

Temporal Experiment for Storms and Tropical Systems Technology Demonstration (TEMPEST-D)

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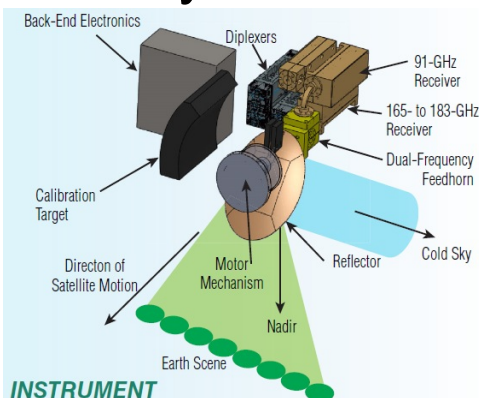
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TEMPEST In-Space Technology Demonstration

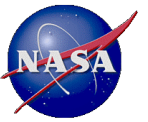


- TEMPEST-D provides first in-space technology demonstration of a millimeter-wave radiometer (90-183 GHz) based on InP HEMT low-noise amplifier front-end aboard a 6U-Class satellite.
- TEMPEST-D raises the instrument TRL from 5 to 7.
- TEMPEST-D is a partnership among CSU, JPL and spacecraft provider (TBD with process in place to determine)
- Managed by NASA ESTO with funding from Earth System Science Pathfinder in Earth Science Division
- Start in July 2015 for three-year technology demonstration task

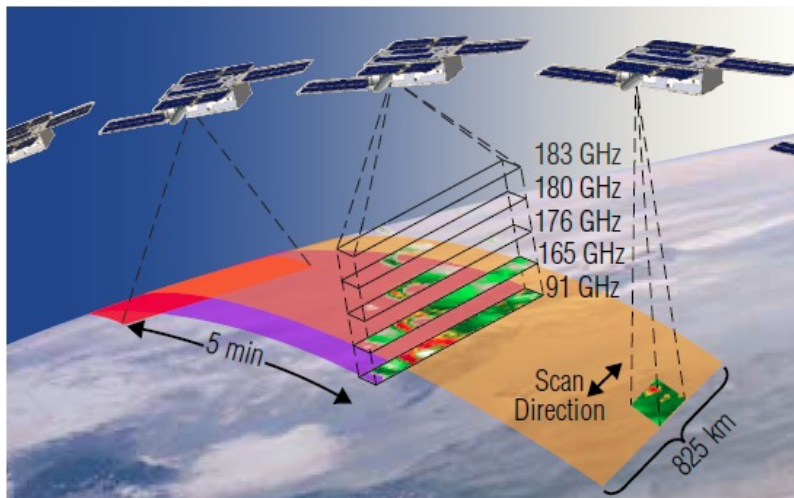




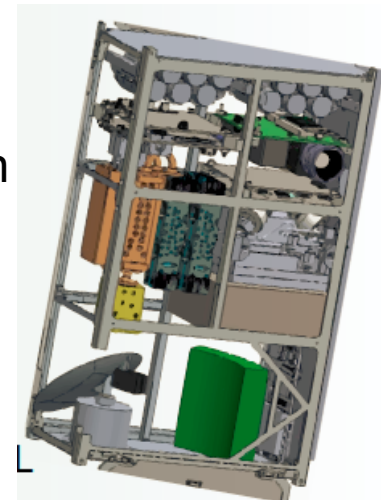
TEMPEST-D Scientific Motivation and Objectives



- Demonstrate capability of U-Class Satellites to provide suitable measurements to advance NASA's Earth Science Goals
- Reduce cost, risk and development time of future constellations of NASA's small satellites to perform Earth Science measurements
- Demonstrate drag-adjusting altitude maneuvers required to provide time separation to a U-Class constellation with common deployment
- Demonstrate precision intercalibration with existing conically-scanning satellite radiometers with similar frequencies within 1-2 K



6U-Class
satellite with
millimeter-
wave
radiometer
instrument





Reference: Global Precipitation Mission Microwave Imager



- Global Precipitation Mission launched Feb. 27, 2014.
- GPM Microwave Imager is a conically-scanning radiometer measuring at 10, 18, 23, 36, 89, 166, 176, and 180 GHz

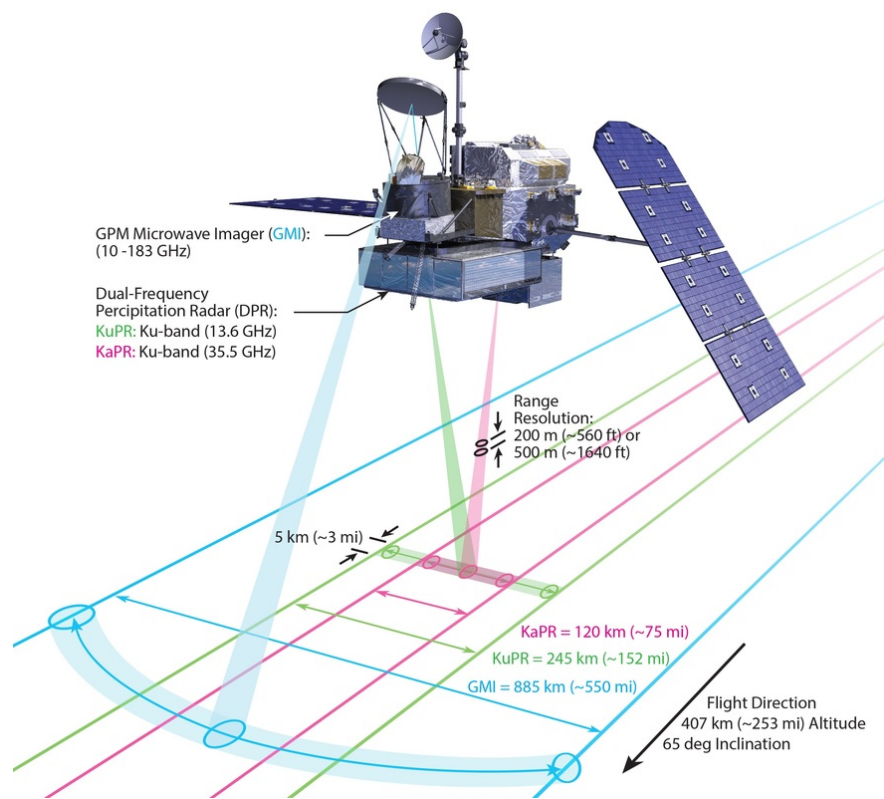


TABLE I
GMI PERFORMANCE REQUIREMENTS

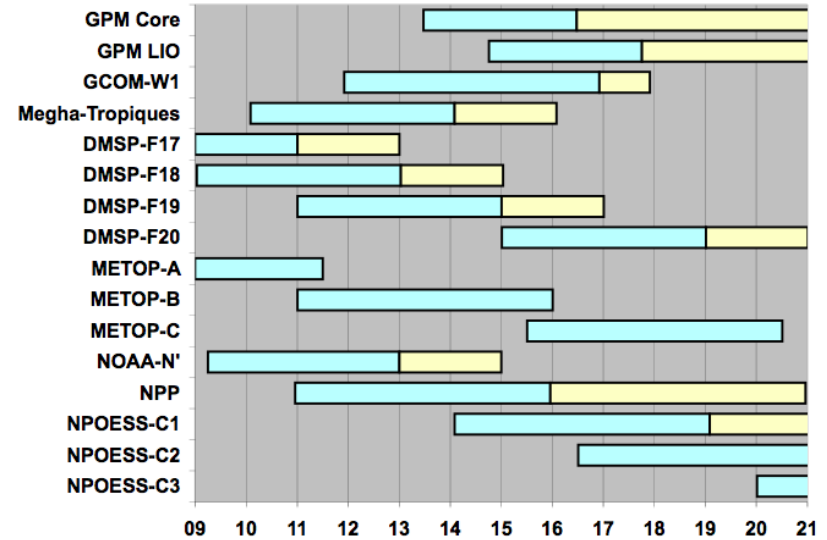
| Ch | Center freq (GHz) | Pol | Freq stability (MHz) | Bandwidth (MHz) | Sampling interval (ms) | NEDT (K) ** |
|-------------------------|----------------------|-----|-----------------------------------|--------------------|------------------------------|-------------------|
| 1 | 10.65 | V | 10 | <100 | 3.6 | <0.96 |
| 2 | 10.65 | H | 10 | <100 | 3.6 | <0.96 |
| 3 | 18.7 | V | 20 | <200 | 3.6 | <0.84 |
| 4 | 18.7 | H | 20 | <200 | 3.6 | <0.84 |
| 5 | 23.8 | V | 20 | <400 | 3.6 | <1.05 |
| 6 | 36.64* | V | 50 | <1000 | 3.6 | <0.65 |
| 7 | 36.64* | H | 50 | <1000 | 3.6 | <0.65 |
| 8 | 89 | V | 200 | <6000 | 3.6 | <0.57 |
| 9 | 89 | H | 200 | <6000 | 3.6 | <0.57 |
| 10 | 166 | V | 200 | <4000 | 3.6 | <1.5 |
| 11 | 166 | H | 200 | <4000 | 3.6 | <1.5 |
| 12 | 183.31 ± 3 | V | 200 | <2000 | 3.6 | <1.5 |
| 13 | 183.31 ± 7 | V | 200 | <2000 | 3.6 | <1.5 |
| Calibration uncertainty | | | <1.35K (Ch 1–9); <1.5K (Ch 10–13) | | | |
| Calibration stability | | | <0.5K (All Ch's) | | | |
| Dynamic range | | | 3 to 340K (All Ch's) | | | |
| Nonlinearity | | | <0.5K (All Ch's) | | | |

Online at www.spaceflight101.com

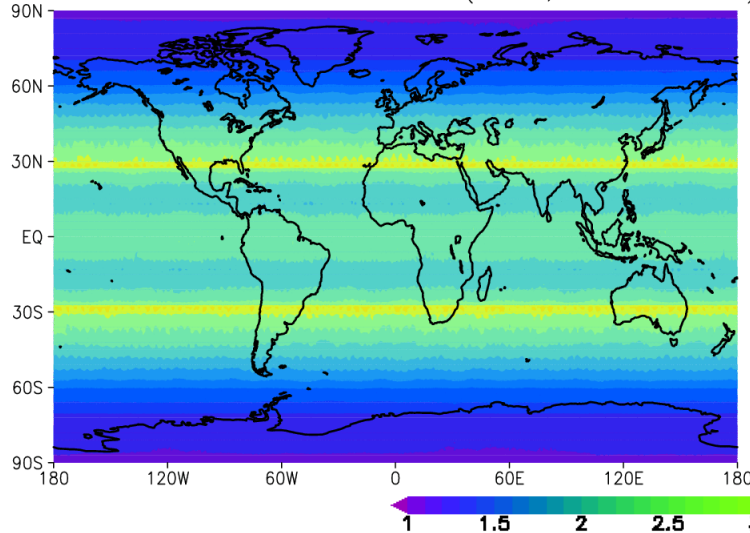
Draper et al., *IEEE J-STARS*, in press, 2015.



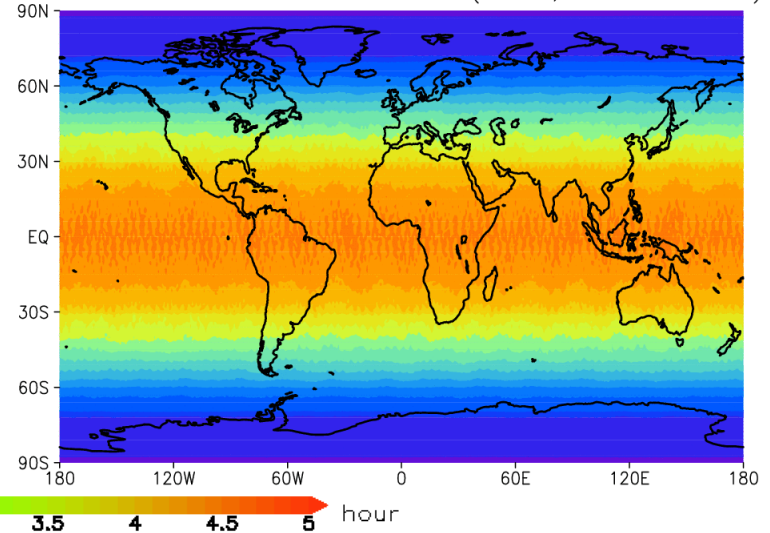
Reference: Sampling/Coverage of GPM Constellation



GPM Constellation Revisit Time (2017, without GM12)

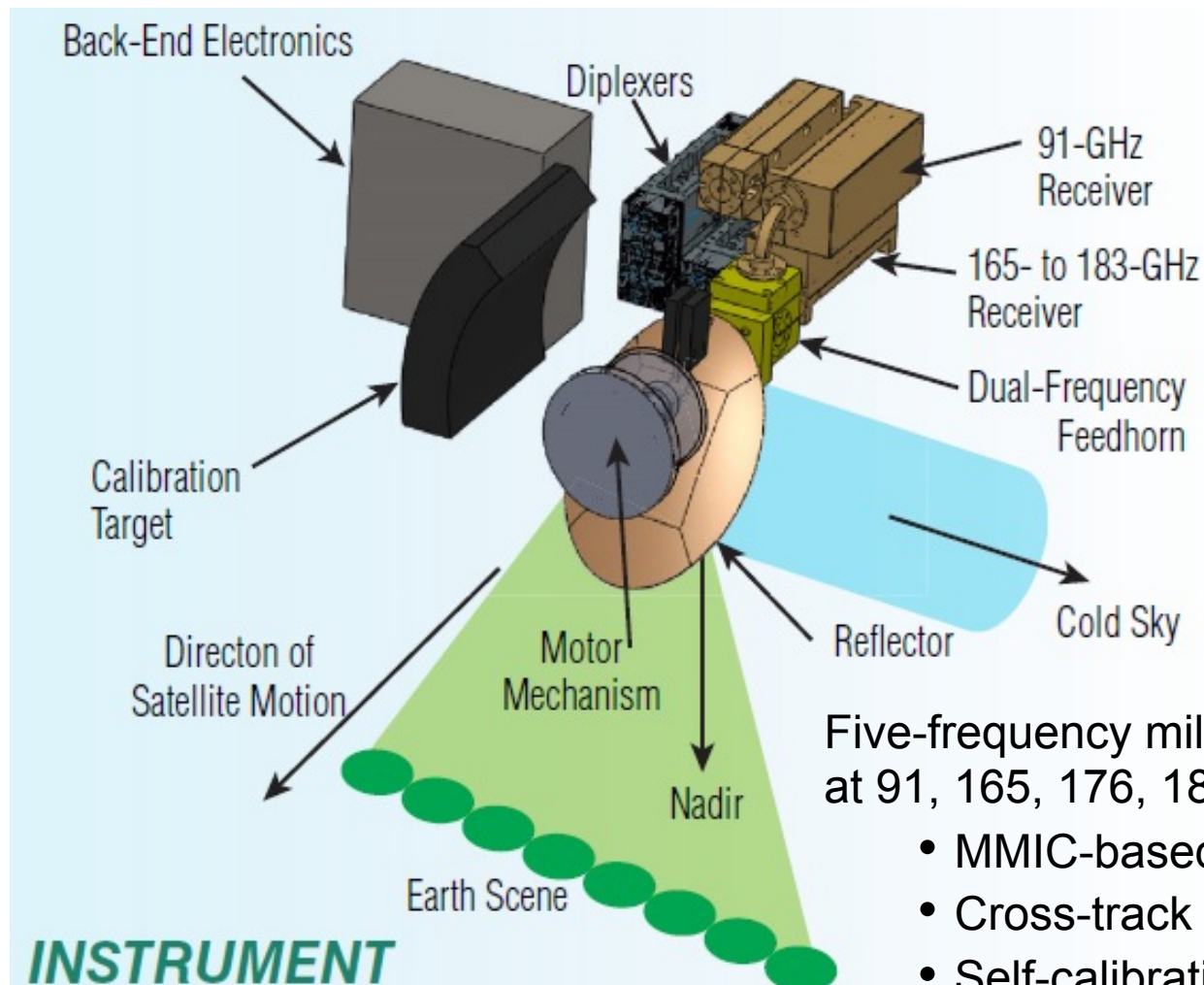


GPM Constellation Revisit Time (2018, without GM12)





TEMPEST-D Instrument Design

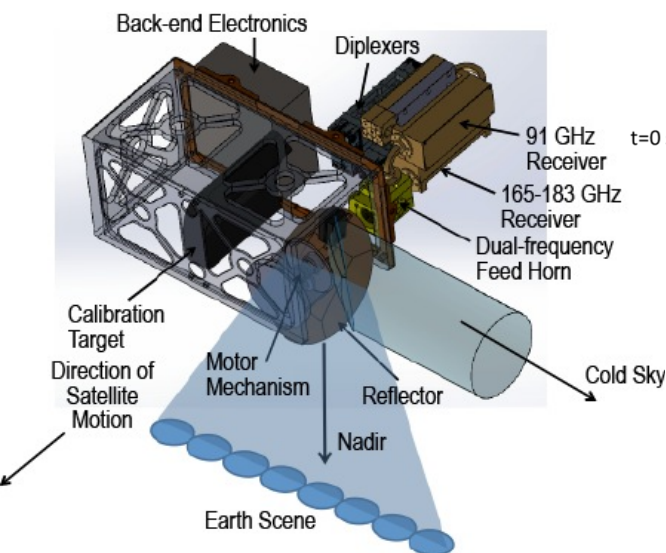


Five-frequency millimeter-wave radiometer at 91, 165, 176, 180 and 183 GHz

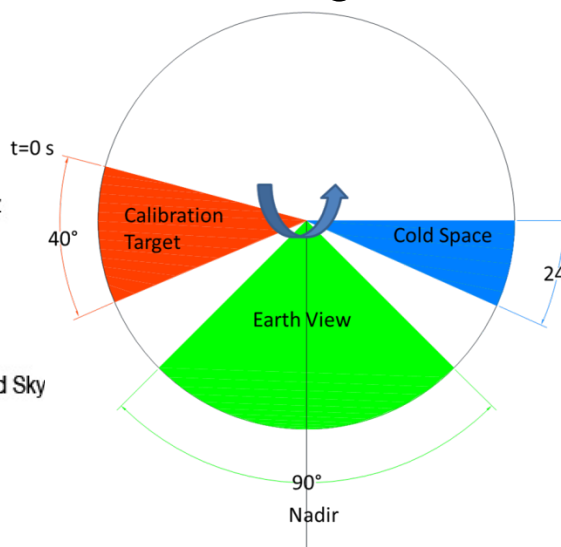
- MMIC-based
- Cross-track scanning
- Self-calibrating



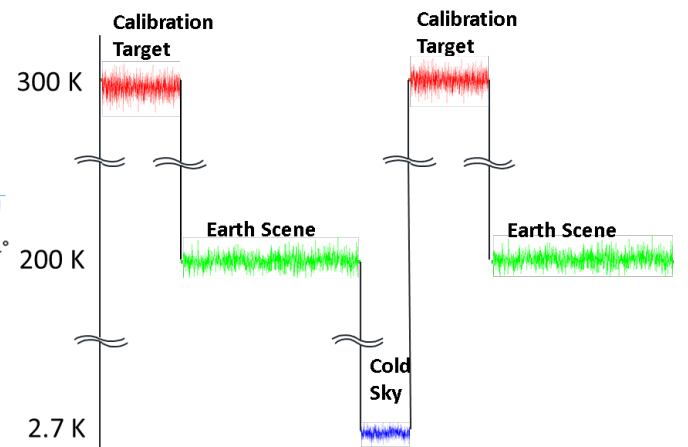
TEMPEST-D Instrument



Observing Profile



Time Series of Output Data



- Five-frequency millimeter-wave radiometer measures Earth scene over $\pm 45^\circ$ incidence angles from nadir.
- This provides an 825-km swath width from a nominal altitude of 400 km.
- Each pixel is sampled for 5 ms.
- Space view observes cold sky, i.e. cosmic microwave background at 2.7 K.
- Ambient calibration target provides the second measurement for two-point calibration to determine instrument gain and offset.



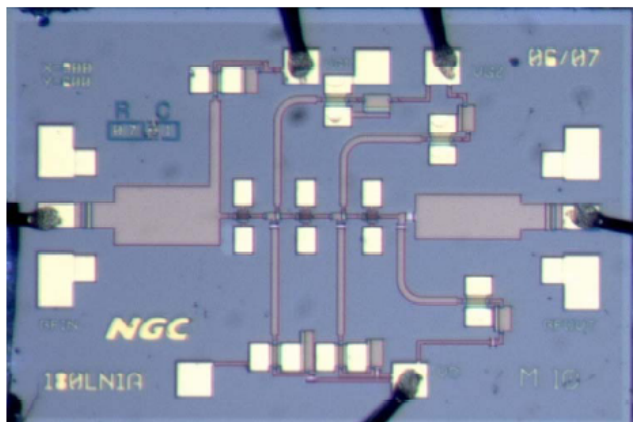
TEMPEST-D Instrument Specifications



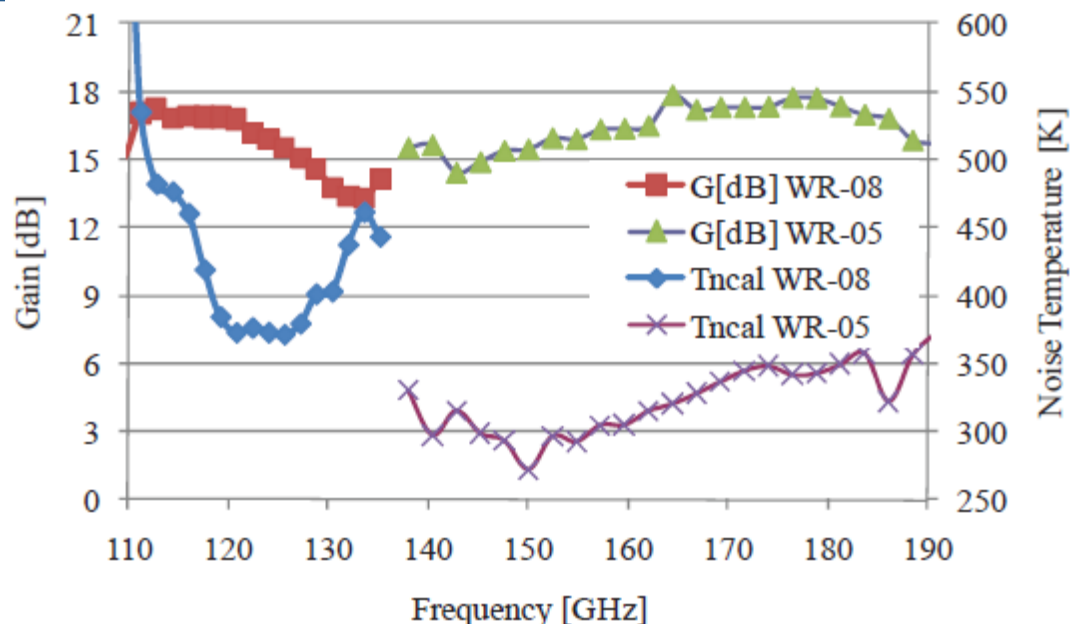
| Parameter | Specification | |
|----------------------------|---|-----------------|
| System noise temperature | < 600 K | |
| Number of channels | 5 | |
| Bandwidth | 4 GHz at 91 and 166 GHz; 2 GHz at 173, 180 and 183 GHz | |
| Minimum spatial resolution | 13 km at 183 GHz | 24 km at 91 GHz |
| Minimum beam efficiency | > 90% | > 90% |
| Mass | < 3 kg | |
| Power | < 8 W | |
| Volume | 3U | |
| Precision (Stability) | 0.5 K (CBE) | |
| Accuracy (Calibration) | 1 K (CBE) | |



Enabling Technology: 35-nm InP HEMT Low-Noise Amplifiers



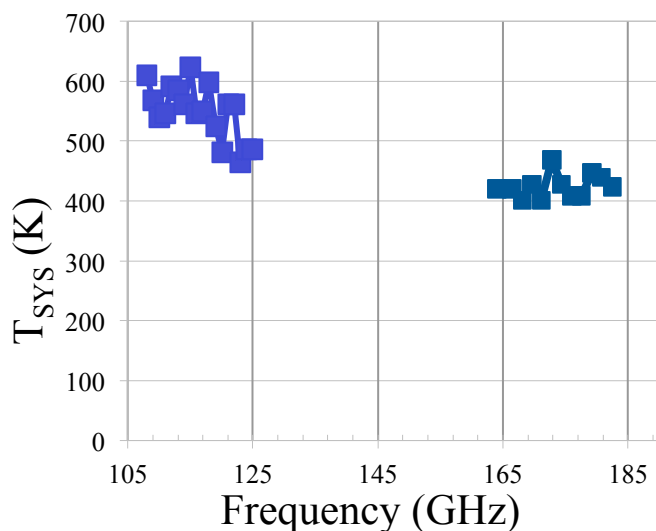
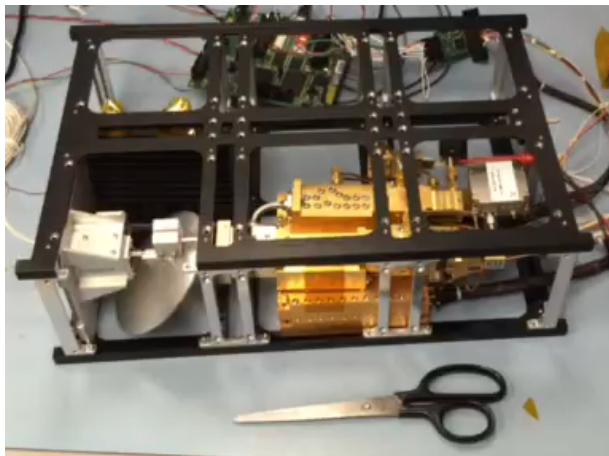
From Kangaslahti et al.,
Proc. Int'l Microwave Symp. 2012



- 35-nm InP HEMT process
- Record low noise temperature of 300 - 350 K from 140 - 190 GHz (at room temperature)
- Chip area of 900 x 560 μm^2
- Three-stage design with separate gate bias for the first stage to optimize low-noise performance
- The LNA was mounted in optimized WR-08 and WR-05 waveguide housings for testing over a broad bandwidth.



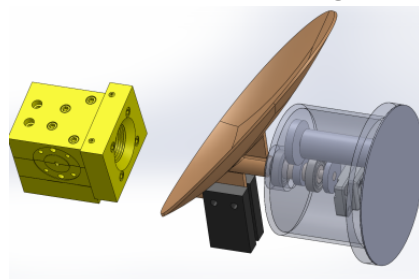
Microwave Atmospheric Sounder on CubeSat (MASC)



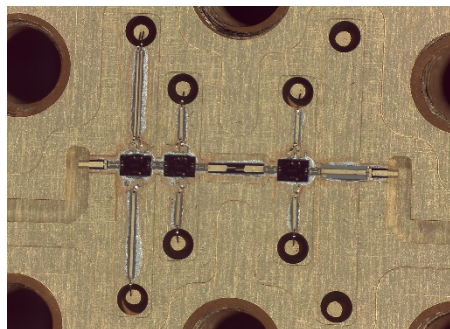
- Prototype millimeter-wave sounding radiometer for 6U CubeSat; JPL R&TD led by Sharmila Padmanabhan
- Instrument occupies 3U of a 6U CubeSat, with mass of 3 kg and power consumption of 6.5 W
- Temperature and water vapor profiling using 4 channels each near absorption lines at 118.75 GHz and 183.31 GHz oxygen and water vapor absorption lines, respectively
- Receiver noise temperatures were measured using standard gain horns.
- Heritage from NASA ESTO-funded ACT, IIP and AITT programs as well as JPL RACE CubeSat mission.

MASC Instrument Architecture and Hardware Heritage

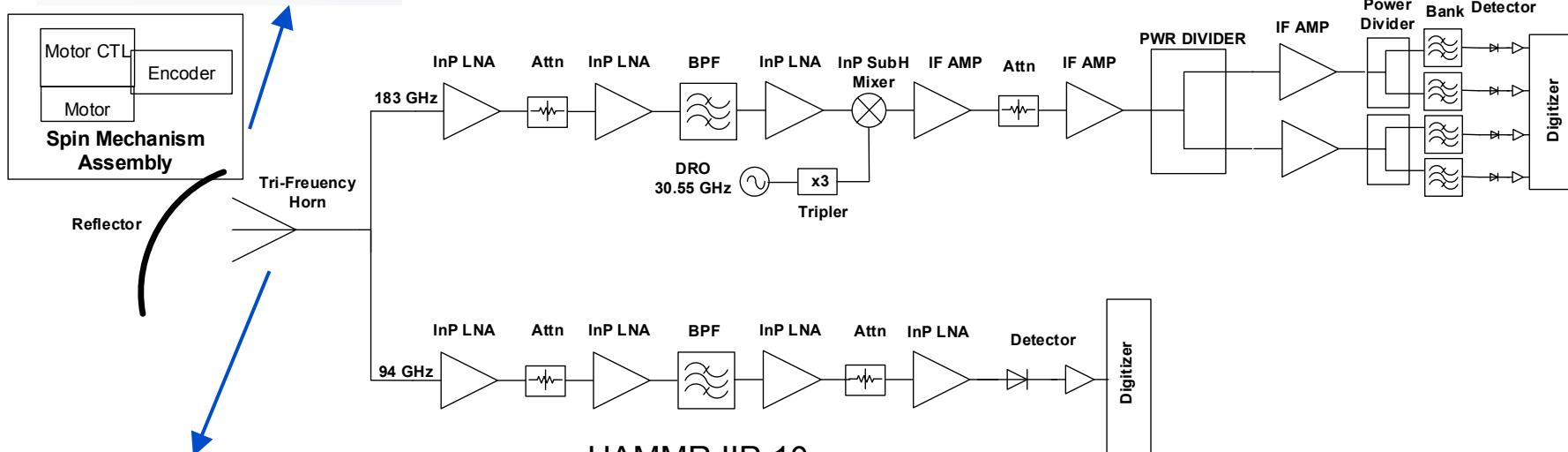
JPL R&TD Offset Paraboloid Reflector design



MASC 183 GHz front-end



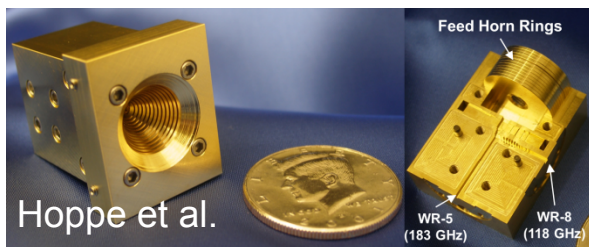
RACE Digitizer



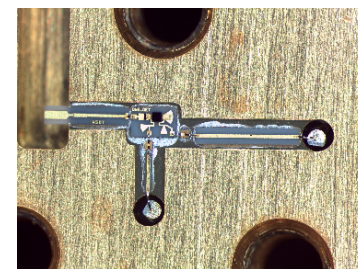
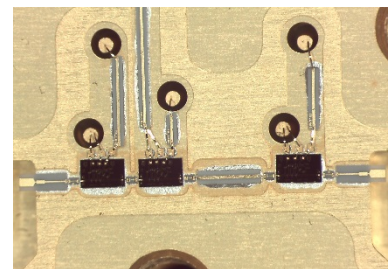
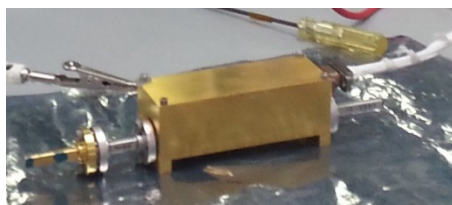
HAMMR IIP-10

ACT-08 (Reising, Hoppe, Kangaslahti) (Reising et al.; CSU+JPL)

MASC 118 GHz front-end IIP 90 GHz Detector



Hoppe et al.





TEMPEST-D Selected in Feb. 2015 by NASA CubeSat Launch Initiative



- TEMPEST-D was selected by NASA CSLI (ELaNa) in Feb. 2015
- Assigned priority #7 of 14 (based on predicted launch readiness, launch priority orbital requirements, etc.)

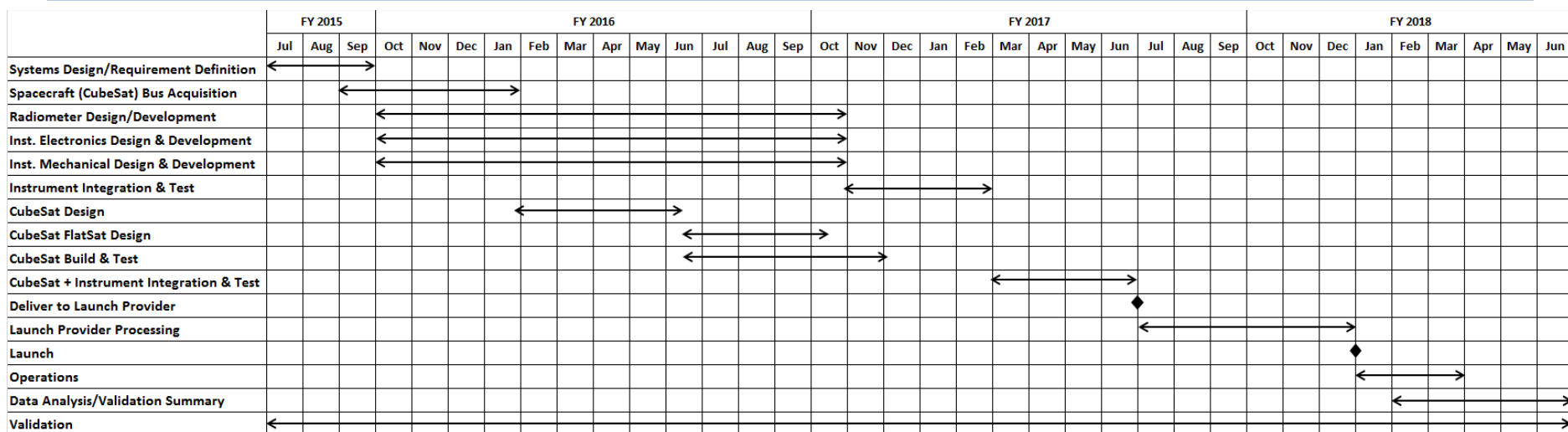
| February 2015 Selection | | | | | | |
|-------------------------|------------|---|----------------------------|---------------------|-----------|---------|
| CSLI Rank | Short Name | Project | Organization | POC | Cube Size | Status |
| 1 | OPEN | OPEN: Open Prototype for Educational NanoSats | University of North Dakota | Ronald Marsh | 1U | In Work |
| 2 | RadSat | RadSat: Satellite Demonstration of a Radiation Tolerant System | Montana State University | Dr. Brock LaMeres | 3U | In Work |
| 3 | ALBus | Advanced ELectrical Bus (ALBus) | NASA Glenn Research Center | Katie Shaw | 3U | In Work |
| 4 | MiTEE | MiTEE: Miniature Tether Electrodynamics Experiment | University of Michigan | Dr. Brian Gilchrist | 3U | In Work |
| 5 | KickSat-2 | KickSat-2: A Technology Demonstration Mission for the Sprite ChipSat | Cornell University | Zachary Manchester | 3U | In Work |
| 6 | TBEx | TBEx – The Tandem Beacon Experiment | The University of Michigan | Dr. James Cutler | 3U | In Work |
| 7 | TEMPEST-D | TEMPEST-D: Temporal Experiment for Storms and Tropical Systems - Demonstrator | Colorado State University | Dr. Steven Reising | 6U | In Work |

From https://www.nasa.gov/directorates/heo/home/CSLI_selections.html#2015

TEMPEST-D Key Roles

| Name | Role | Institution |
|----------------------|--|---------------------------------|
| Steven Reising | Principal Investigator | Colorado State University (CSU) |
| Todd Gaier | JPL Lead Co-I | Jet Propulsion Laboratory (JPL) |
| Ron Steinkraus | Project Manager; Spacecraft and Mission Operations | JPL |
| V. Chandrasekar | CSU Co-I; Validation / Engineering Interface | CSU |
| Sharmila Padmanabhan | Instrument Engineer | JPL |
| Boon Lim | Systems Engineer | JPL |
| Christian Kummerow | Validation Lead | CSU |
| Ted Sweetser | Orbital Architecture | JPL |

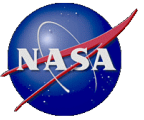
TEMPEST-D Schedule



- July 2015: Project Start
- Nov. 2016: Deliver radiometer instrument for integration & test
- Dec. 2016: Complete 6U-Class satellite bus
- July 2017: Deliver TEMPEST-D spacecraft to launch provider
- July 2017 to Jan. 2018: 6-month window for launch availability
- After launch: 3 months of operations; 3 months analysis & validation
- Jun. 2018: Project Completion

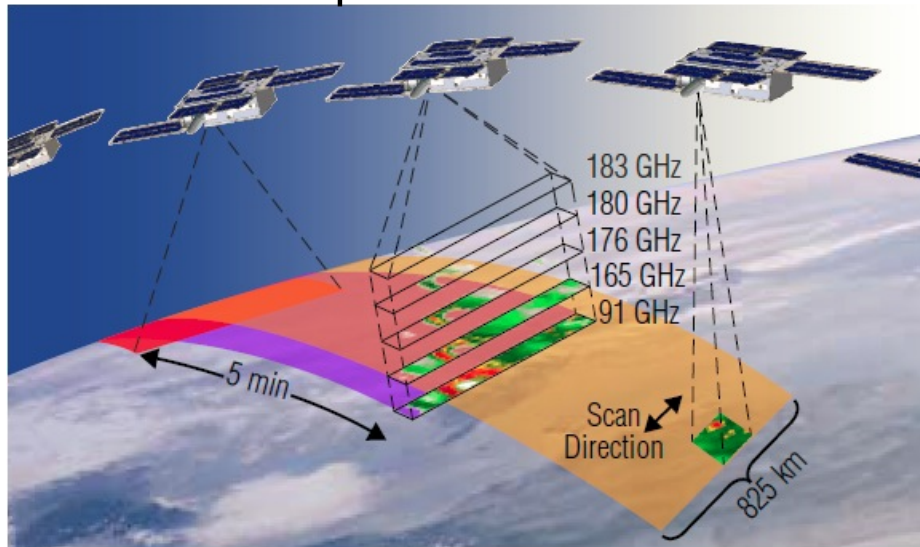


Temporal Experiment for Storms and Tropical Systems (TEMPEST)

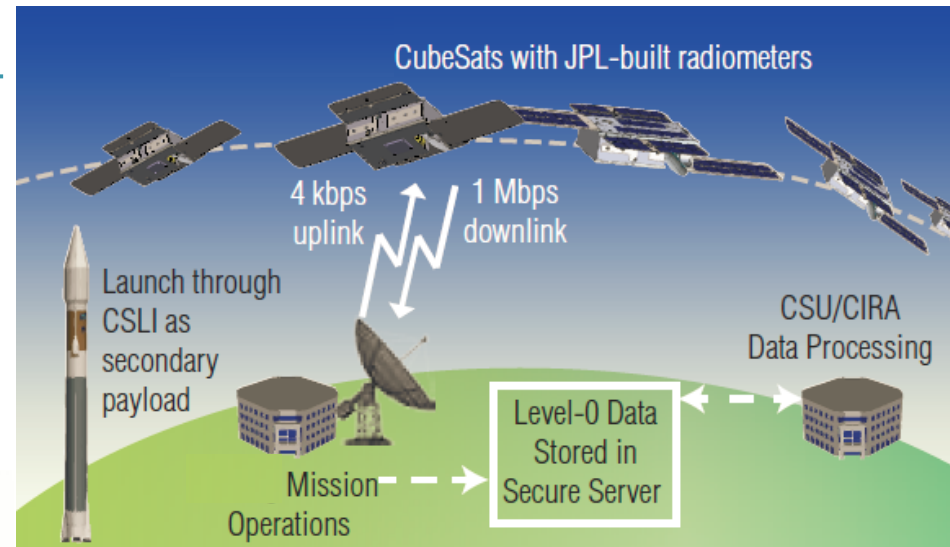


BENEFITS AND STRENGTHS

- First global observations of time evolution of precipitation
- Low-cost approach and rapid development using 6U CubeSats
- Unique data sets to improve weather and climate prediction models



5 identical 6U CubeSats, each with an identical 5-channel radiometer, flying 5 minutes apart

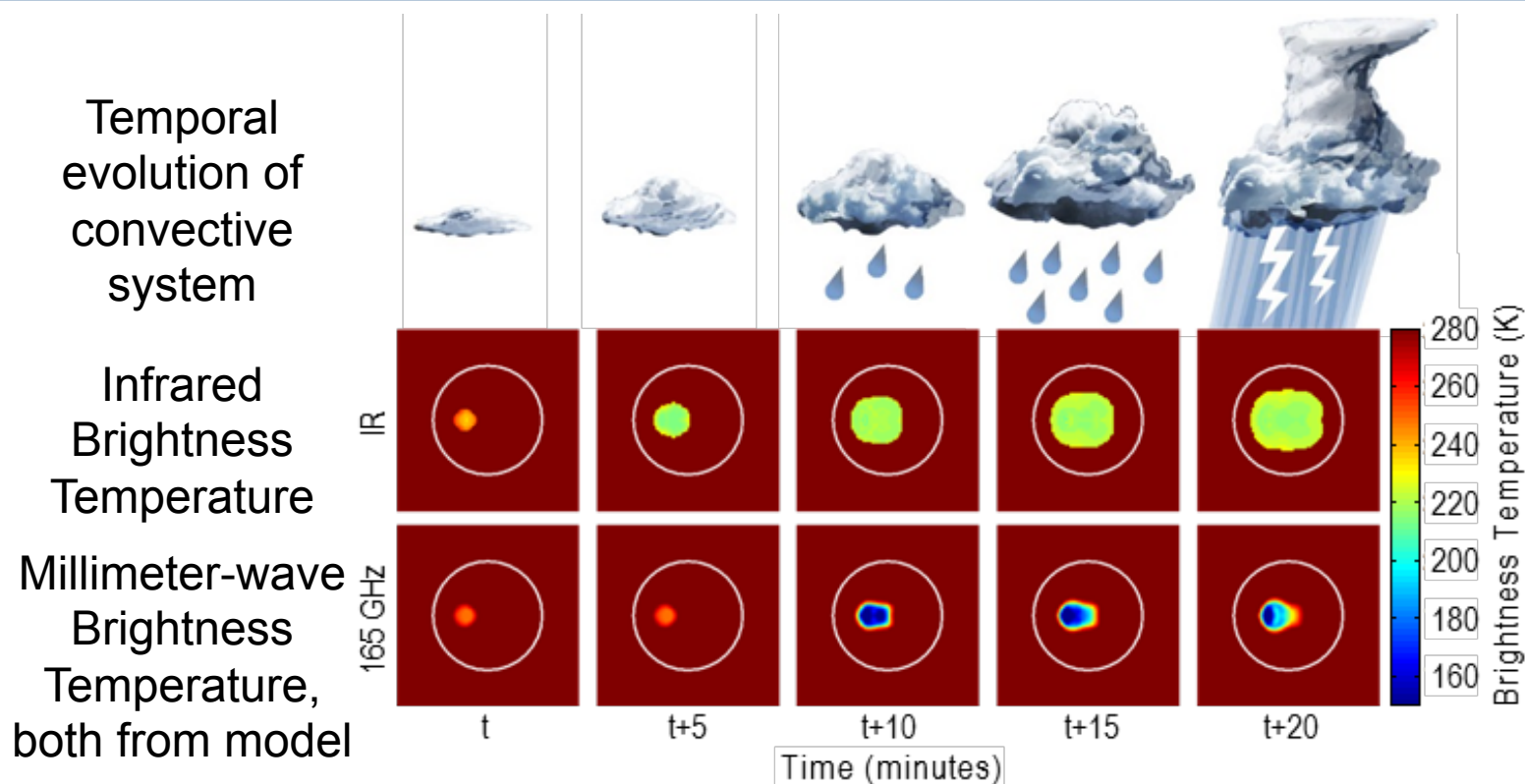


IMPORTANCE TO NASA

- Constrain climate models through improved understanding of cloud processes, transition from clouds to precipitation and impact on Earth's energy balance
- Characterize temporal variability of precipitation globally to improve understanding of water cycle



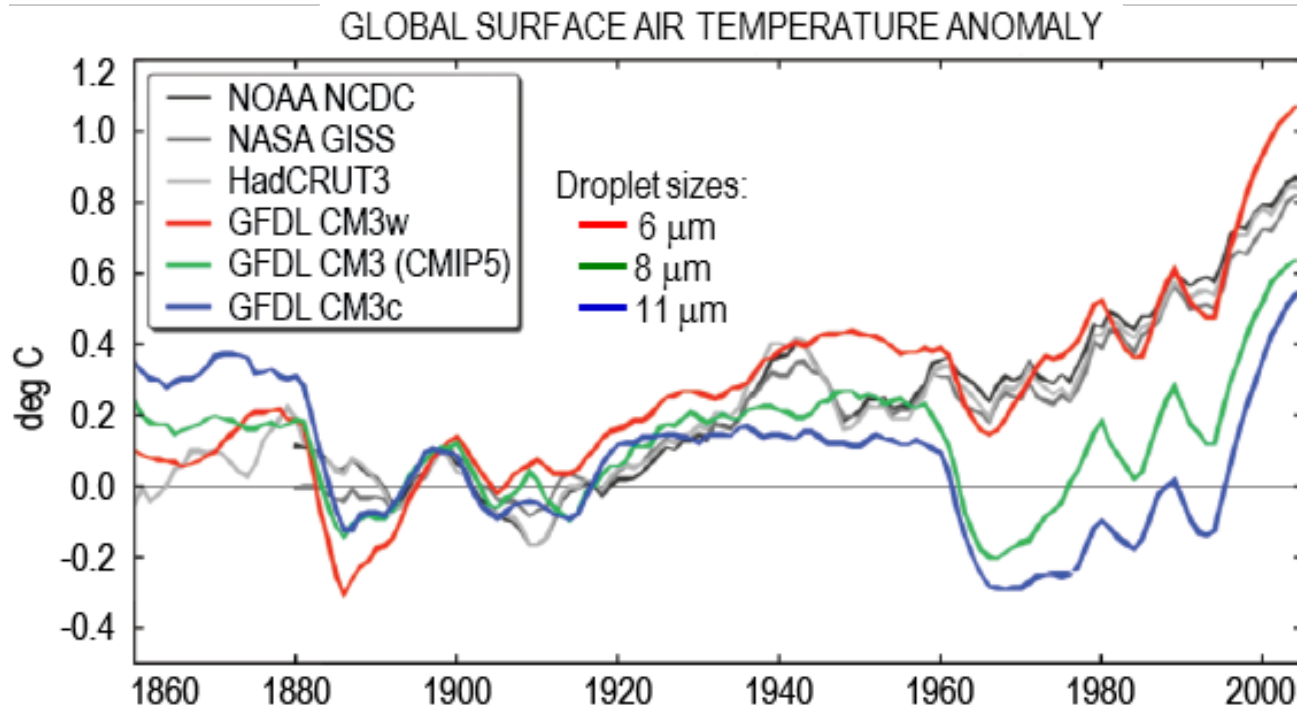
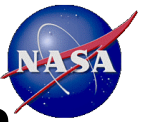
Future Vision: Observation of Transition from Clouds to Precipitation



- Infrared brightness temperatures (GEO) show only cold cloud top temperatures.
- Transition from clouds to precipitation is clearly detected at millimeter-wave frequencies on TEMPEST constellation, including 165 GHz.
- TEMPEST spatial resolution of 25 km at lowest frequency is shown (circles).



Sensitivity of Climate Model Predictions to Onset of Precipitation

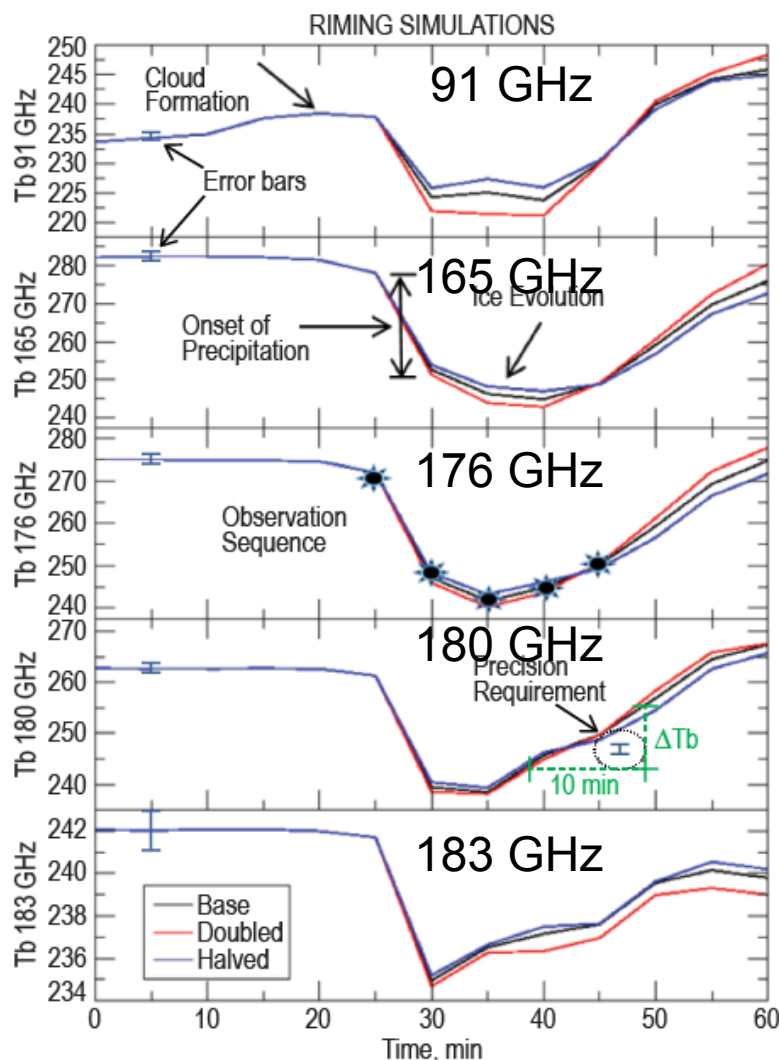


[Adapted from Golaz et al., *GRL*, 2013; Suzuki et al., *GRL*, 2013]

- Global climate model temperature anomaly projections for onset of precipitation at cloud droplet sizes of 6 μm (red), 8 μm (green) and 11 μm (blue).
- TEMPEST constellation could provide the first global sample of the onset of precipitation, constraining climate prediction models.



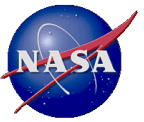
Temporal Development of Ice in Cloud-Scale Models



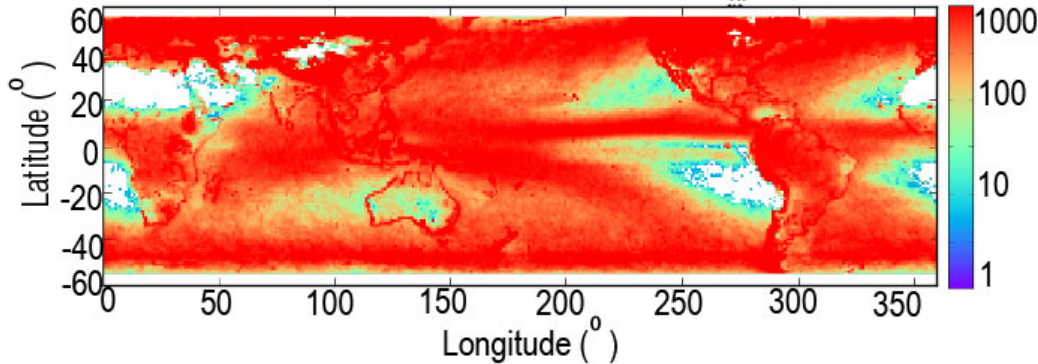
- Modeled brightness temperatures at the five TEMPEST frequencies with 25-km spatial resolution
- Simulations compare different rates of supercooled water droplets collecting on ice crystals (riming efficiency).
- Efficiency (rate) varies from baseline (black) to twice (red) and half (blue).
- Measurable difference between curves is 4 K or greater in 5 minutes at onset of ice formation. Precision requirement is 1 K in 5 minutes.
- Ice remaining in clouds after precipitation can have significant effects on climate system. Residual ice can be compared to W-band radar observations from CloudSat or ESA's EarthCARE.



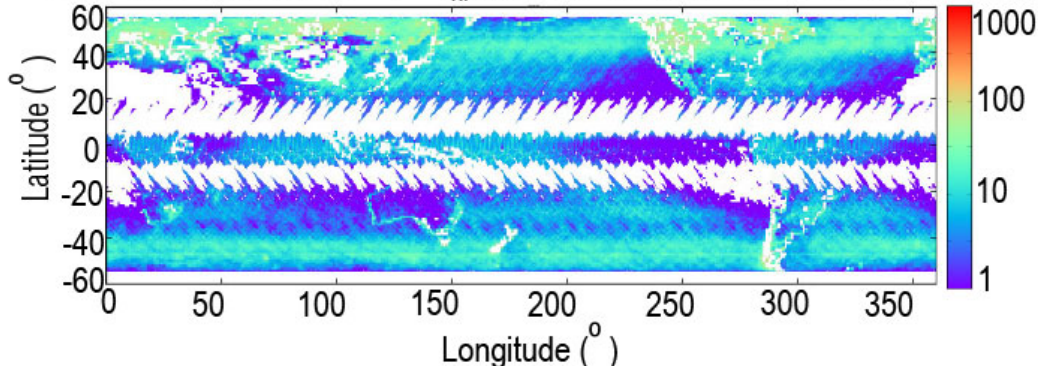
Global Observations of Temporal Evolution of Precipitation



TEMPEST Number of Rain Events $> 1 \text{ mm/hr}$ in each $1^\circ \times 1^\circ$



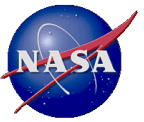
TEMPEST Rain Events $> 1 \text{ mm/hr}$ in each $1^\circ \times 1^\circ$, seen by GPM within 30 min



- In a potential one-year mission, TEMPEST constellation could make more than 3,000,000 temporal observations of precipitation ($> 1 \text{ mm/hr}$), including 100,000+ deep-convection events
- Could perform more than 50,000 coincident precipitation observations within 30 minutes of NASA's Global Precipitation Mission (GPM) for a nominal TEMPEST orbit for ISS launch at 400-km altitude and 51° inclination.
- Precipitation estimates from AMSR-E satellite radiometer data with oceanic observations only.



TEMPEST 6U-Class Satellites and Passive Constellation Maneuvers



KEY FLIGHT CHARACTERISTICS

- 5 identical 6U CubeSats

Attitude:

3-axis stabilization

0.13° (1 σ) control

0.15° (1 σ) knowledge

Mass:

5.8 kg (Margin: 38%)

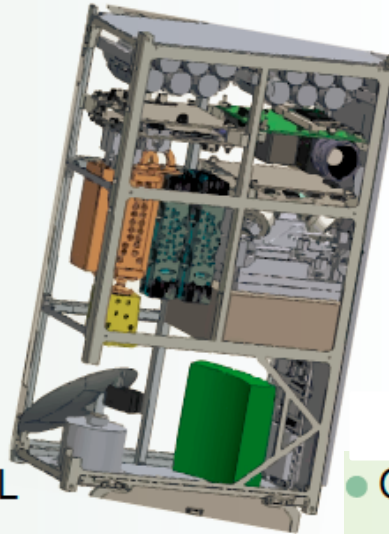
Power:

13 W (Margin: 23%)

Peak Power: 65 W EOL

Communications:

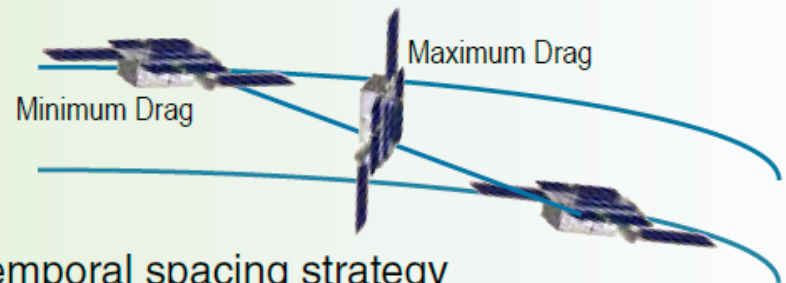
1 Mbps S-band (Margin: 22%)



- Orbital characteristic (CSLI compatible)

Altitude: 390 – 450 km

Inclination: 50°– 65°



- Temporal spacing strategy
Drag-adjusting attitude maneuvers used to achieve temporal separation between CubeSats

Summary



- TEMPEST-D provides first in-space technology demonstration of a 90-183 GHz millimeter-wave radiometer based on InP HEMT low-noise amplifier front-end aboard a 6U-Class satellite
- TEMPEST-D raises the instrument TRL from 5 to 7.
- TEMPEST-D partners are CSU, JPL and spacecraft provider (TBD).
- To demonstrate capability of U-Class Satellites to provide suitable measurements to advance NASA's Earth Science Goals
- To reduce cost, risk and development time of future constellations of NASA's small satellites to perform Earth Science measurements
- To demonstrate drag-adjusting altitude maneuvers required for time separation for U-Class constellation with common deployment
- To provide precision intercalibration with existing conically-scanning Earth-viewing radiometers with similar frequencies to within 1-2 K
- Start in July 2015 for launch by Jan. 2018; 3 months of operations