

# SLI-T Integrated Photonic Spectrometer

*THE VALUE OF PERFORMANCE.*

***NORTHROP GRUMMAN***

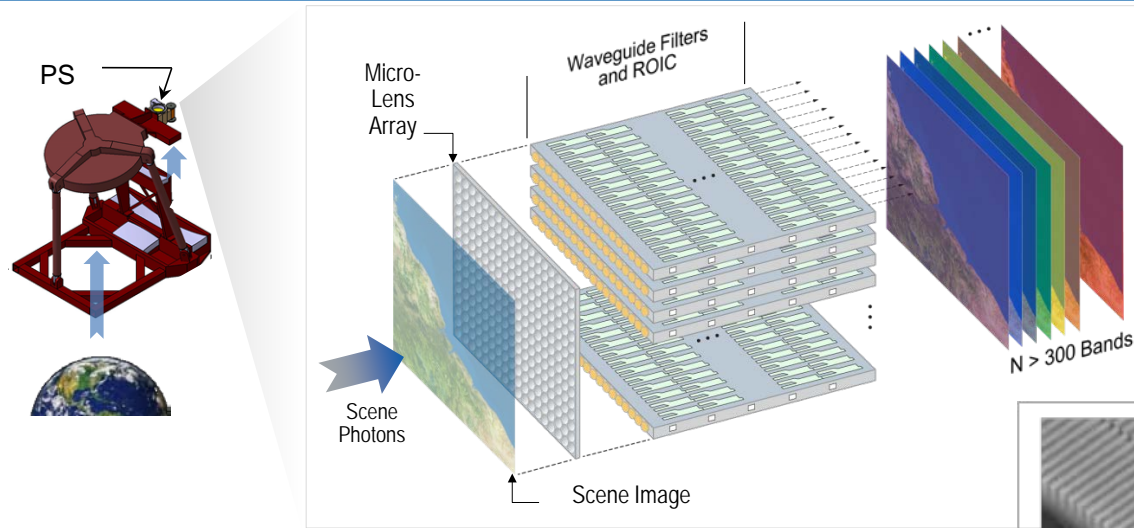
June 14, 2017

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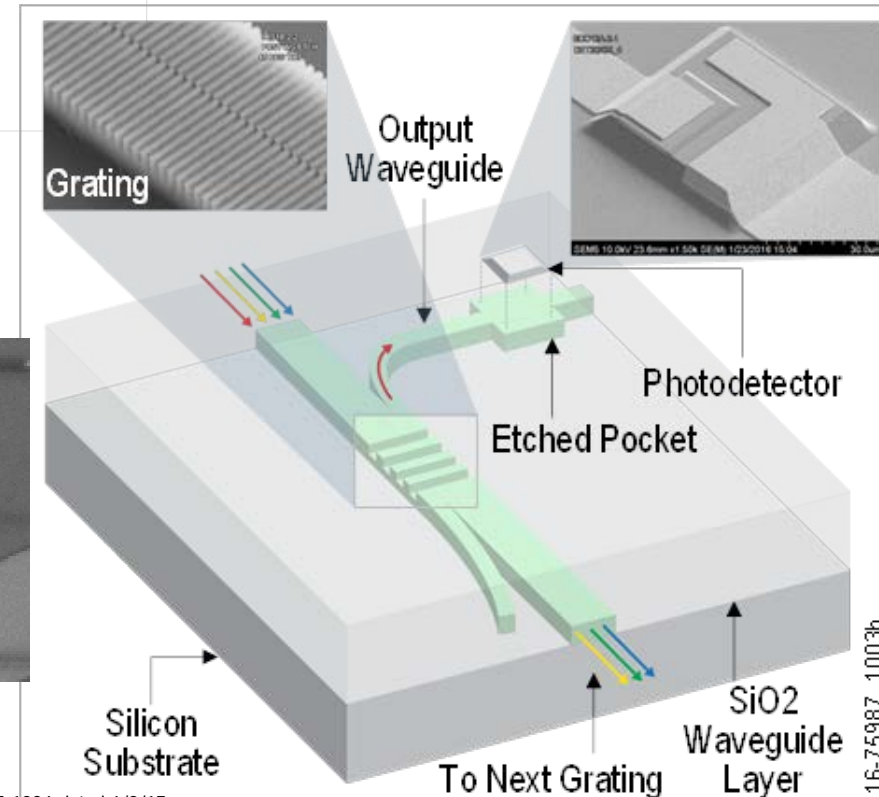
# Photonic Spectrometer Overview

- Approach and component demonstrations
- ESTO study – EVEREST simulations
- ESTO SLI-T status

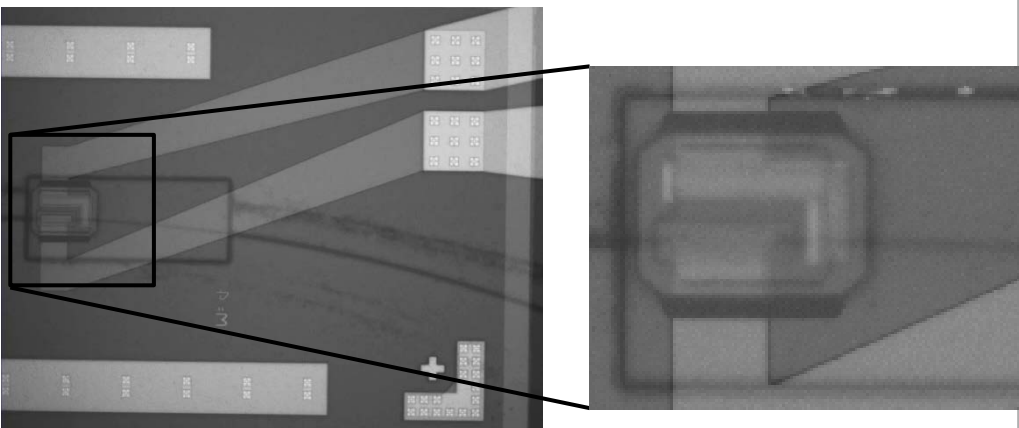
# Photonic Spectrometer Approach



- Micro-fabricated photonic filters and DAHI integrated photodetectors replace free-space optics
- Enables new image acquisition modes (staring, TDI)

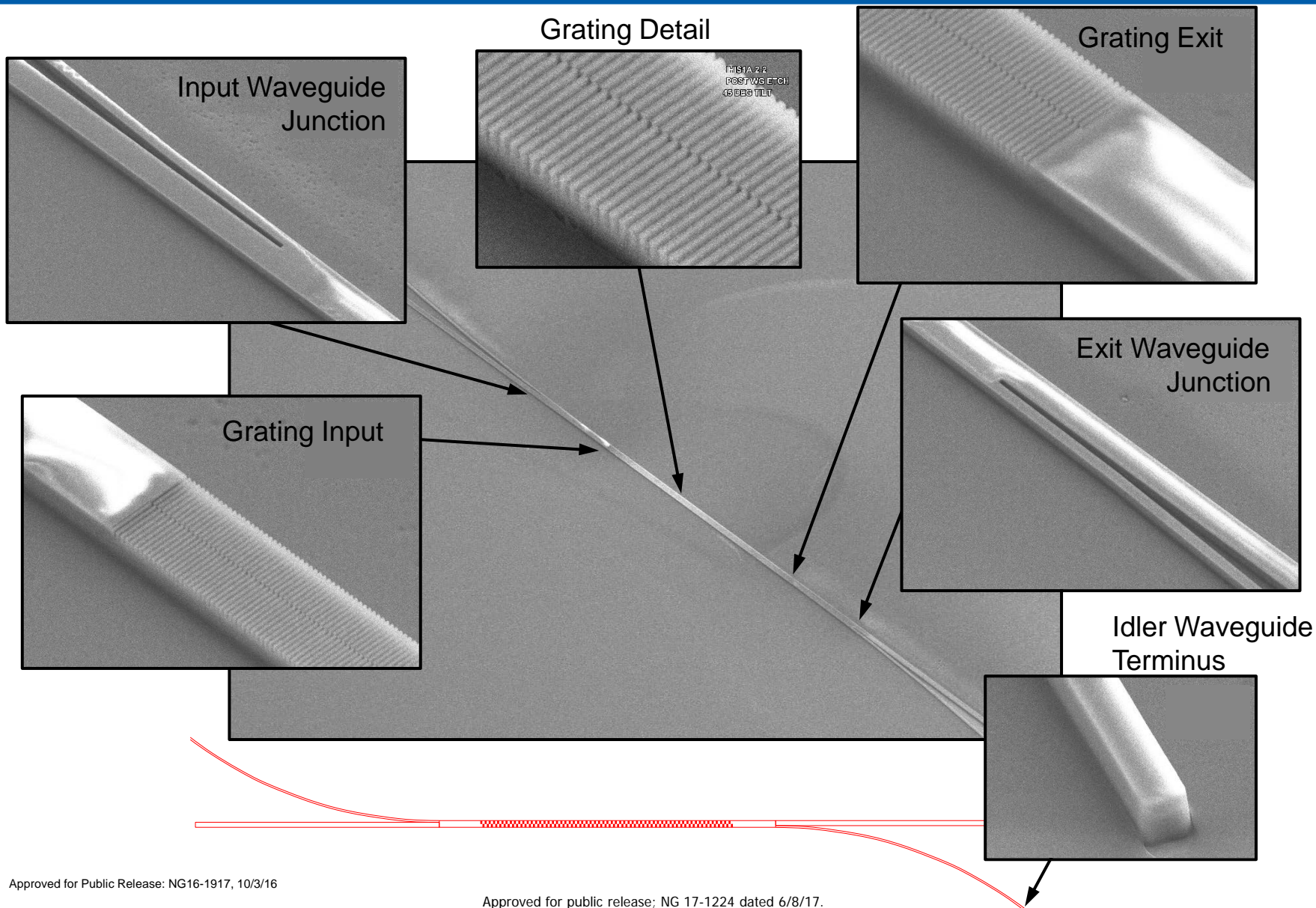


16-75987\_1003b



**DAHI Integration**

# NGAS Program for Waveguide Filter Development

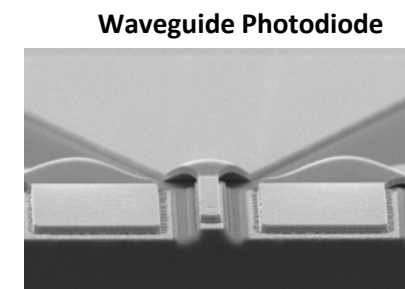
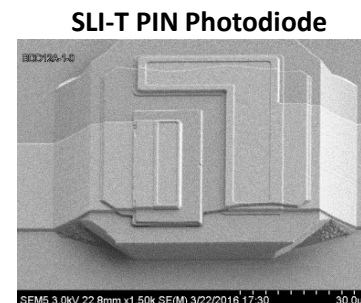
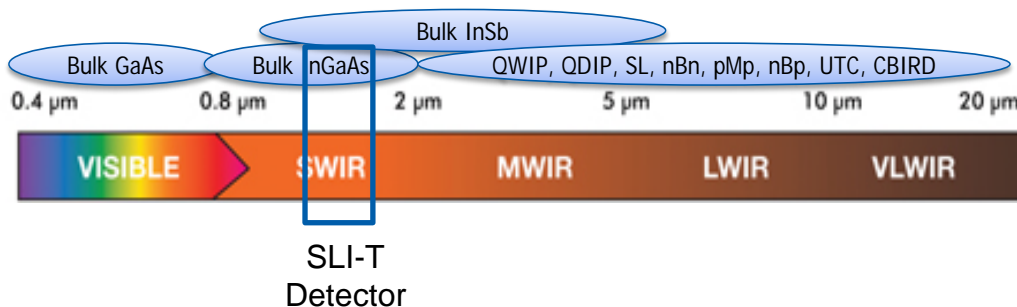


# NG Detector and Integration Capabilities

	NG Demonstrated
Device type / configuration	<ul style="list-style-type: none"> <li>• <b>PIN photodiodes</b></li> <li>• Schottky diodes</li> <li>• Avalanche photodiodes</li> <li>• Photo-transistors</li> <li>• Quantum wells and graded superlattices</li> <li>• Type-II tunnel diodes</li> <li>• Waveguide filters and gratings</li> </ul>
Material systems	Lattice-matched, strained, and metamorphic materials: <ul style="list-style-type: none"> <li>• <b>InP/InGaAs</b></li> <li>• GaAs/AlGaAs</li> <li>• InP/InAlAs/InGaAs/InAlGaAs/ InGaAsP/InGaAsSb</li> <li>• InAs/GaSb/AlSb/AlGaSb/InAlAs</li> </ul>

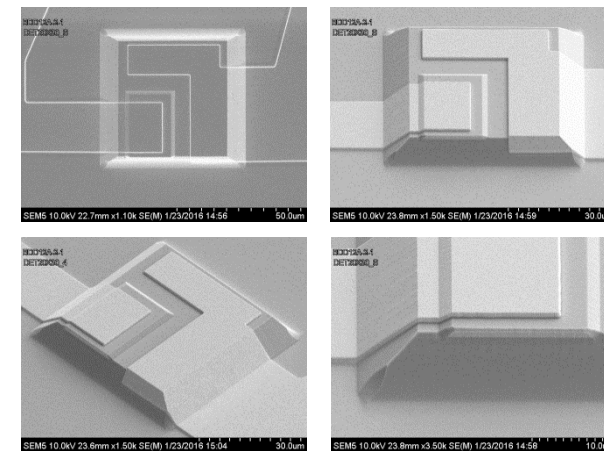
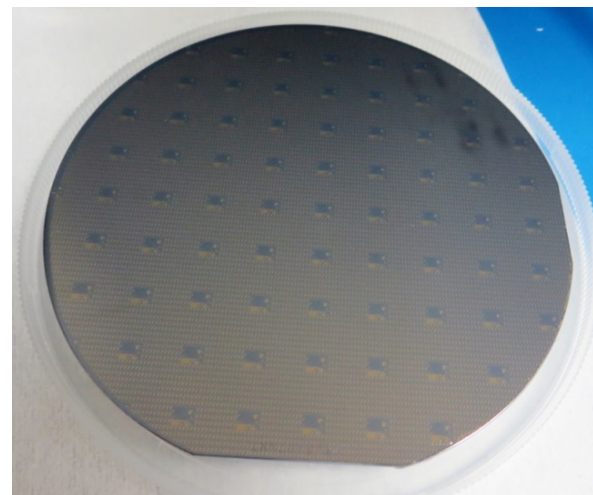
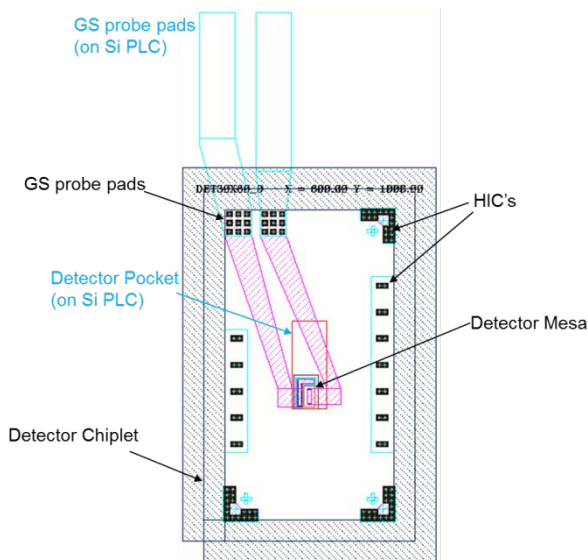
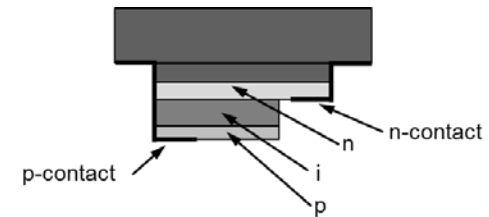
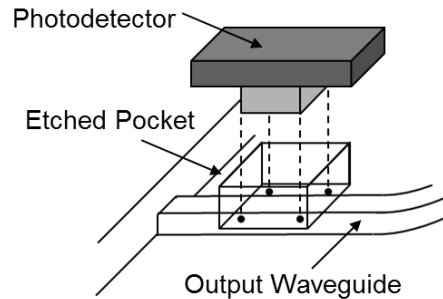
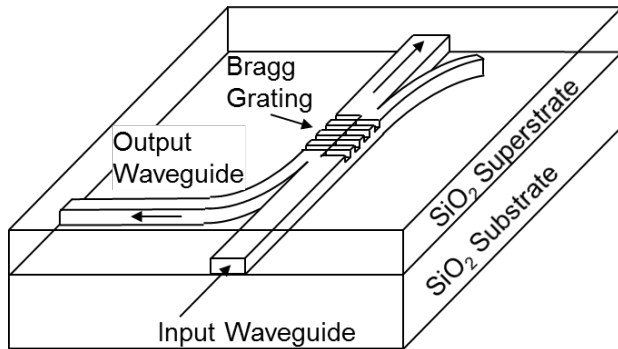
	NG Demonstrated
Device layout	<ul style="list-style-type: none"> <li>• Mesa diode</li> <li>• Surface normal coupled with front- or back-illumination</li> <li>• <b>Edge coupling to diode or waveguide</b></li> <li>• Discrete and array of diodes</li> </ul>
Integration capabilities	<ul style="list-style-type: none"> <li>• <b>DAHI integration</b></li> <li>• Monolithic integration through epitaxial design</li> <li>• Regrowth of multiple epitaxial device structures on same substrate</li> </ul>

## Spectrum Coverage



# Black Diamond: Waveguide-Detector Integration

- Black Diamond program demonstrated waveguide/detector integration using NG heterogeneous integration processes

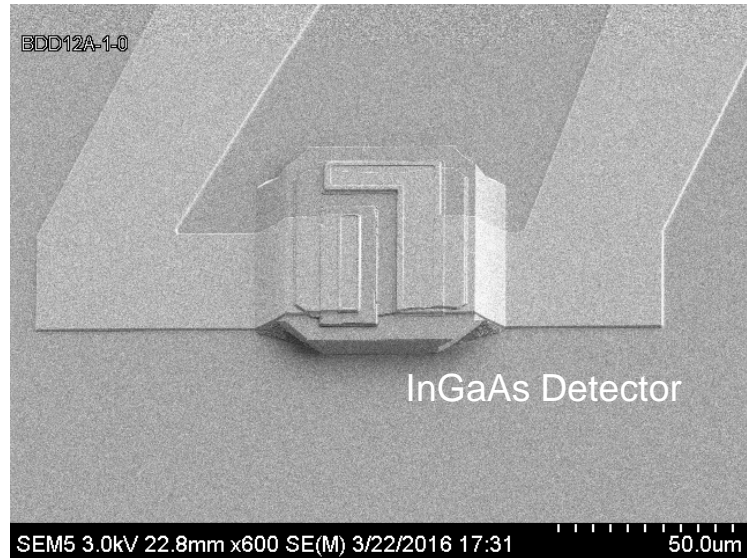
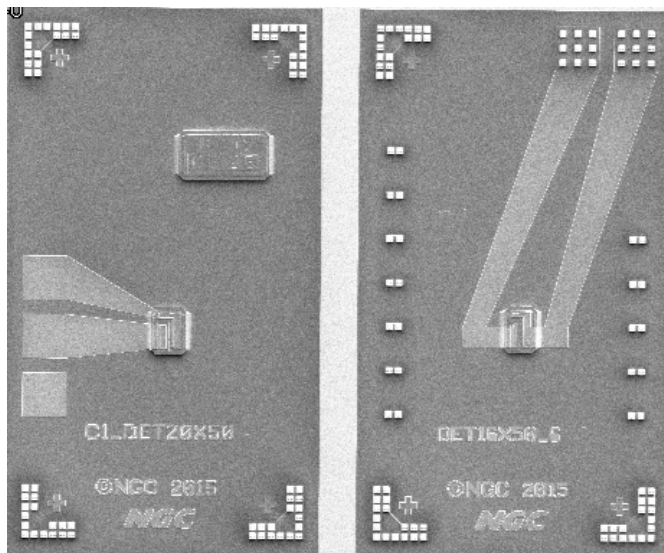


SEMs – detector topology

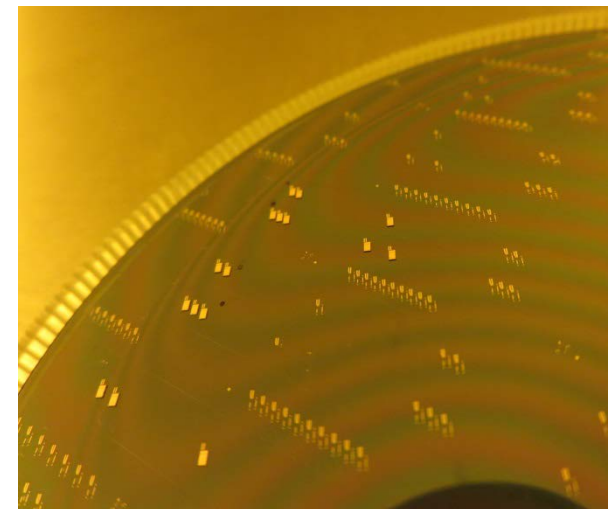
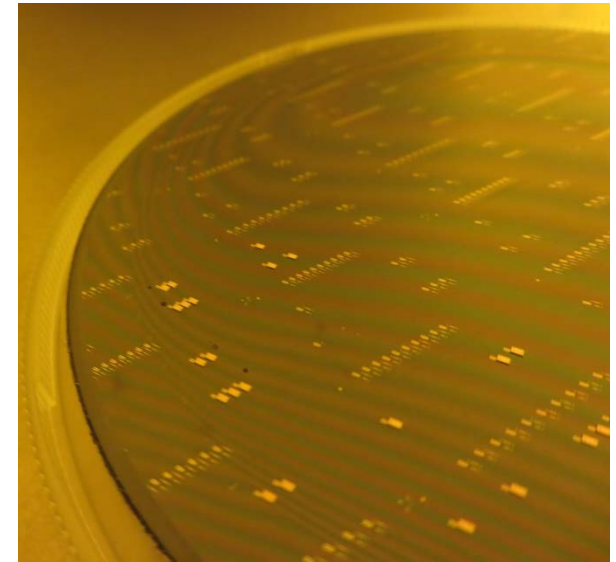
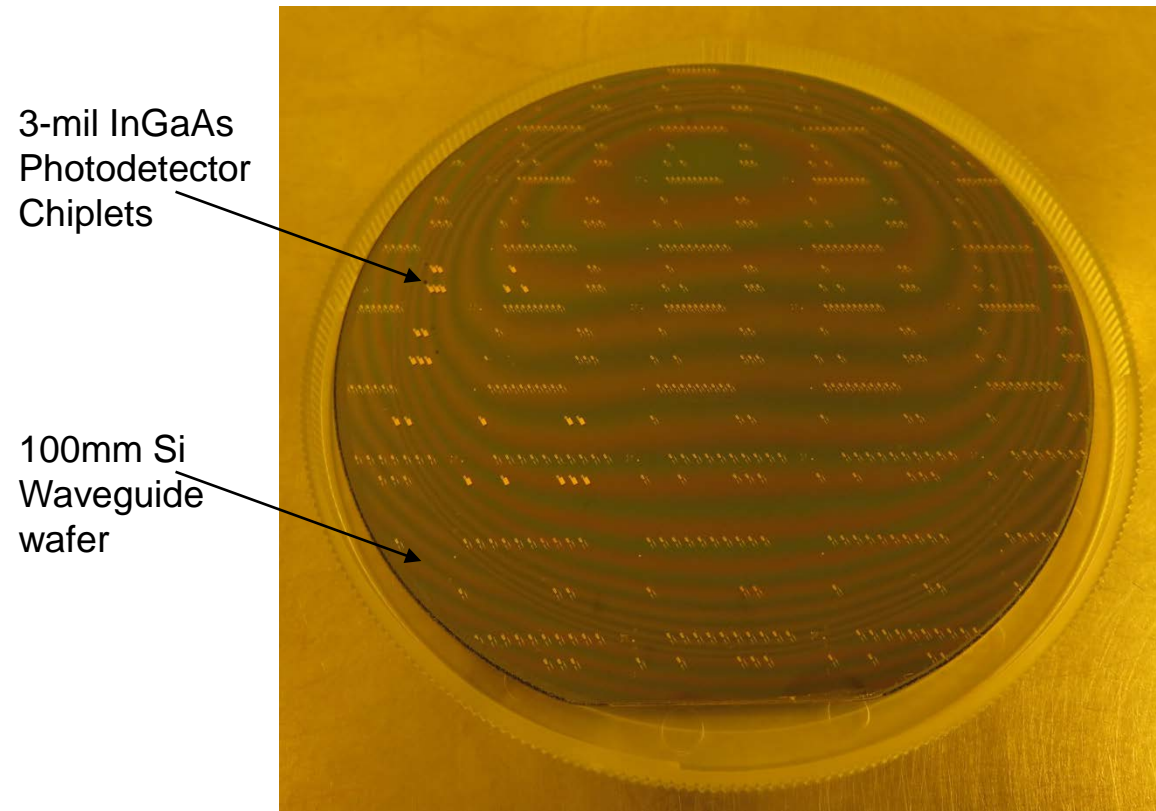
BDD12A-2-1: short-flow minesweeper lot using 100mm GaAs mechanicals

# Detector SEM Image

Detector MMIC



# Initial Integration of Detector on Si Waveguide

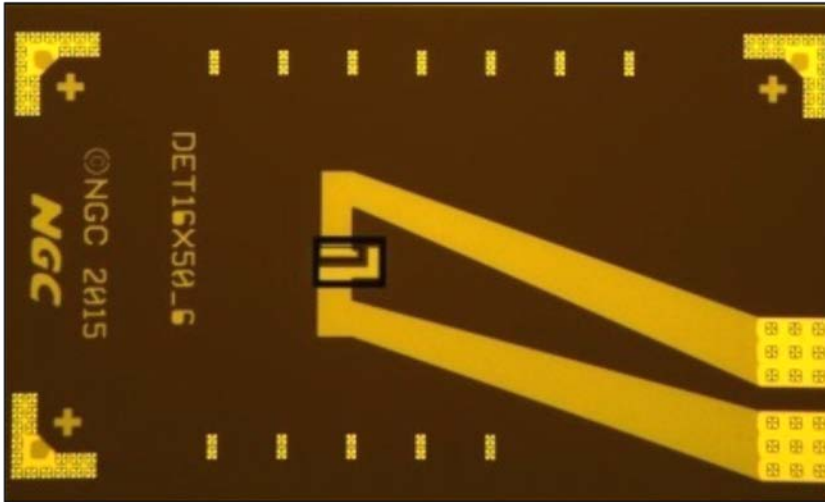


Successfully integrated photodetector chiplets on to the Si waveguide wafer

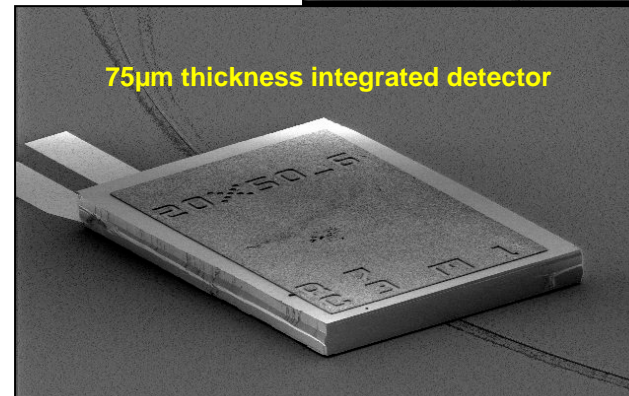
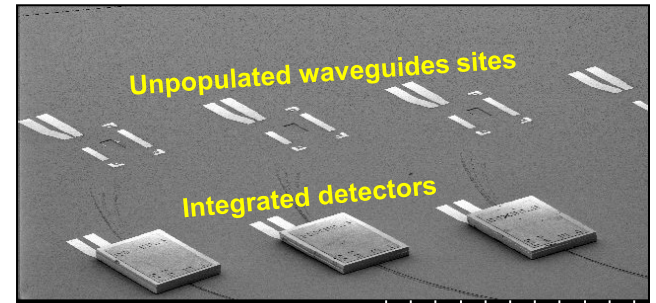
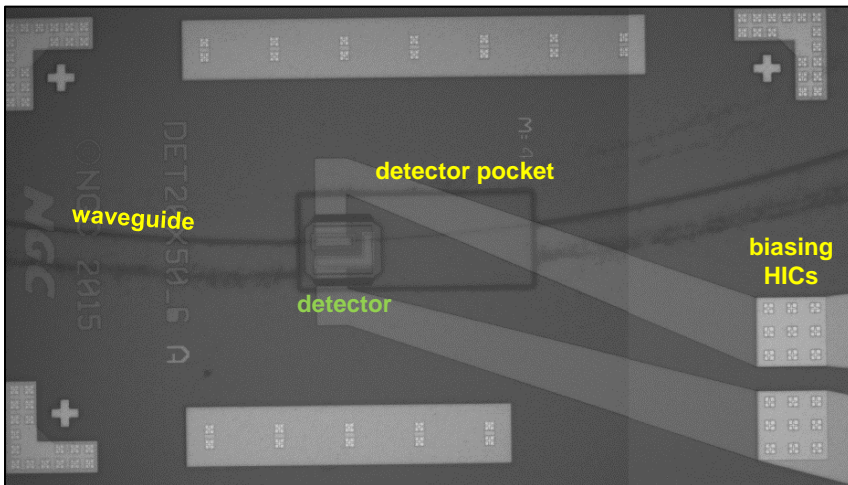


# Photodiode Integration Demonstrations

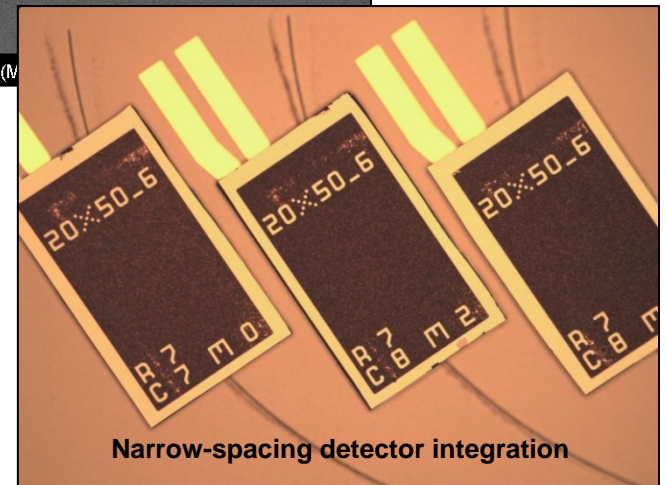
**Photodiode Microscope Image**



**Detector-Waveguide Overlay Image**

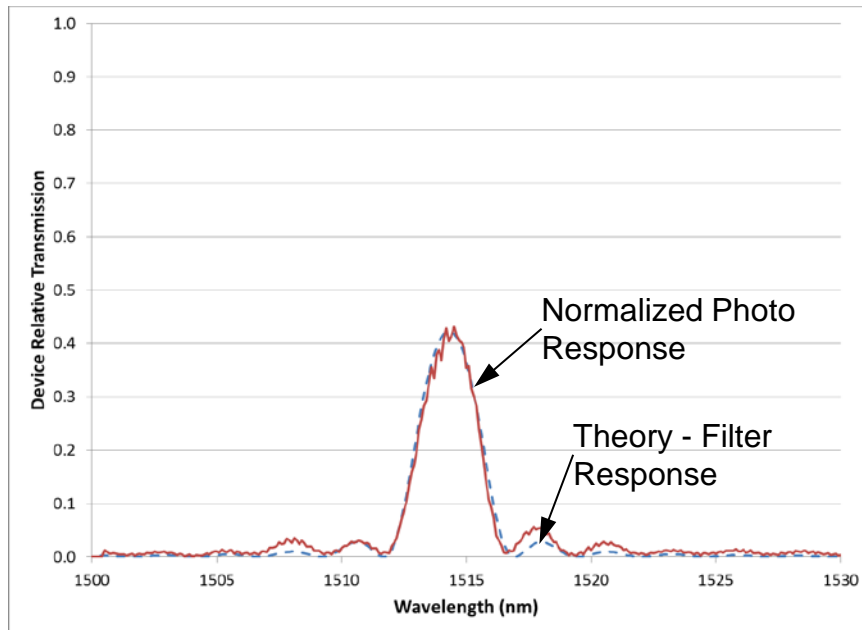
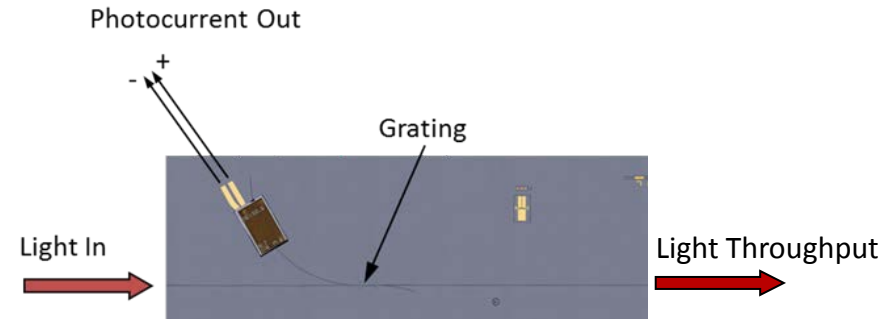
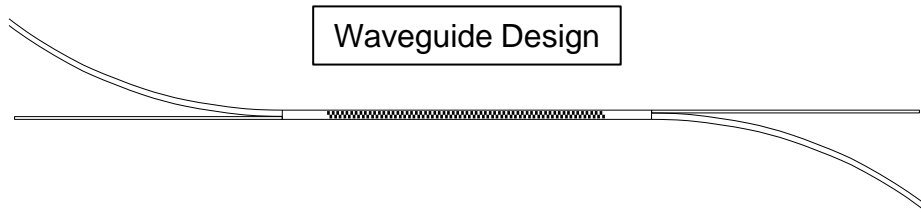


SEM4 2.0kV 44.5mm x100 SE(M)

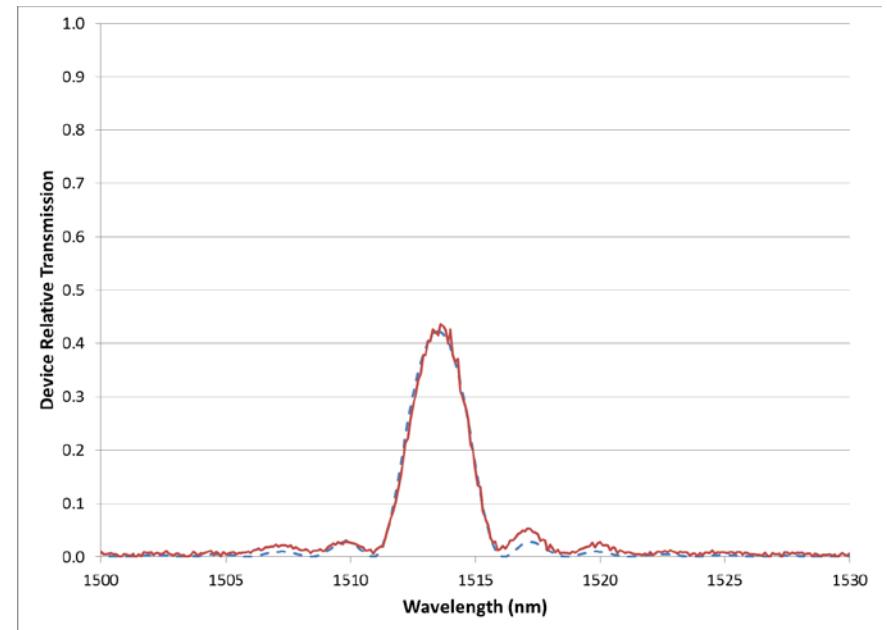


**Narrow-spacing detector integration**

# Filter Response for Single-Filter Devices



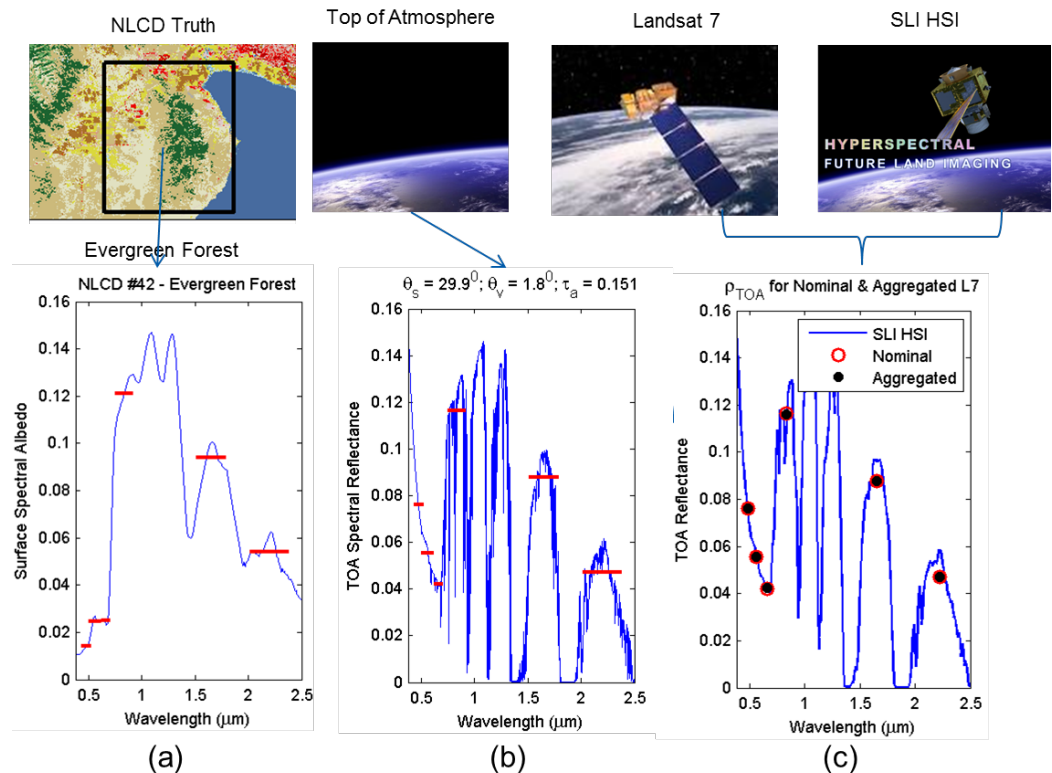
Device W10-v-3-A



Device W10-v-1-A

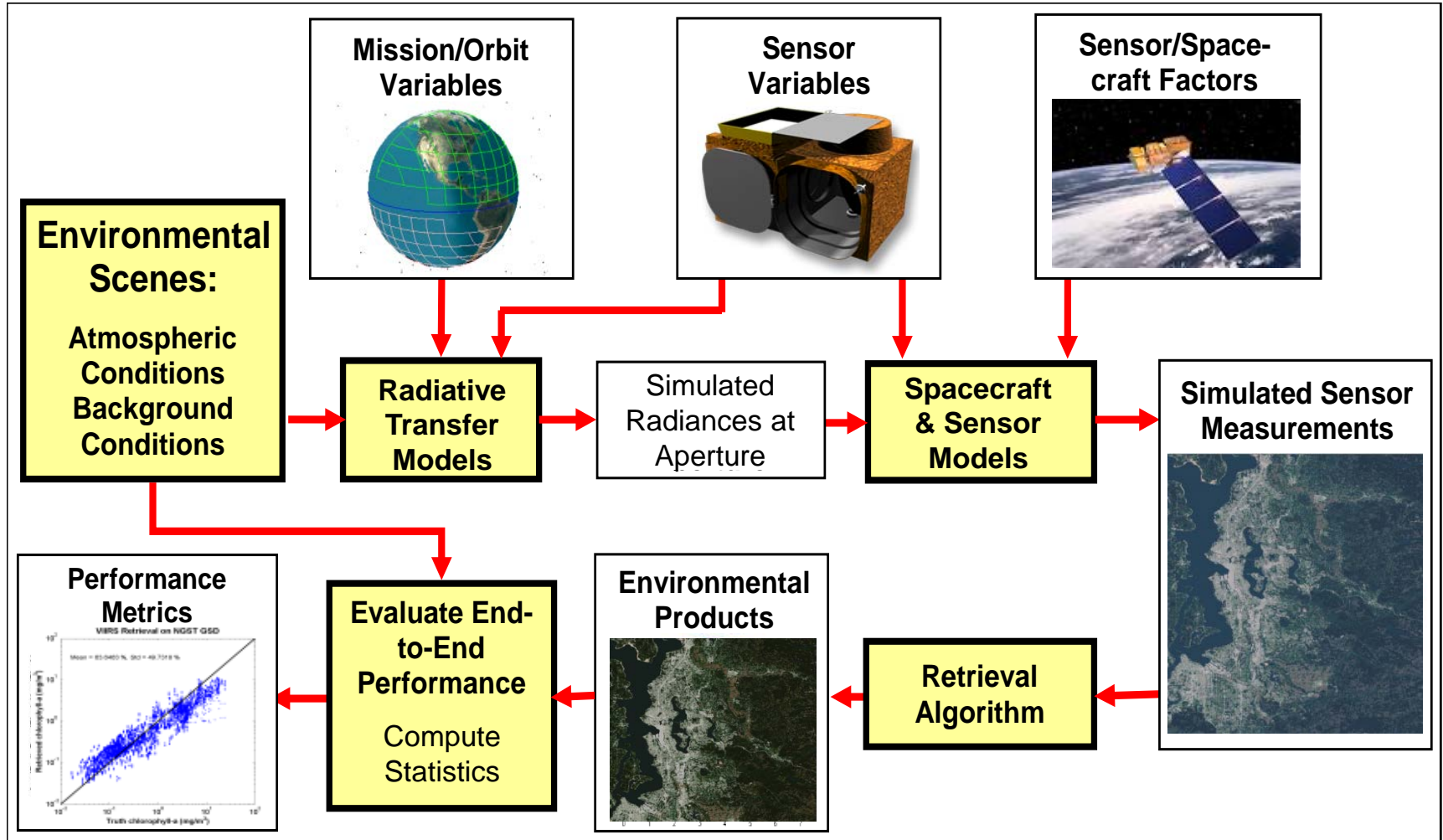
- Study focused on developing designs and performance estimates for key system elements:
  - Telescope
  - Coupling efficiency as a function of wavelength and lenslet focal length
  - Optical throughput and system Signal-to-Noise
  - NGAS band aggregation algorithms

## EVEREST HSI Band Aggregation Simulation



# NGAS Core Mission Evaluation Capability

## Environmental Verification and Remote Sensing Testbed




**EVEREST is a Core Capability to Support Mission Concept Evaluation and System Development**

# EVEREST for Land Imaging


## Basis for High Value End to End Simulations

### Surface




The National Land Cover Database

USGS science for a changing world




USGS Digital Spectral Library 06

<http://speclab.cr.usgs.gov/spectral.lib06>



Jet Propulsion Laboratory  
California Institute of Technology

ASTER Spectral Library



MODIS  
UCSB  
Emissivity  
Library

NASA/GSFC SBRC

### Atmosphere


NCEP Global Data Assimilation System (GDAS)  
P, T, RH, Ozone

NASA MERRA

MODIS 8-Day Aerosol Product Climatology

NGAS  
Cloud Scene Simulation Model  
Clouds

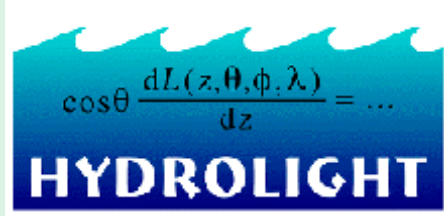
### Radiative Transfer Model



Modtran<sup>5</sup>

Successive Order of Scattering (SOS)

6S(V)



$$\cos\theta \frac{dL(z, \theta, \phi, \lambda)}{dz} = \dots$$

HYDROLIGHT

DISCORD & ISBRDF Snow/Ice

### TOA / SDR Characteristics

Flexible Sensor Observation Geometries

Observation Time Based on Orbit

Spatial Scale Per NLCD

Spectral Resolution < 1nm

Spectral Range ~ 0.38-13um

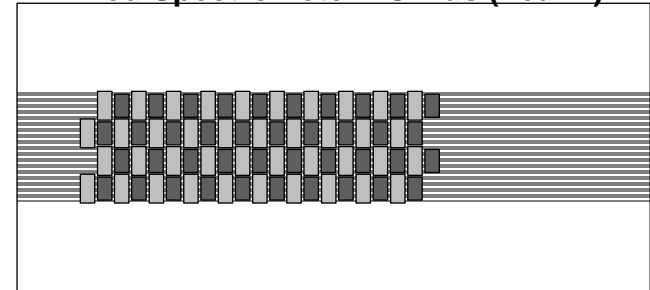
Landsat & HSI Detector RSRs

NGC EVEREST Leverages Many Community Standards

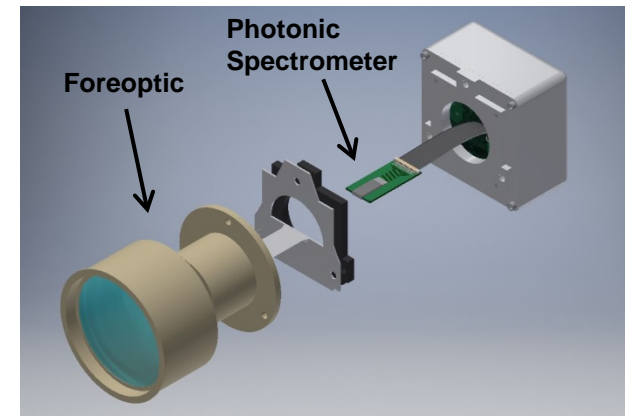
# SLI-T Program

- Currently in first year of a 5-year development program funded by NASA ESTO to build and test a heterogeneously integrated photonic instrument
  - Covers two SLI bands: Band 9 (1.36 – 1.39 $\mu$ m at 3nm resolution) and Band 6 (1.56 – 1.66 $\mu$ m at 6nm resolution)
  - Scalability to SLI VNIR and SWIR bands
  - Integrate NGAS novel ROIC
  - Radiometric performance estimates and testing
- Planned exit TRL = 6

Tiled Spectrometer - Cirrus (Year 1)

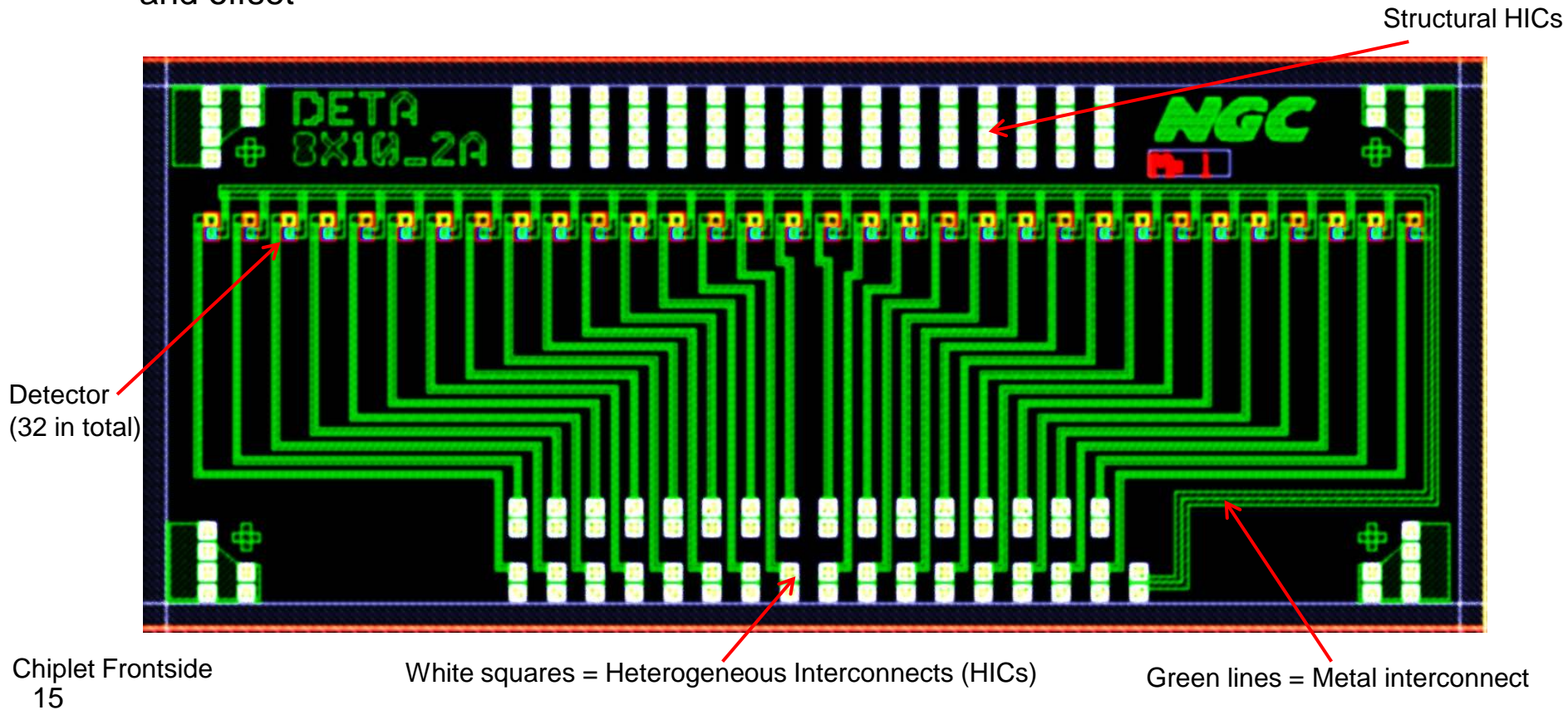


- Spectral range 1360 to 1390 nm
- $\Delta\lambda = 3$  nm, 10 filters per pixel
- # of x-track pixels 128, Pixel spacing 30 $\mu$ m
- Active area: 3.84mm x 13.1mm



# Detector Chiplet Design

- Detector chiplet fabrication initiated
  - Each chiplet contains 32 detectors, metal interconnect, HICs, alignment marks
  - Multiple versions of chiplet per wafer with different detector sizes and offset

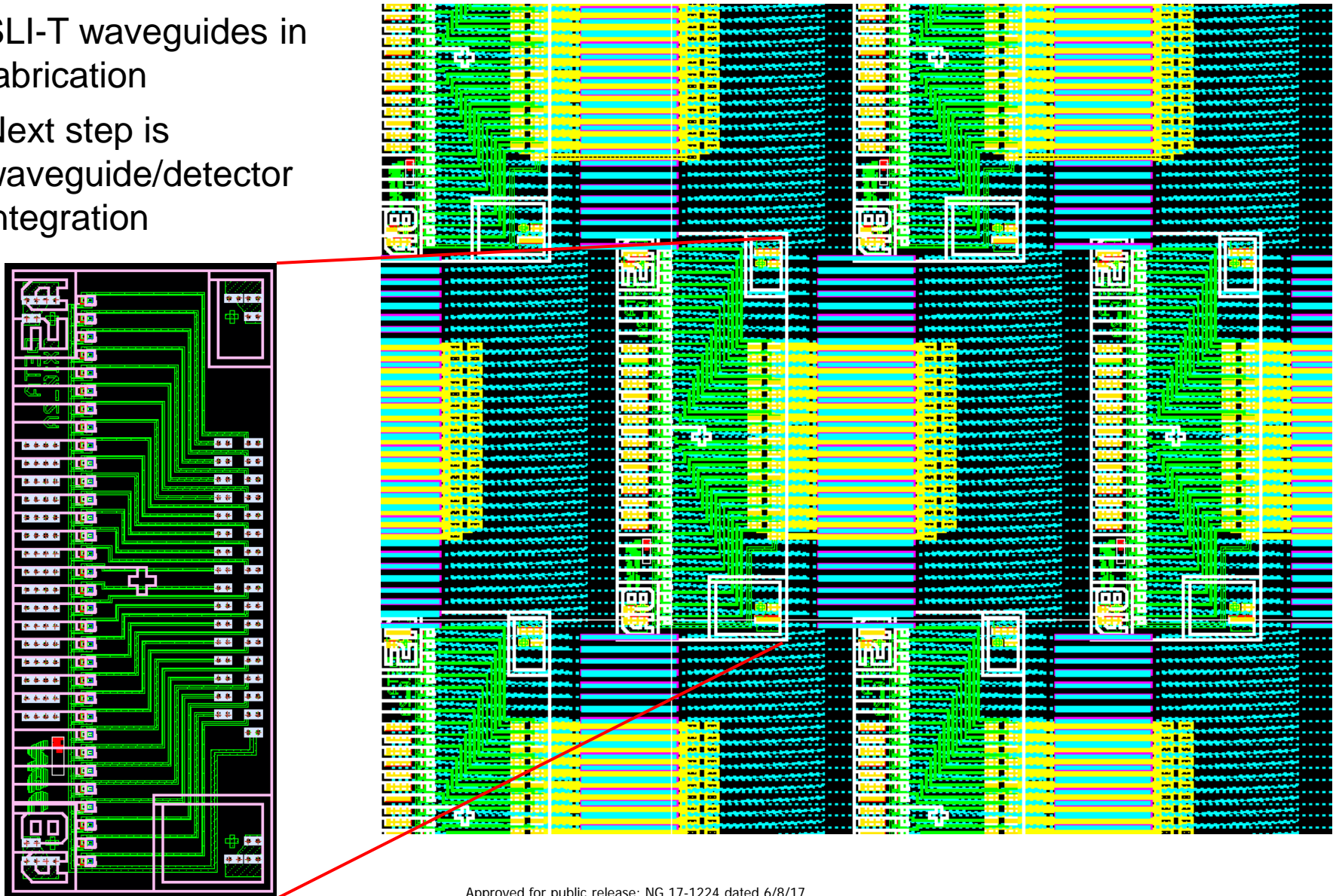


# Detector Tiling Pattern

- SLI-T waveguides in fabrication
- Next step is waveguide/detector integration

Detector Chiplet

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