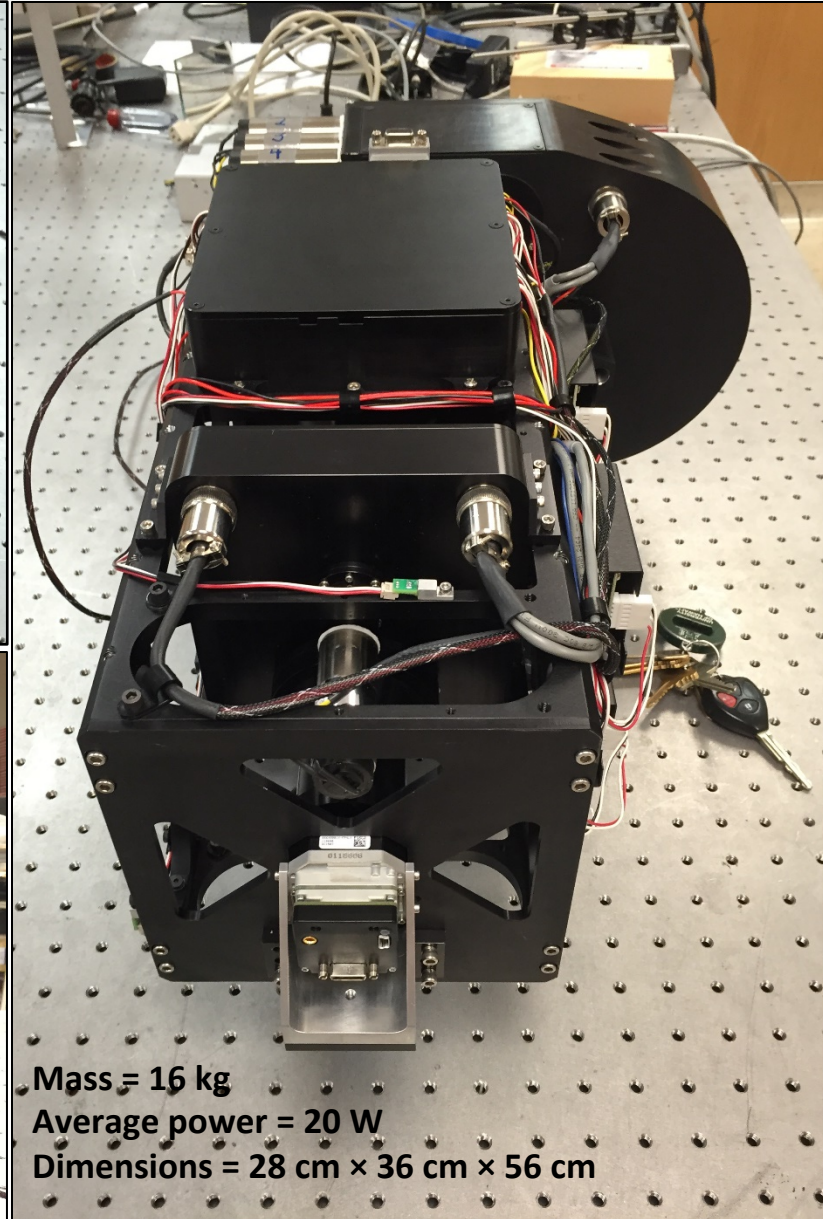
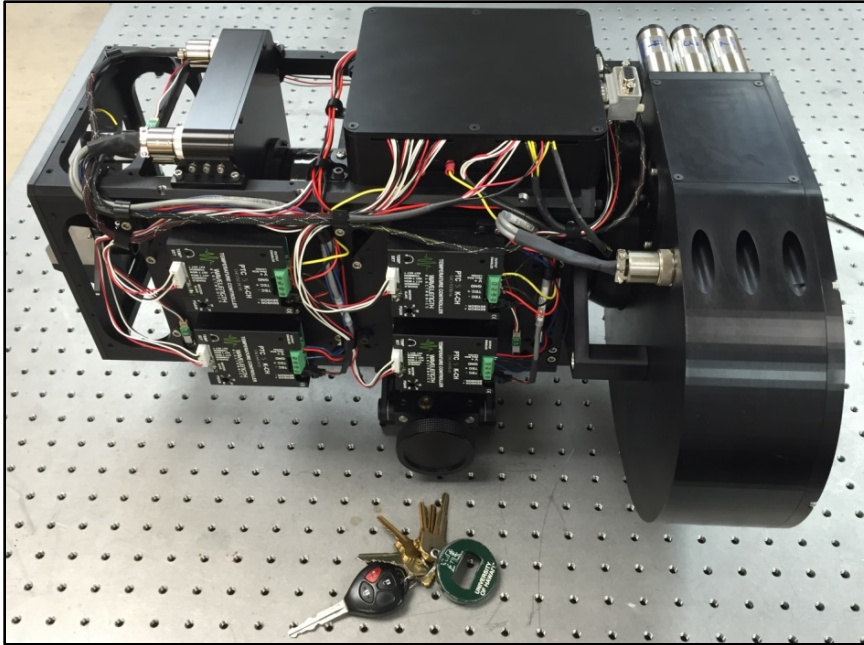


- Many targets of interest to Earth scientists have distinctive spectra at long-wave thermal infrared wavelengths ($\sim 8\text{-}14\text{ mm}$)
- Imaging interferometry has the potential to characterize these targets at high spectral resolution, with high signal-to-noise
- TIRCIS (*Thermal Infra-Red Compact Imaging Spectrometer*) is a compact instrument designed for this purpose, specifically aimed at small- and micro-satellite platforms, funded by NASA's ESTO IIP

TIRCIS - a Spatial Fourier Transform Spectrometer

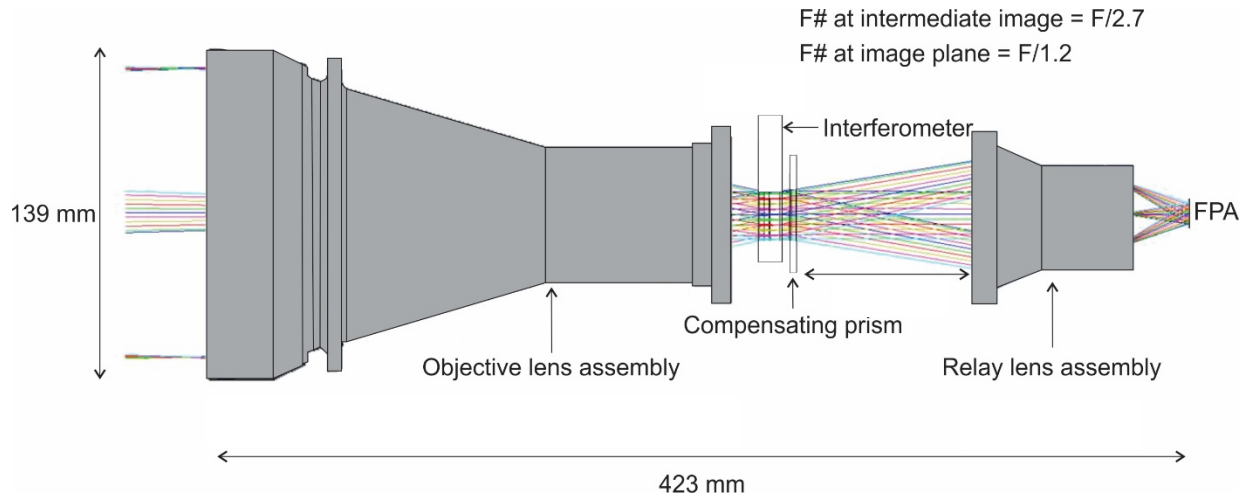


TIRCIS prototype

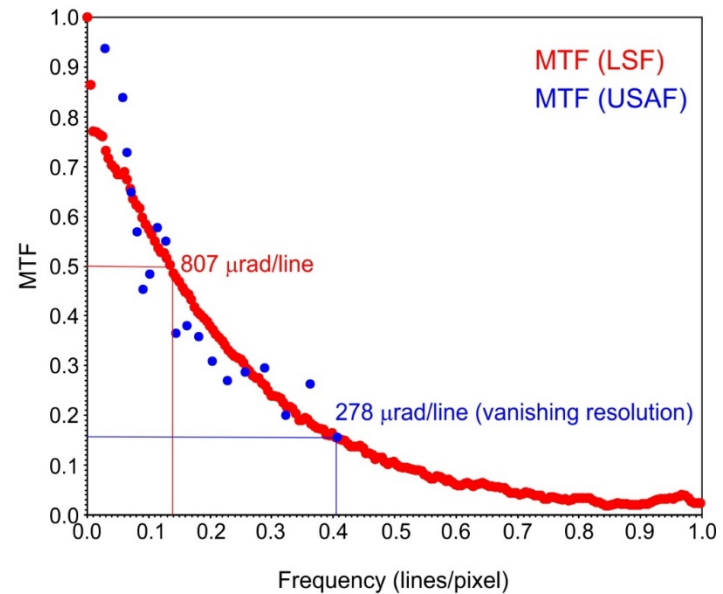
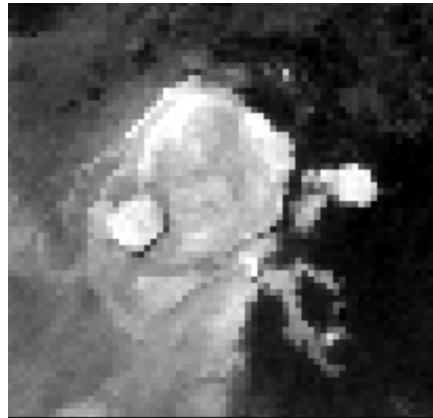


Mass = 16 kg
Average power = 20 W
Dimensions = 28 cm × 36 cm × 56 cm

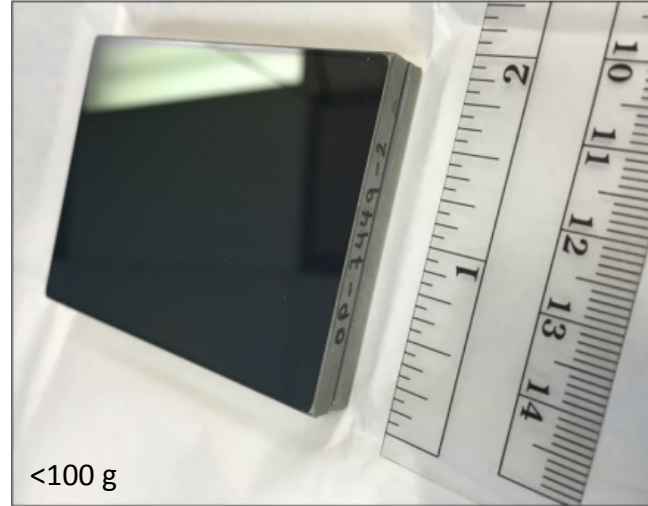
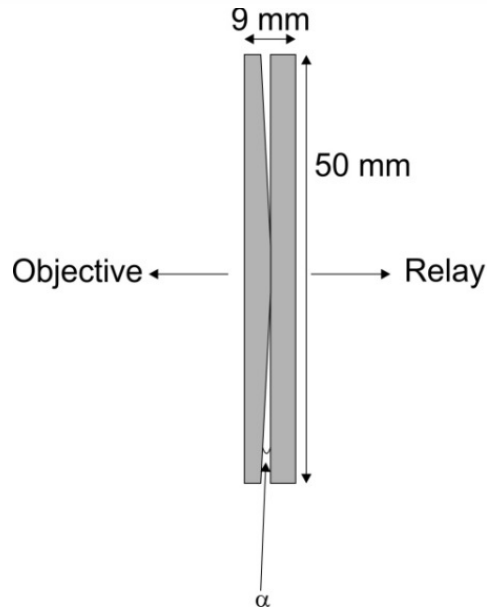
TIRCIS optical design and IFOV



IFOV = 250 mrad; GIFOV = 120 m



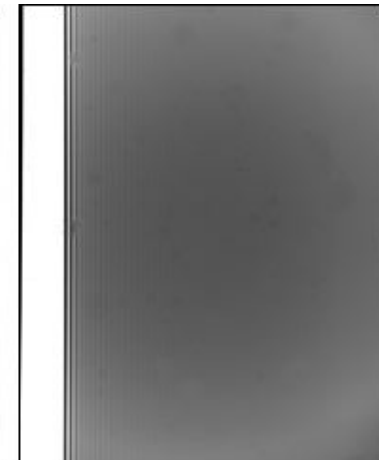
TIRCIS spectral resolution



i. 5 mrad etalon

ii. 15 mrad etalon

iii. 19 mrad etalon



i. ~ 17 bands (8-14 mm) at DI ~ 0.35 mm

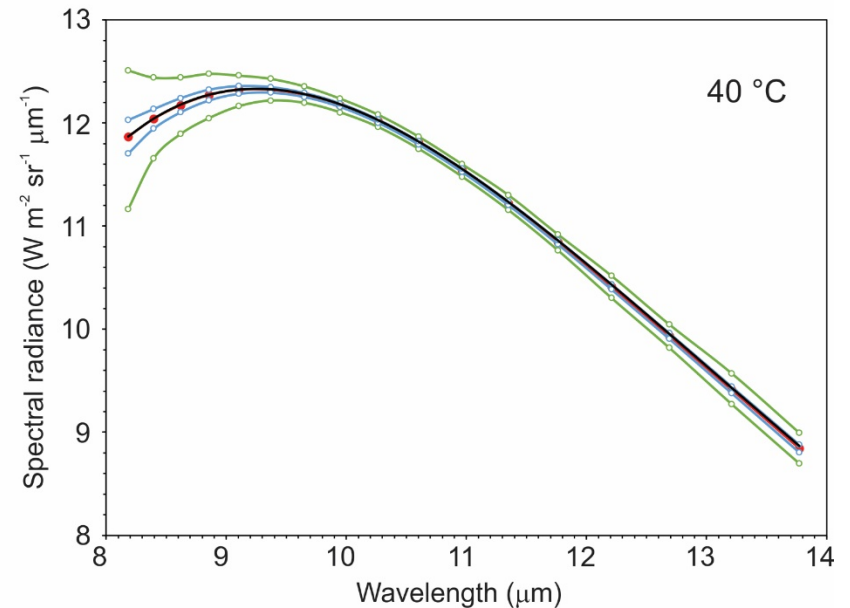
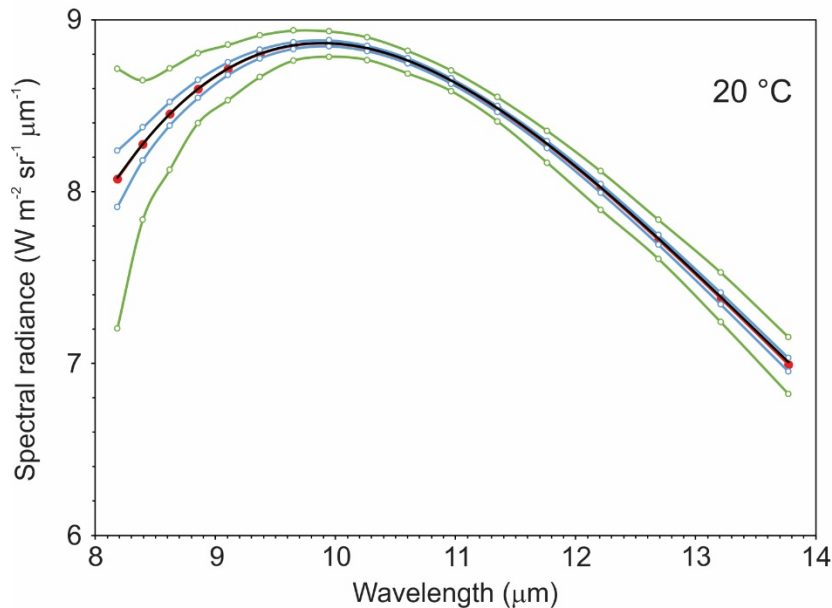
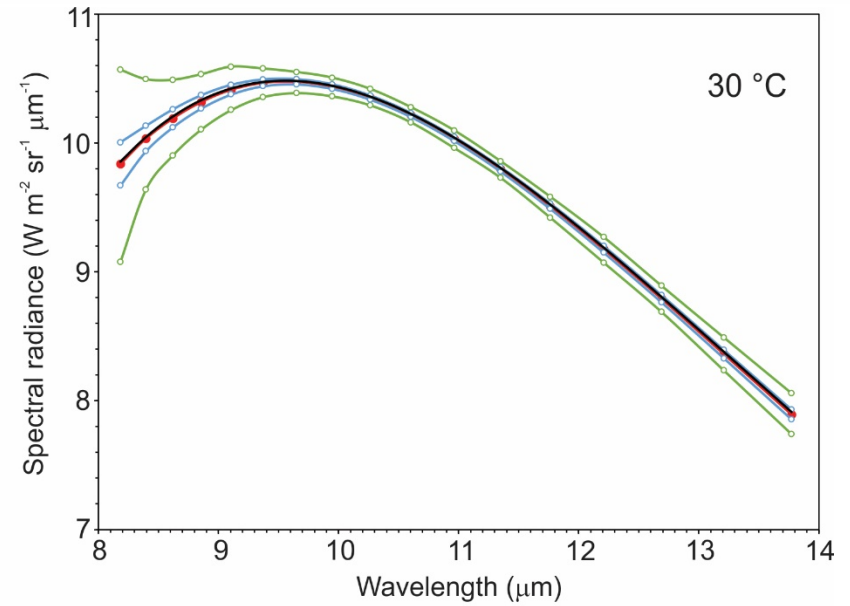
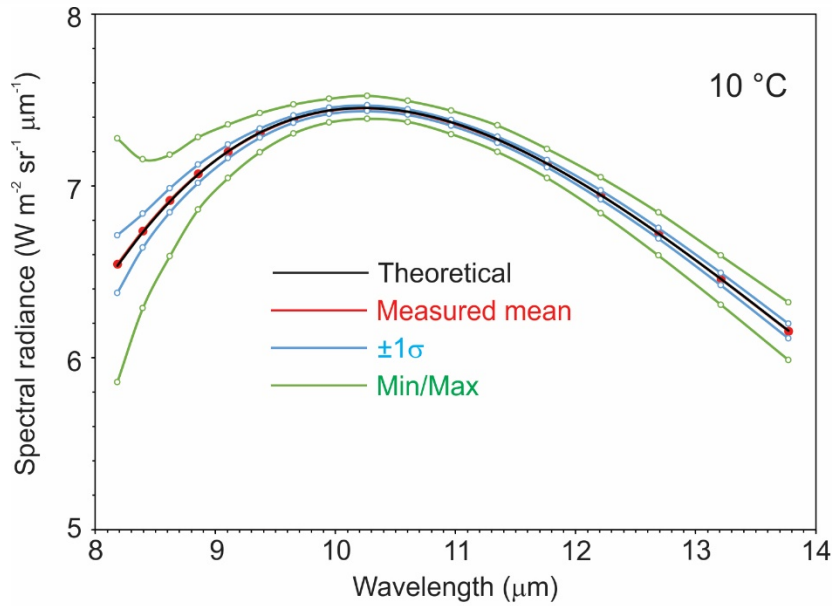
ii. ~ 48 bands (8-14 mm) at DI ~ 0.13 mm

iii. ~ 64 bands (8-14 mm) at DI ~ 0.09 mm



TIRCIS spectro-radiometric performance at 44 cm⁻¹

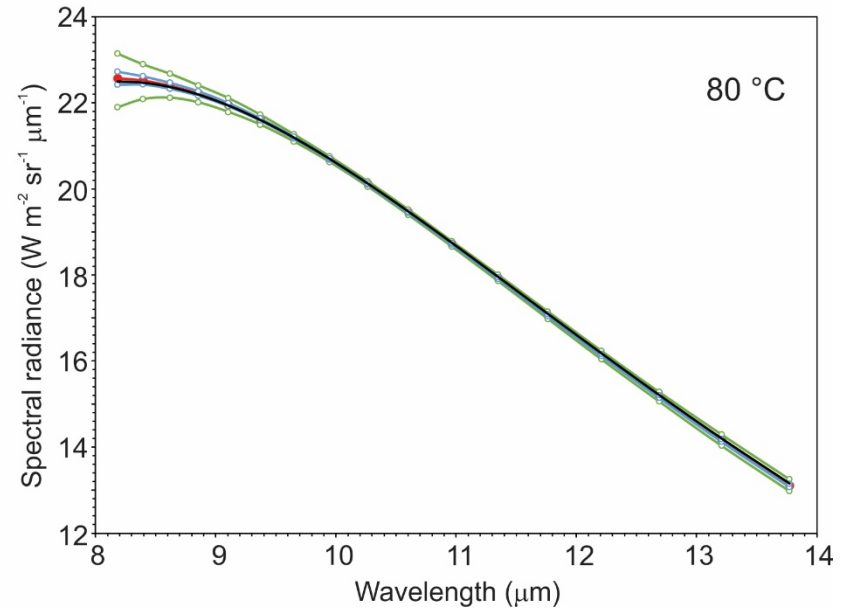
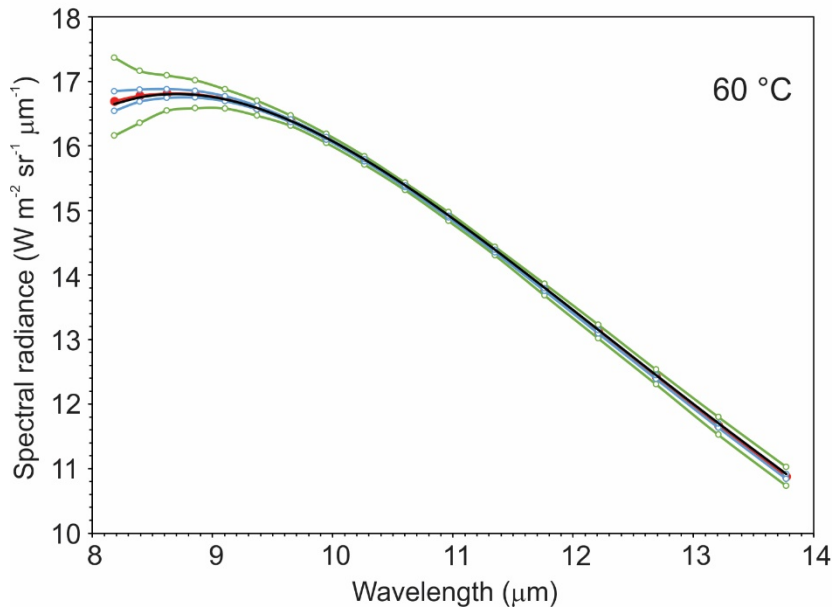
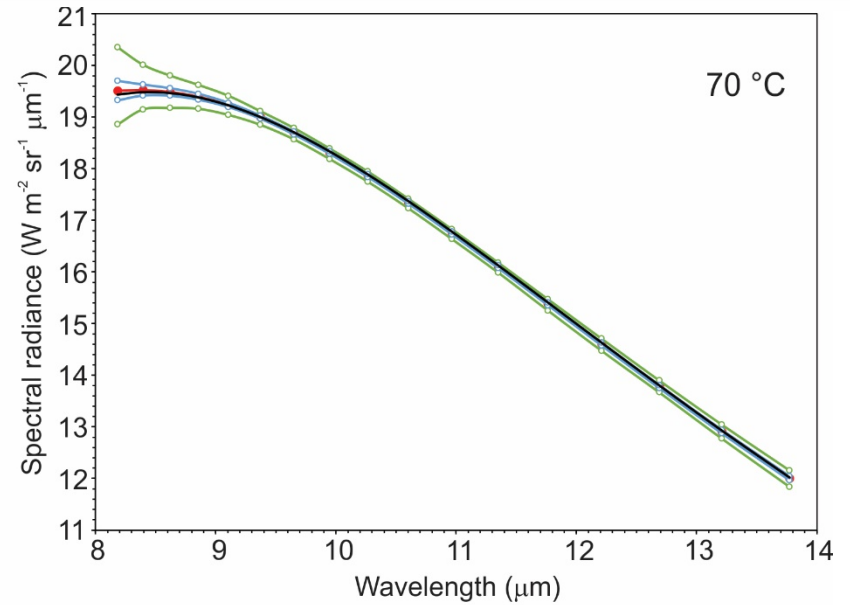
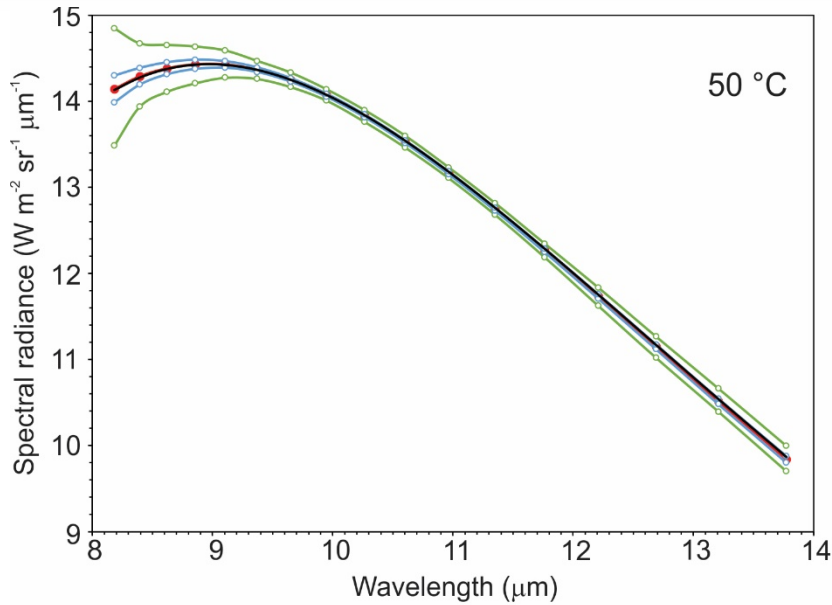
FLIR Photon 320





TIRCIS spectro-radiometric performance at 44 cm⁻¹

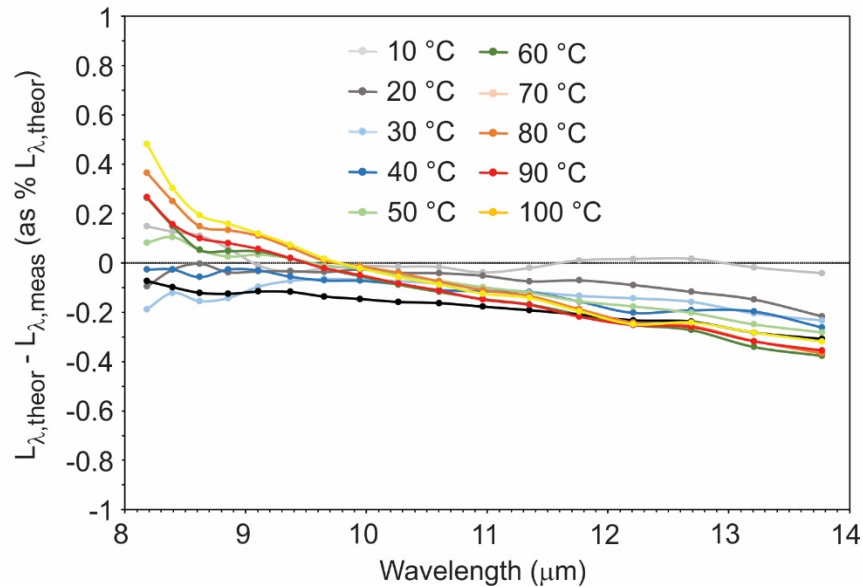
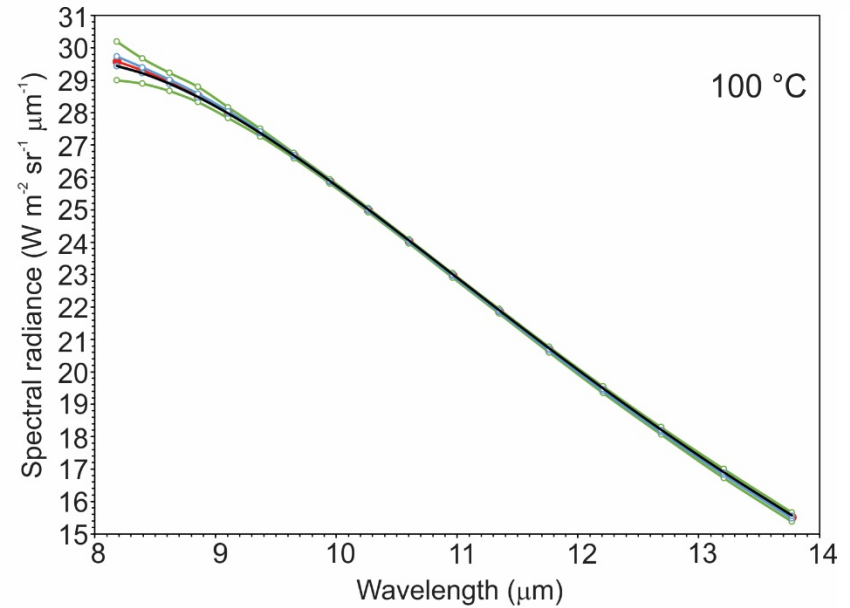
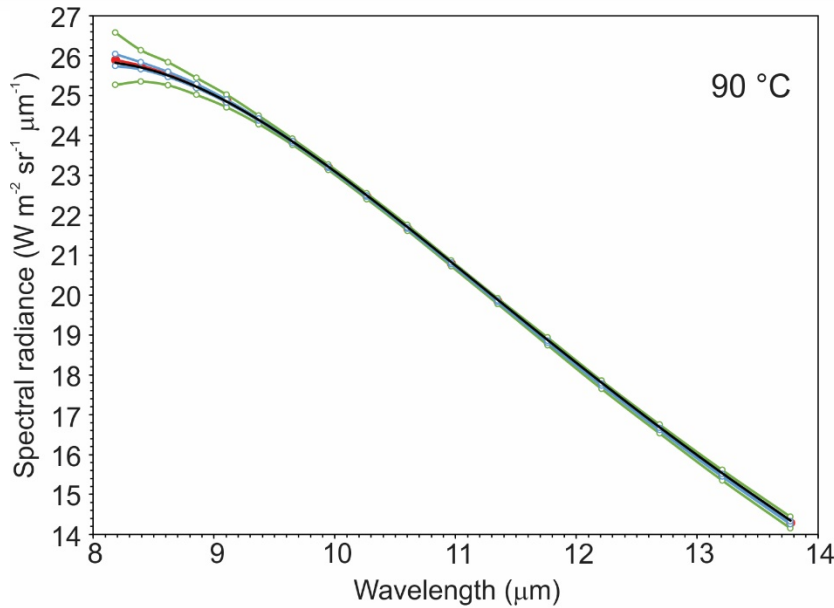
FLIR Photon 320





TIRCIS spectro-radiometric performance at 44 cm⁻¹

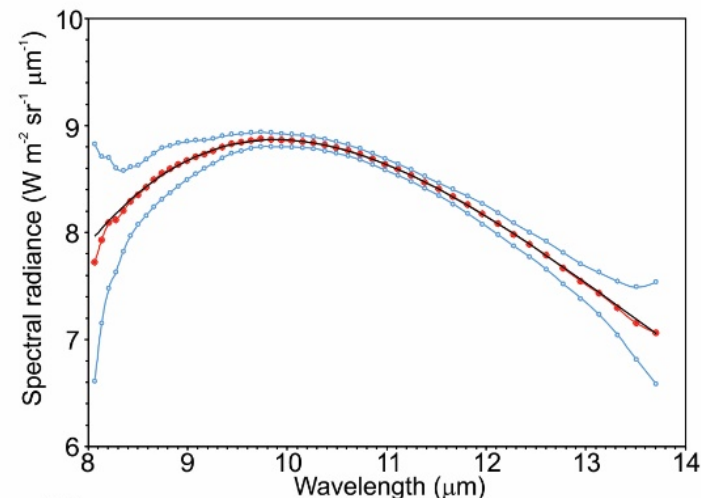
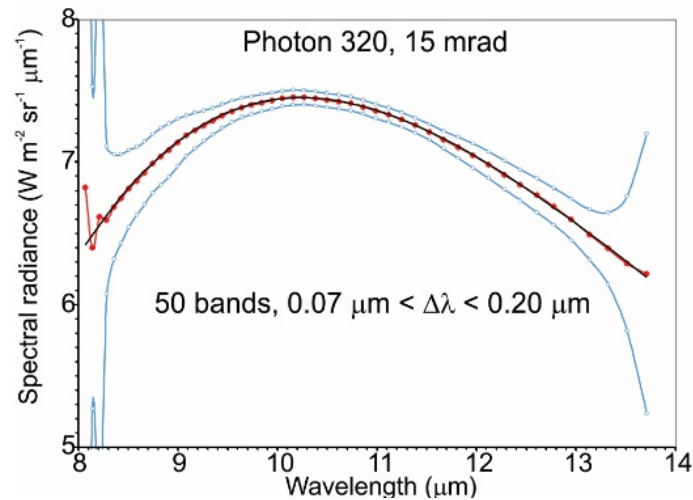
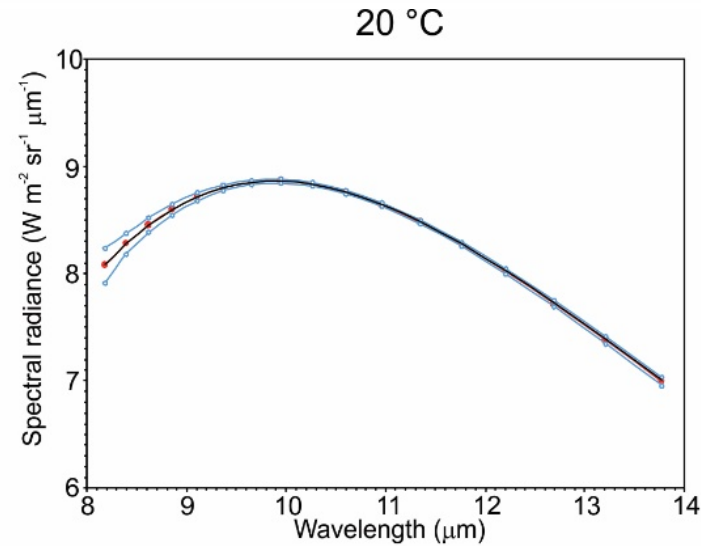
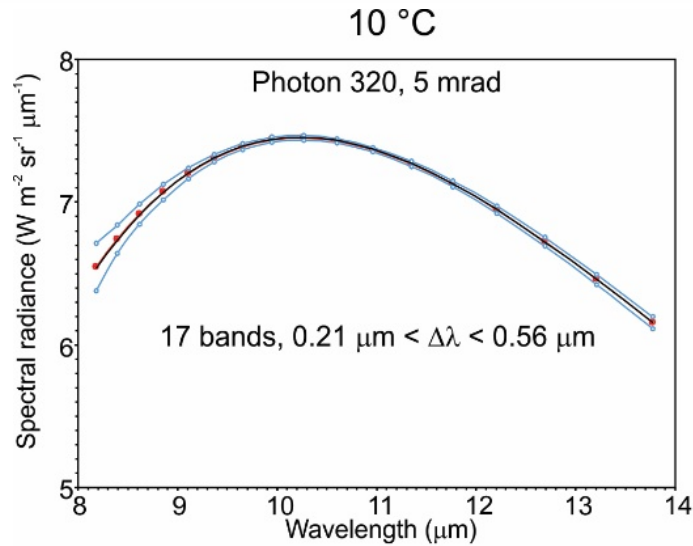
FLIR Photon 320





Comparing the 44 cm^{-1} and 8.7 cm^{-1} interferometers

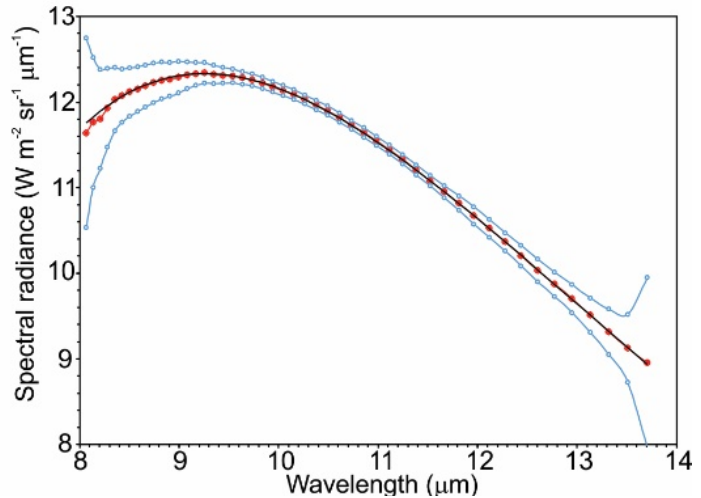
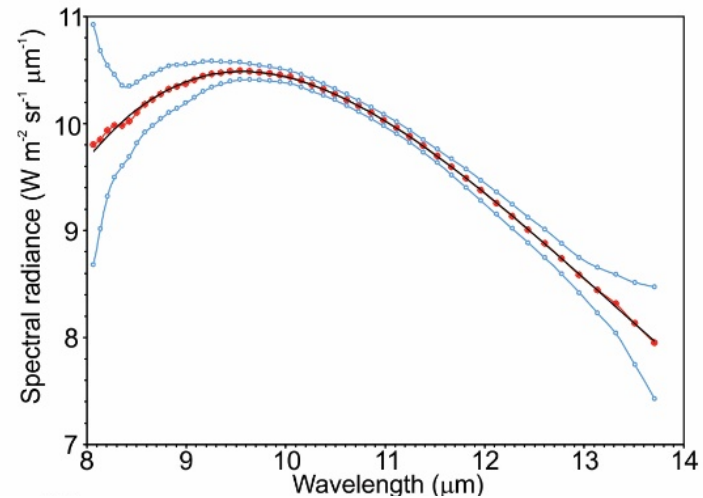
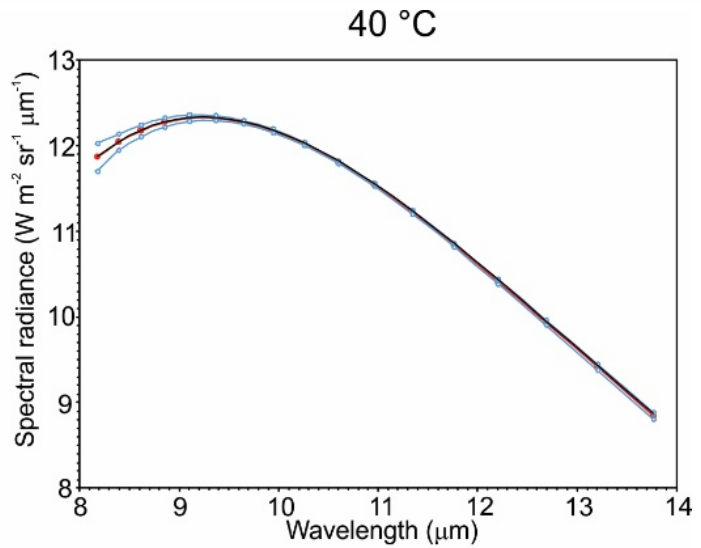
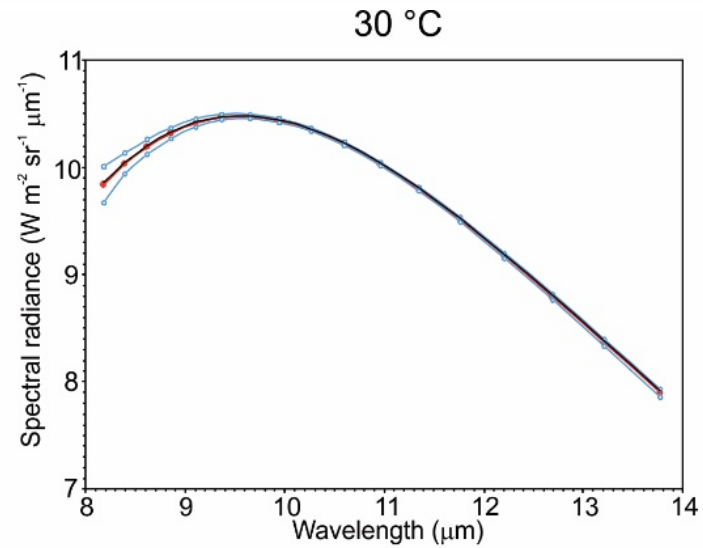
FLIR Photon 320





Comparing the 44 cm^{-1} and 8.7 cm^{-1} interferometers

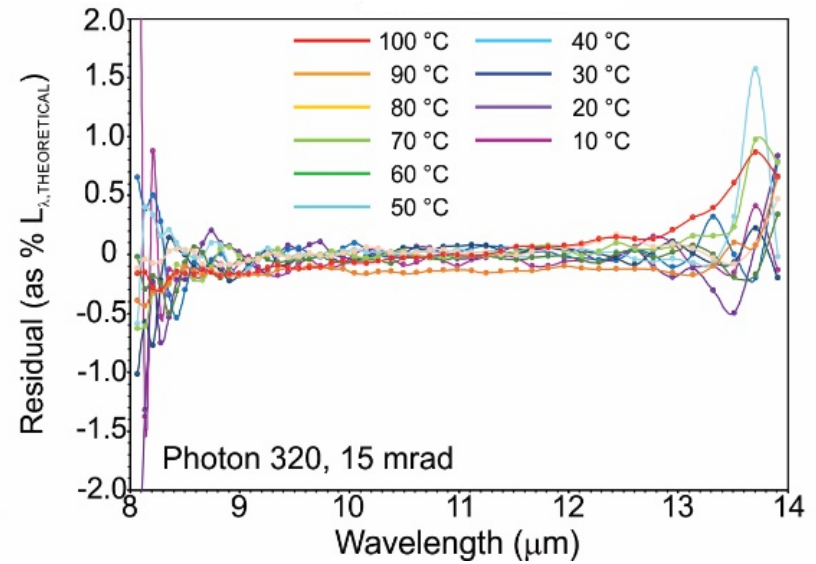
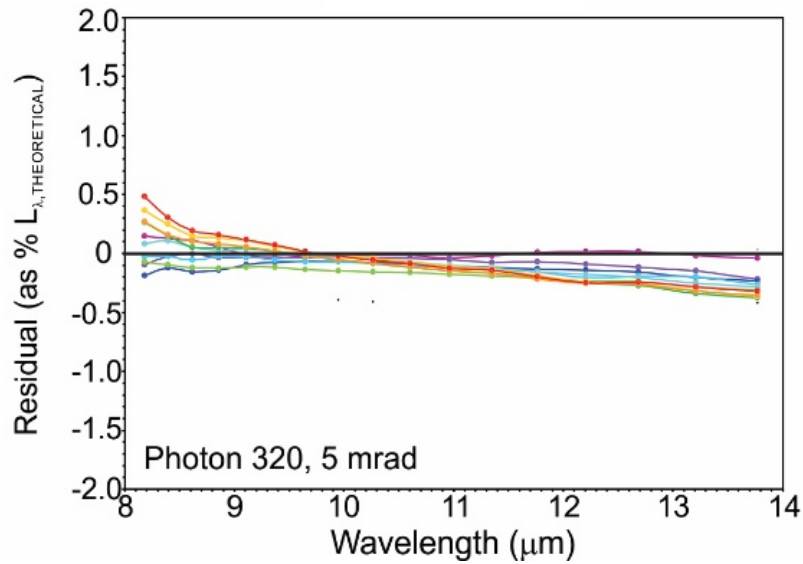
FLIR Photon 320





Noisy, and increasingly noisy, at short and long wavelengths

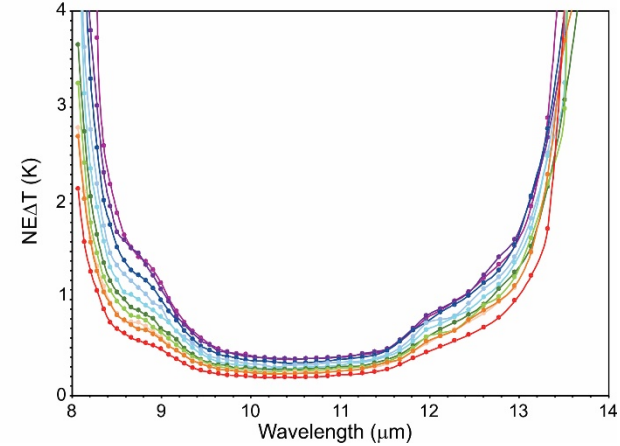
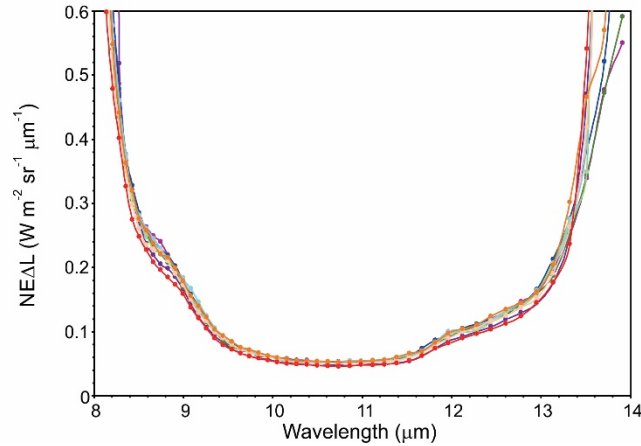
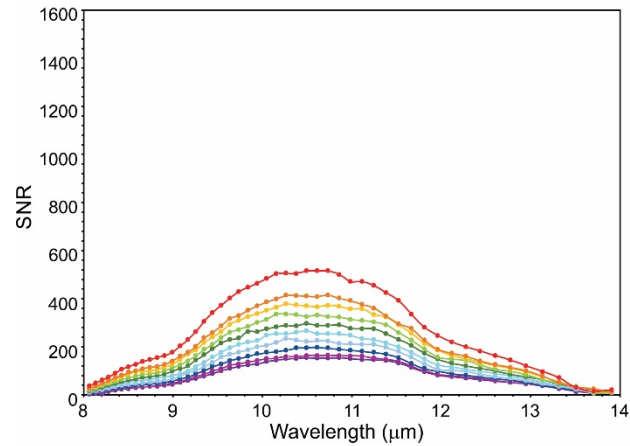
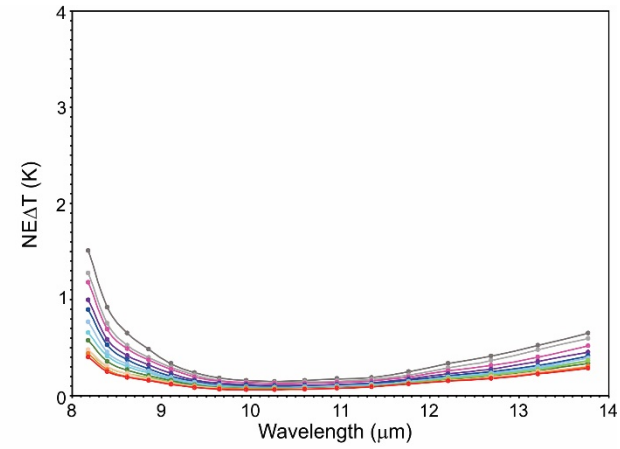
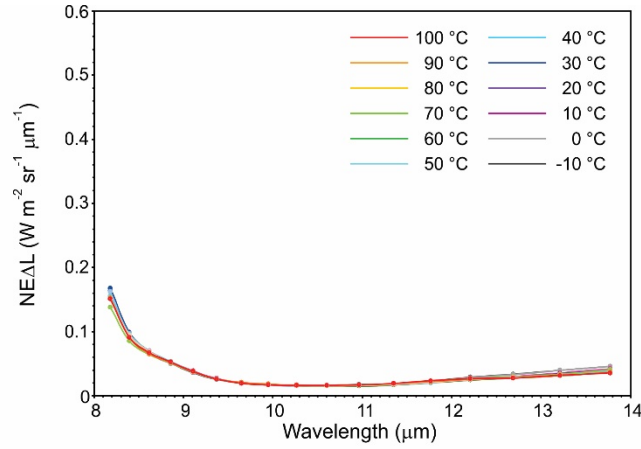
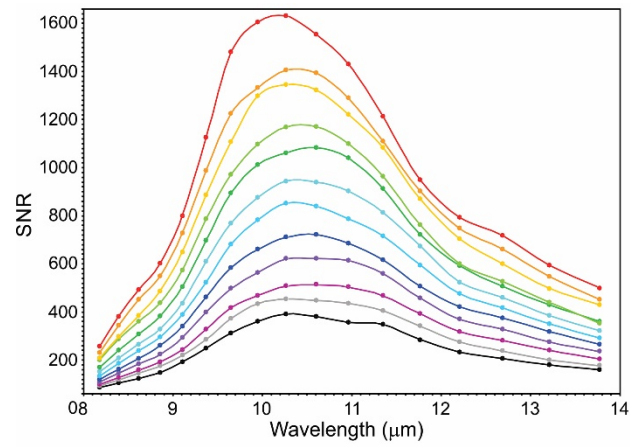
FLIR Photon 320





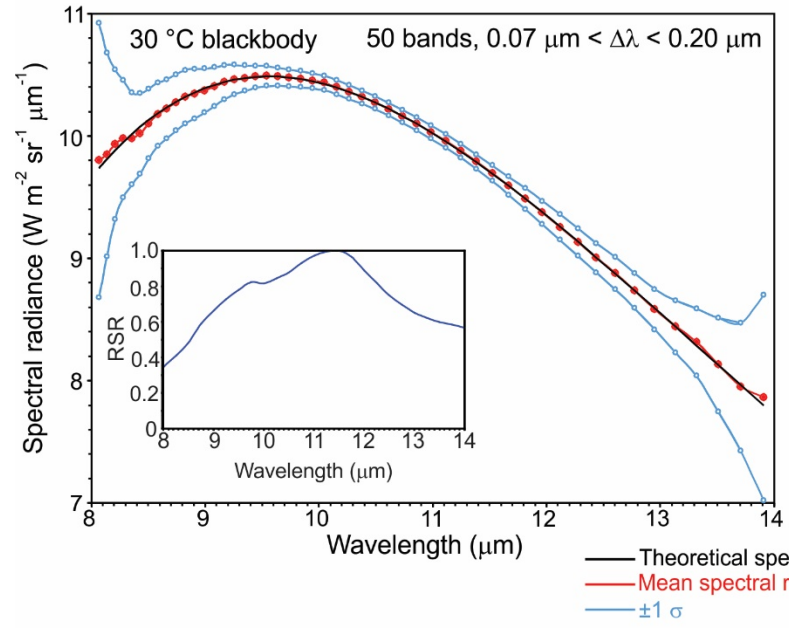
Spectral resolution vs precision at 44 cm^{-1} and 8.7 cm^{-1}

FLIR Photon 320

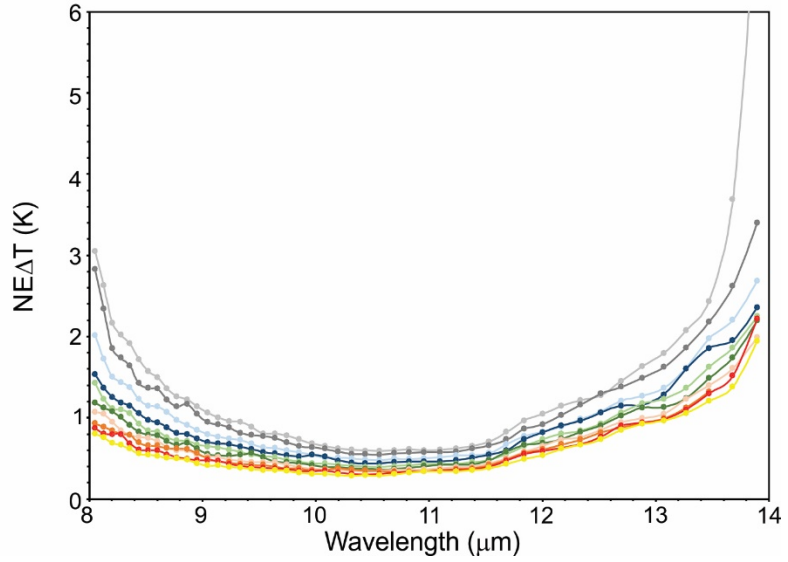
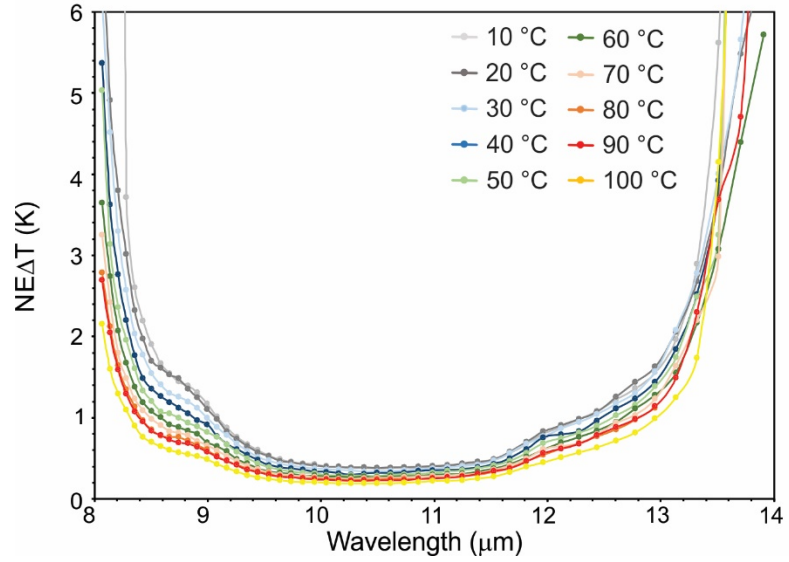
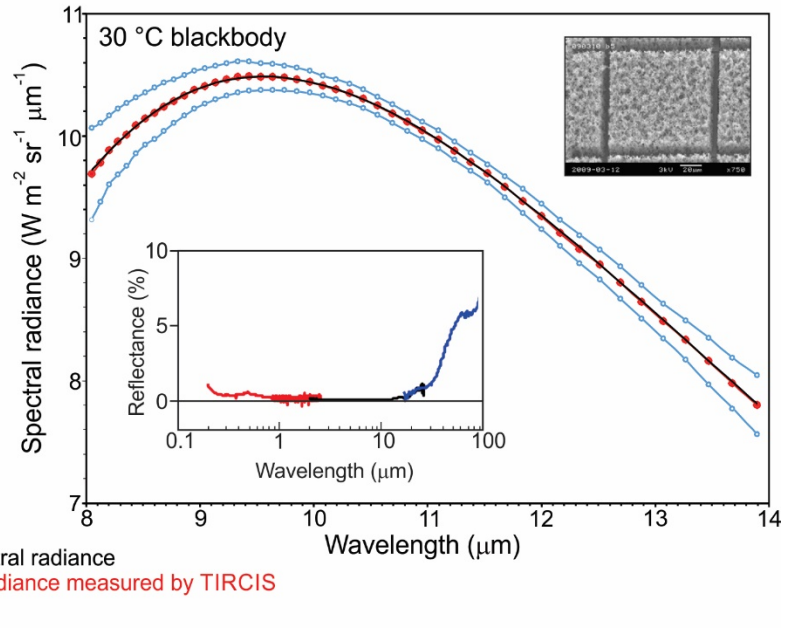


Flatter response can be obtained

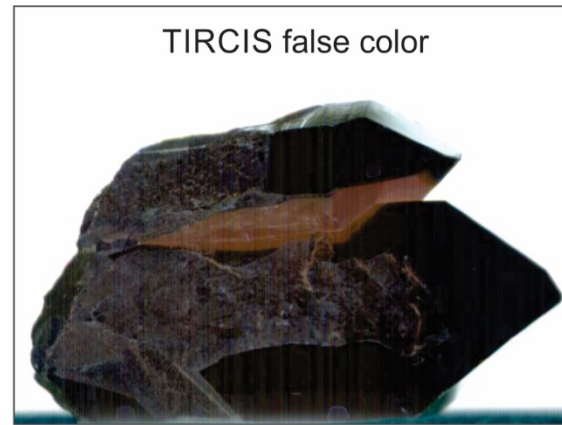
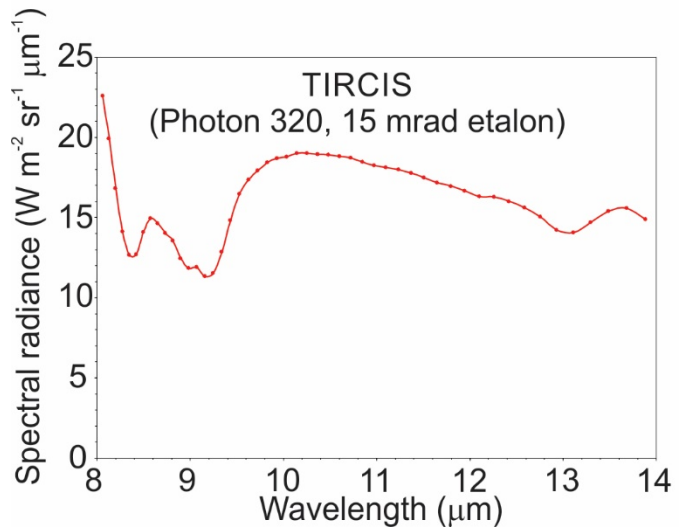
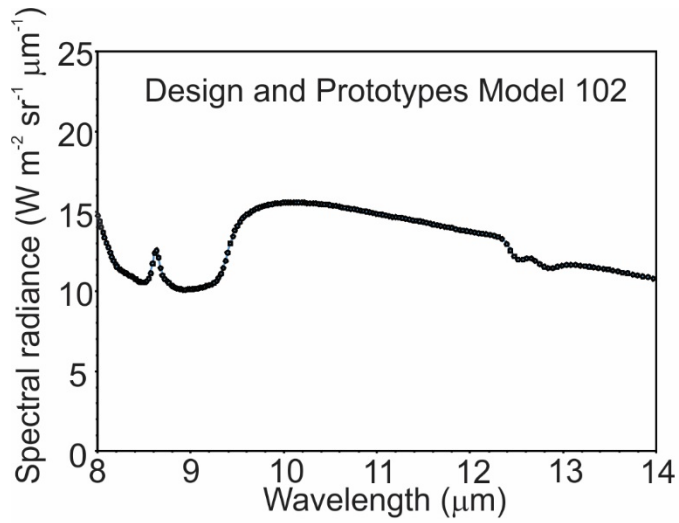
FLIR Photon 320, 15 mrad



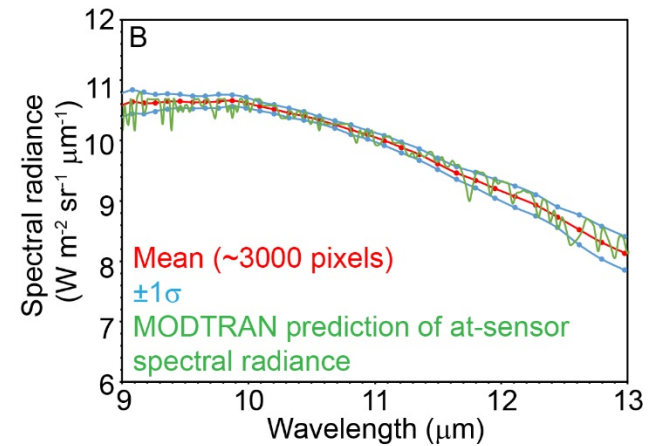
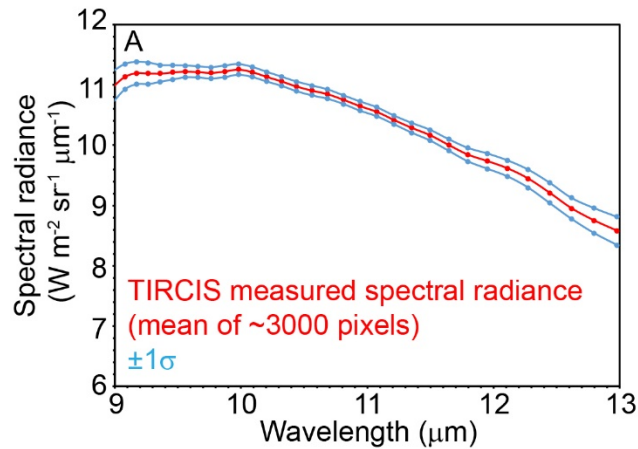
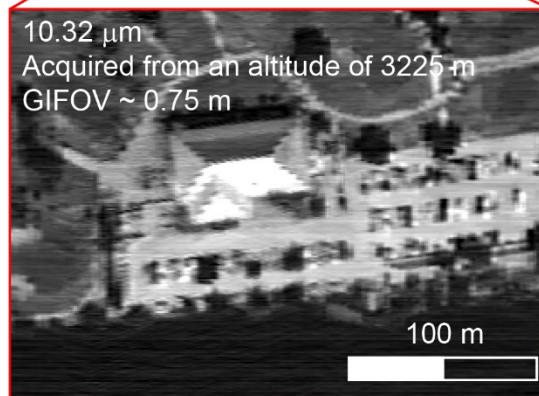
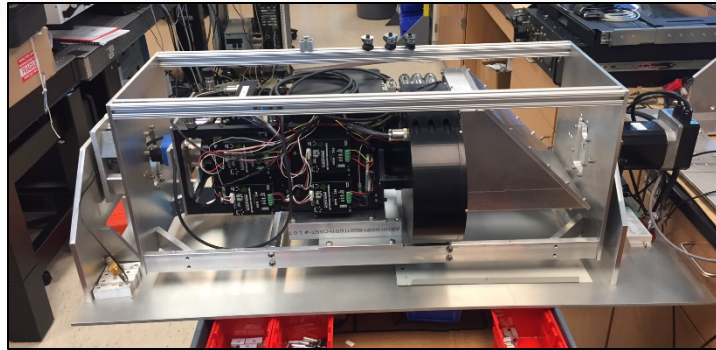
INO μ X-Cam 15 mrad



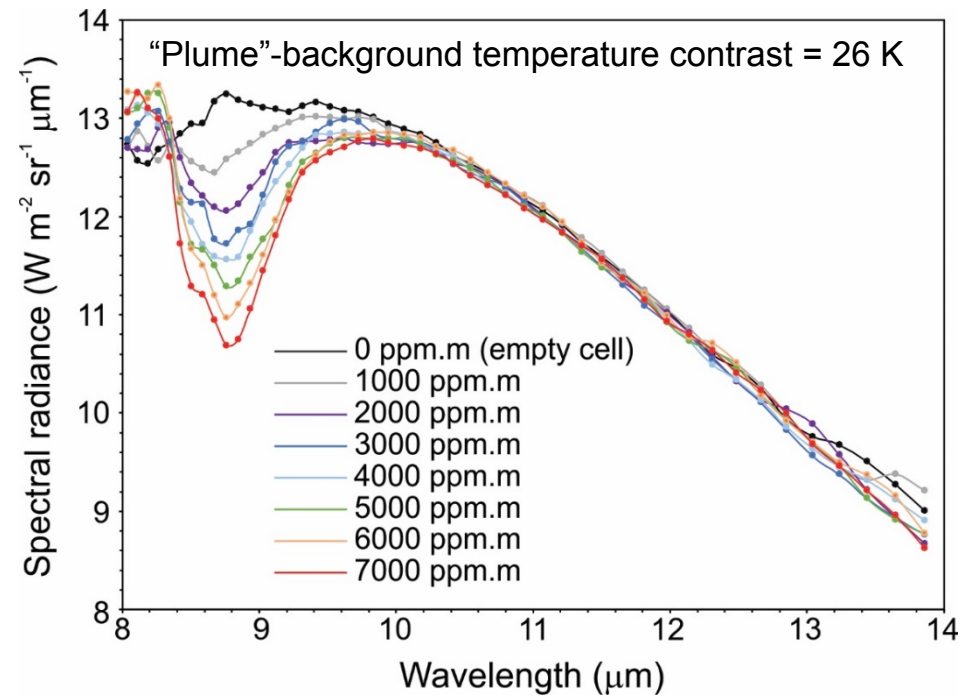
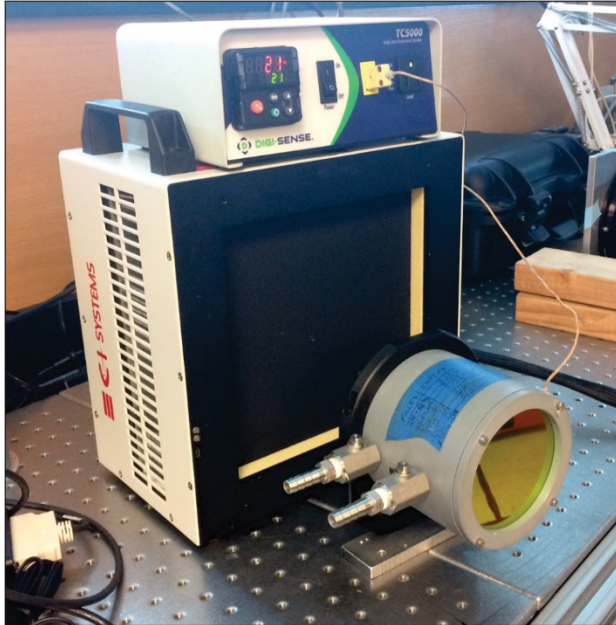
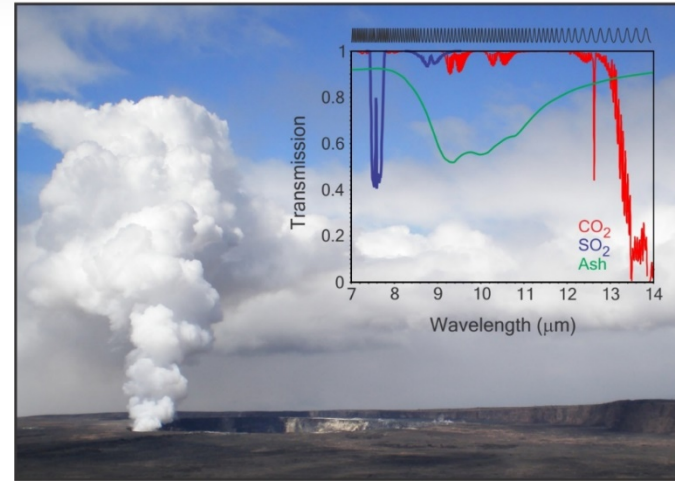
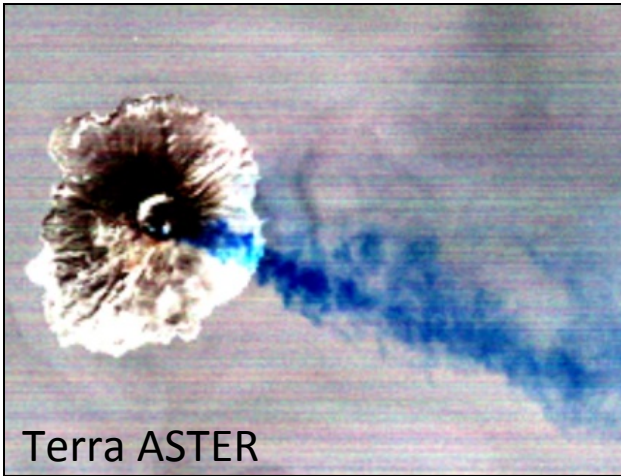
Some initial results



Some results from first test flight



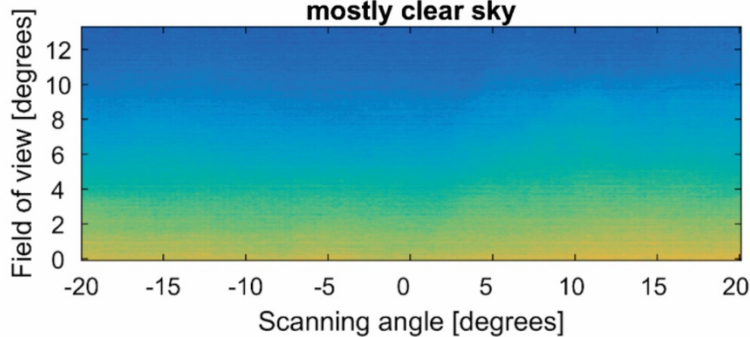
Simulating volcanic SO₂ gas measurements



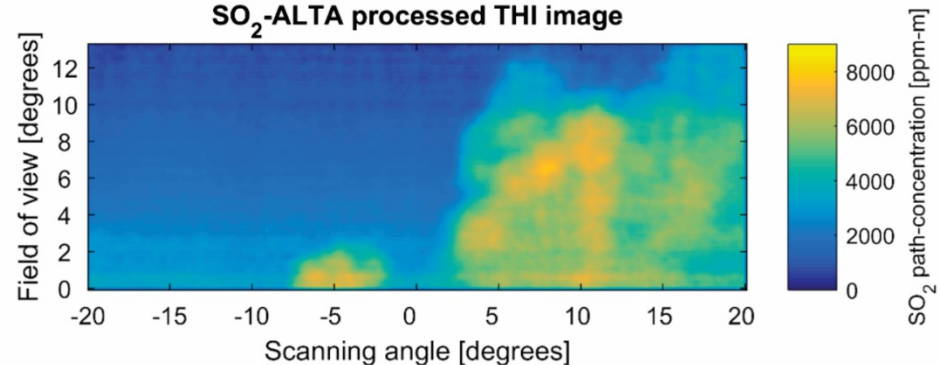
Work in progress is to invert $W/m^2/sr/mm$ to ppm.m



Raw THi image. Kilauea volcano, 7/26/2016, 3:30 PM
mostly clear sky



SO₂-ALTA processed THi image



Gabrieli, A., Wright, R., Lucey, P.G., Garbeil, H., Pilger, E., Porter, J.N., & Wood, M. (2016). Characterization and initial field test of 8-14 μm thermal infrared hyperspectral imager for measuring SO₂ in volcanic plumes. *Bulletin of Volcanology*, 78, DOI: 10.1007/s00445-016-1068-6

Gabrieli, A., Porter, J., Wright, R., Lucey, P.G., (2017). Validation of the accuracy and precision of volcanic SO₂ path concentration retrievals from long-wave infrared hyperspectral images. In review.

Summary

- Imaging interferometry can provide high spatial, high spectral, and high temporal resolution (if incorporated into a constellation of microsattellites) image data for quantifying the chemical composition of Earth's surface and atmosphere
- TIRCIS can provide 17-50 spectral measurements in the 8-14 μm window, with peak SNR of 200-800:1 (at 30 °C)
- TIRCIS has a mass of 16 kg (not lightweighted), a volume of 28 cm \times 36 cm \times 56 cm, and a steady state power consumption of \sim 20 W, making it eminently consistent with integration into a microsattellite (e.g. 50 kg), or constellation of microsattellites.
- Thanks to ESTO IIP for support

