

Introduction

1

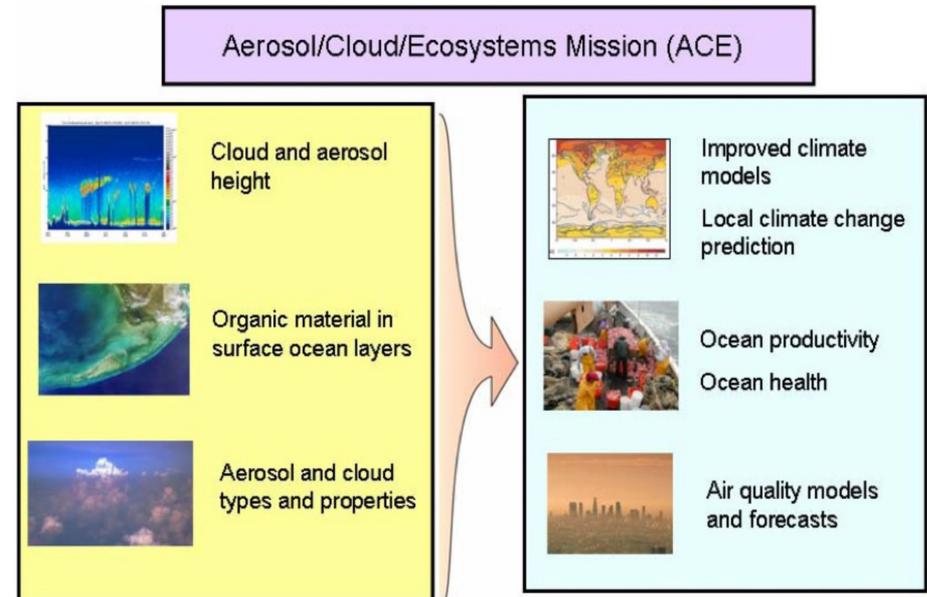
“A “highly accurate multiangle-multiwavelength polarimeter” is a key component of NASA’s future Aerosol-Cloud-Ecosystem (ACE) mission

—NRC Decadal Survey (2007)

JPL’s Multiangle SpectroPolarimetric Imager (MSPI) development effort is maturing key technologies for ACE

Other potential missions include Pre-ACE (PACE) and Earth Venture

- Key design drivers include:
 - ▣ Degree of linear polarization uncertainty ≤ 0.005
 - ▣ Sub-km spatial resolution
 - ▣ Spectral coverage from the UV to SWIR

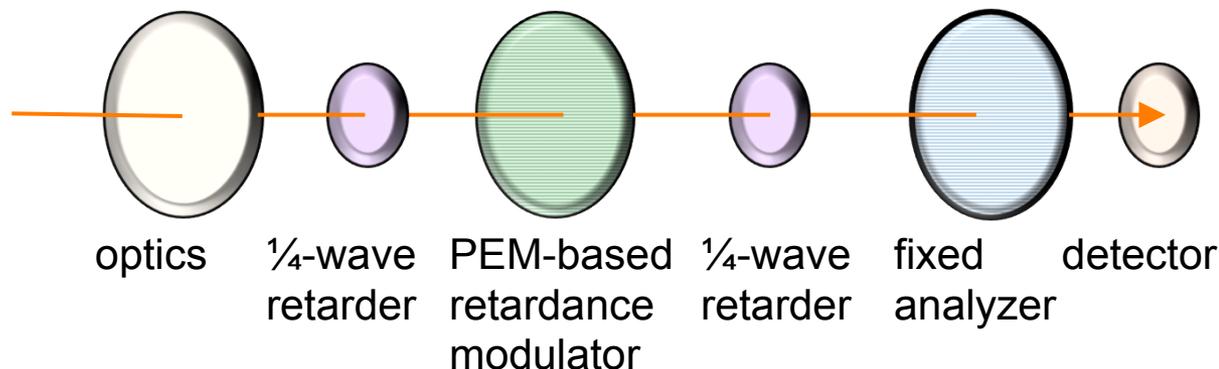
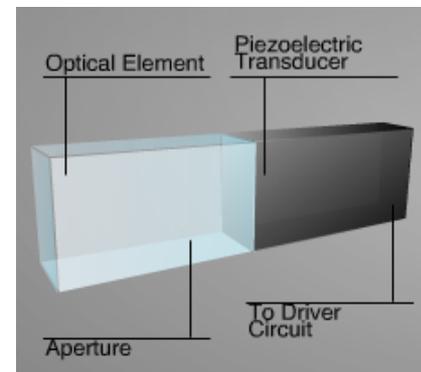


Modulation enables high-accuracy polarimetric imaging

2

- “Polarization modulation is essential to accurate polarimetry in the optical region...one strives to modulate only the polarization preference, leaving the Stokes / sensitivity constant.”
– Tinbergen (2005)

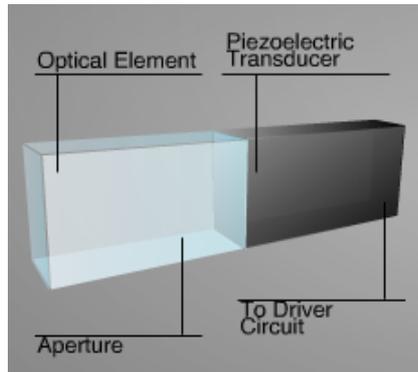
- “The most simple and stable modulators with the best optical properties are the piezoelectric [photoelastic] modulators (PEMs).”
– Povel et al. (1990)



Retardance variation with time rapidly rotates the plane of linear polarization

Current dual-PEM approach

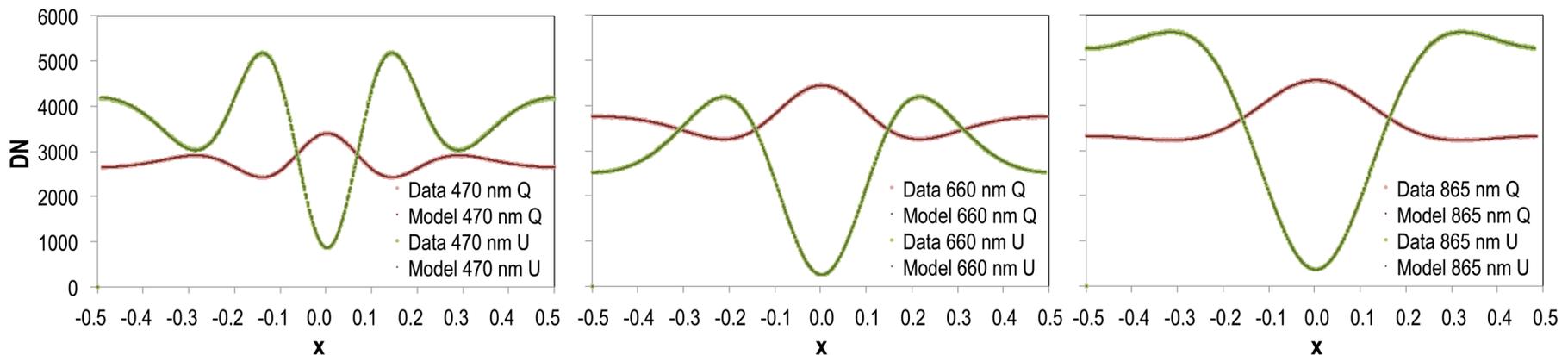
3



Typical PEM modulation frequency is 42 kHz
(24 ms per cycle)

To enable a slower readout, we currently put 2 PEMs in series with slightly different frequencies to generate a 25 Hz beat signal

Modulation patterns from dual PEMs (AirMSPI data)
 x = relative phase within a low frequency (beat) cycle

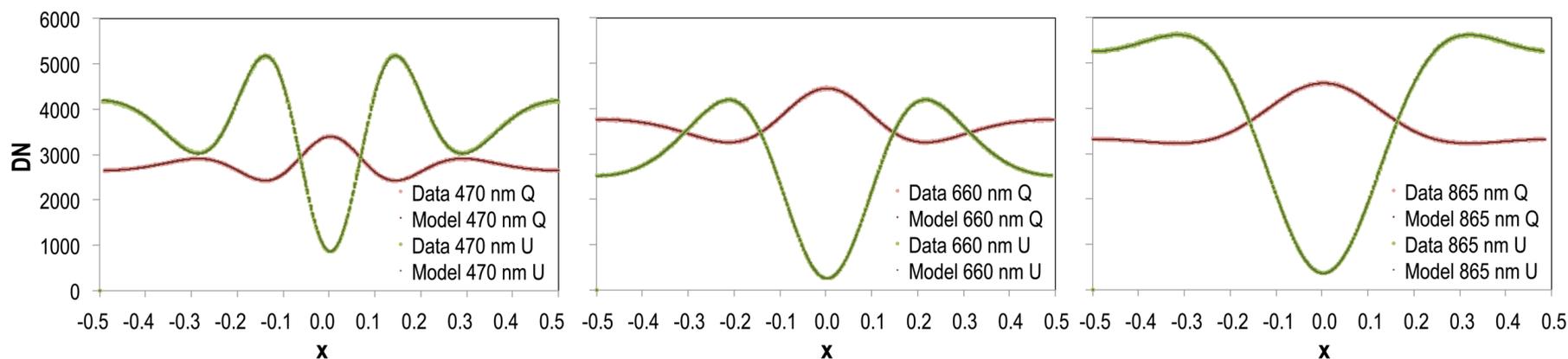


Pushbroom imaging provides I , Q , and U using 2 line arrays for a given spectral band

4



$$S_0(t) = \frac{1}{2} \{I + F(t)Q\} \quad S_{45}(t) = \frac{1}{2} \{I + F(t)U\}$$



Existing Multiangle SpectroPolarimetric Imagers (MSPI)

5



GroundMSPI

Portable field instrument on 2-axis gimbal

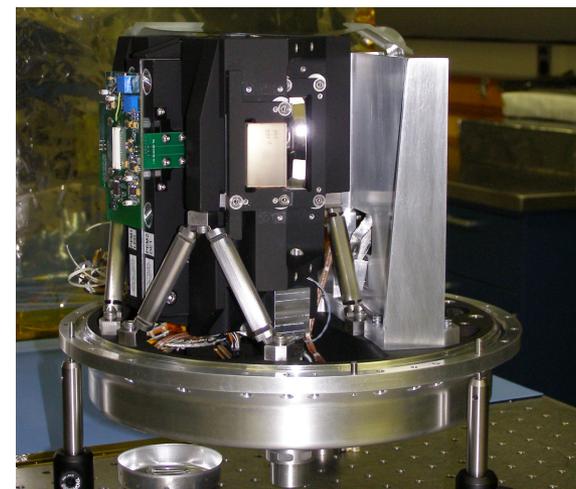
Used for developing surface reflectance models



AirMSPI

Flies in nose of NASA ER-2 with 1-axis gimbal for multiangle viewing $\pm 67^\circ$

Has flown in multiple field campaigns observing aerosols and clouds



AirMSPI-2

Extends spectral coverage into the SWIR and adds O₂ A-band
Currently operating in the lab

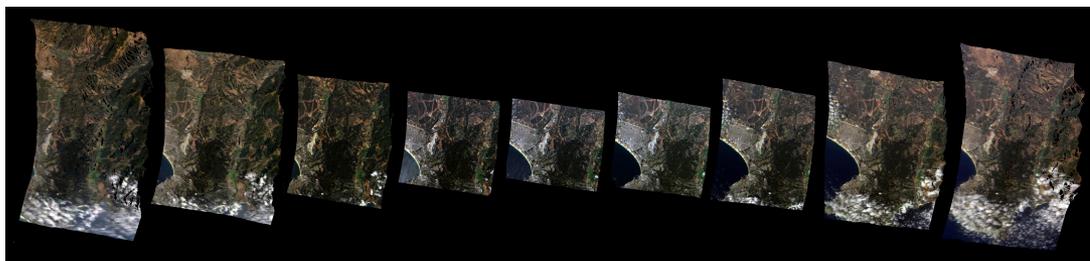
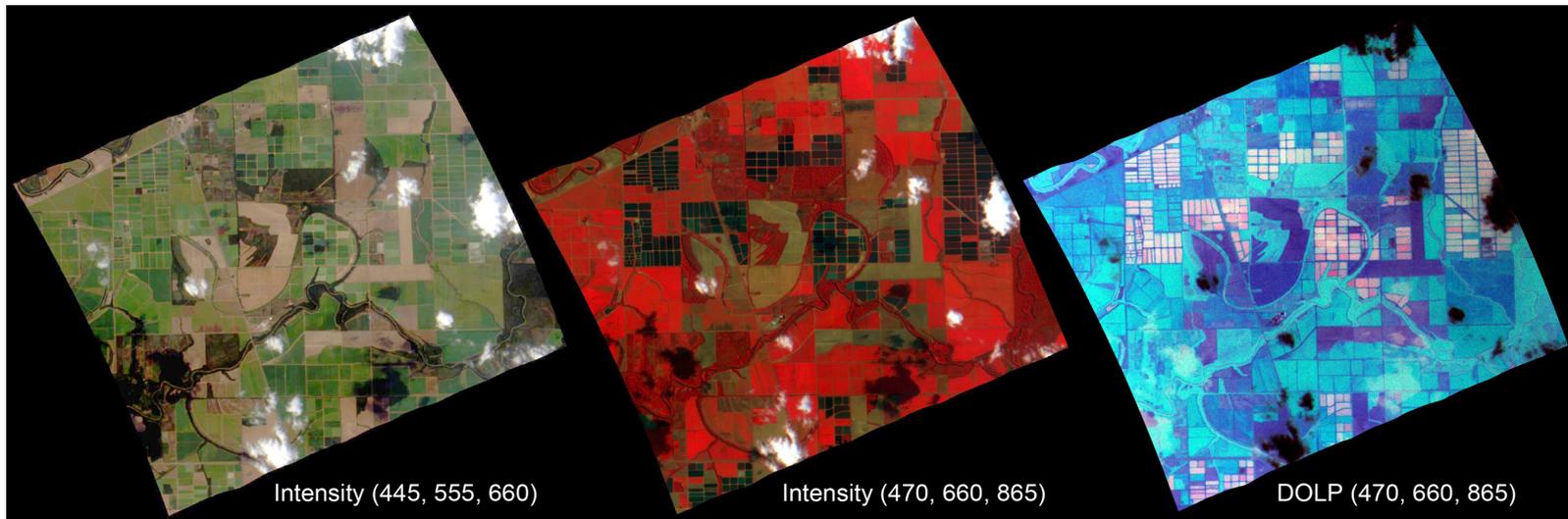
To be installed on the ER-2 in a few months

355, 380, 445, 470*, 555, 660*, 865*, 935 nm

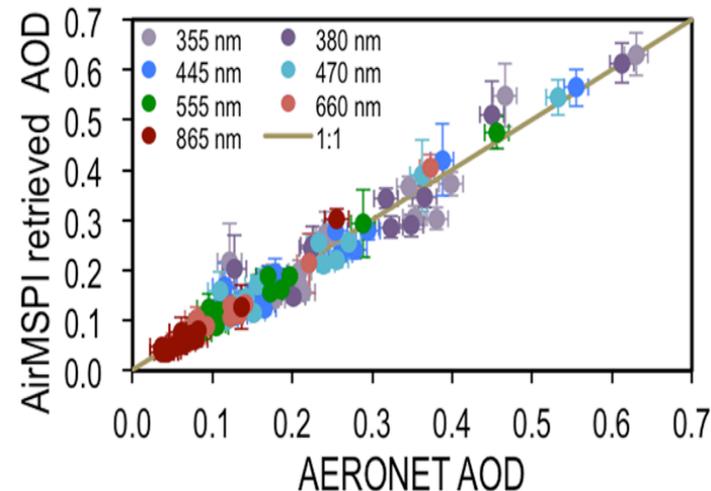
365, 385, 445*, 545, 645*, 751, 763, 865*, 945, 1620*, 1888, 2185* nm

Example AirMSPI aerosol observations over land

6

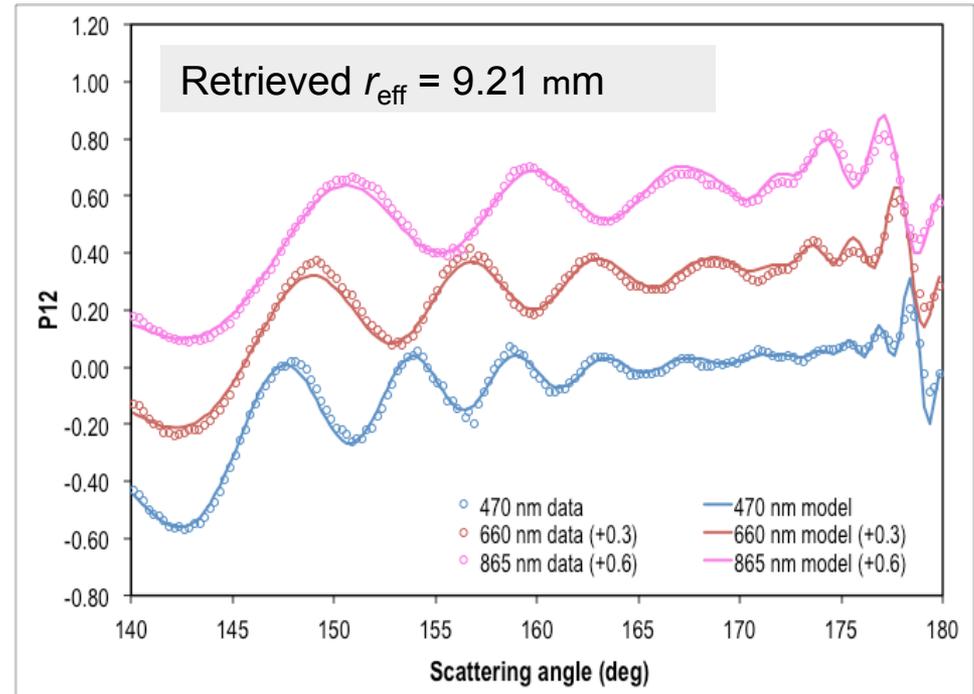
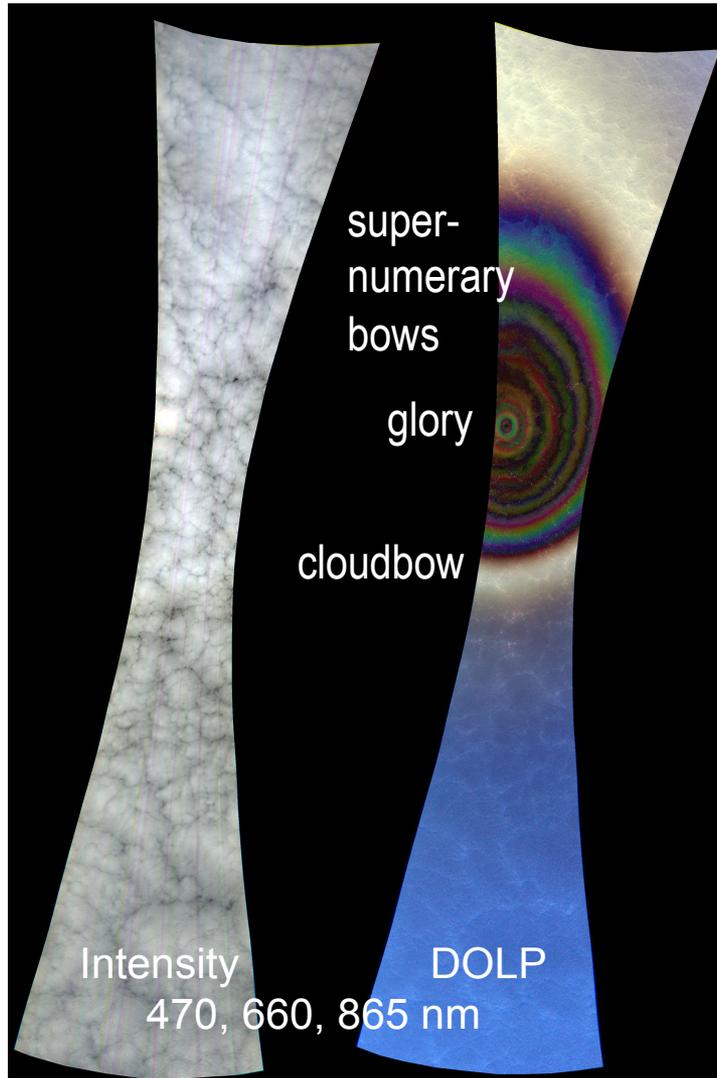


- “Step and stare” mode provides multiangle views of observed targets.
- AirMSPI has flown in the PODEX, SEAC4RS, pre-HyspIRI, and CalWater-2 field campaigns.



Example AirMSPI cloud observations

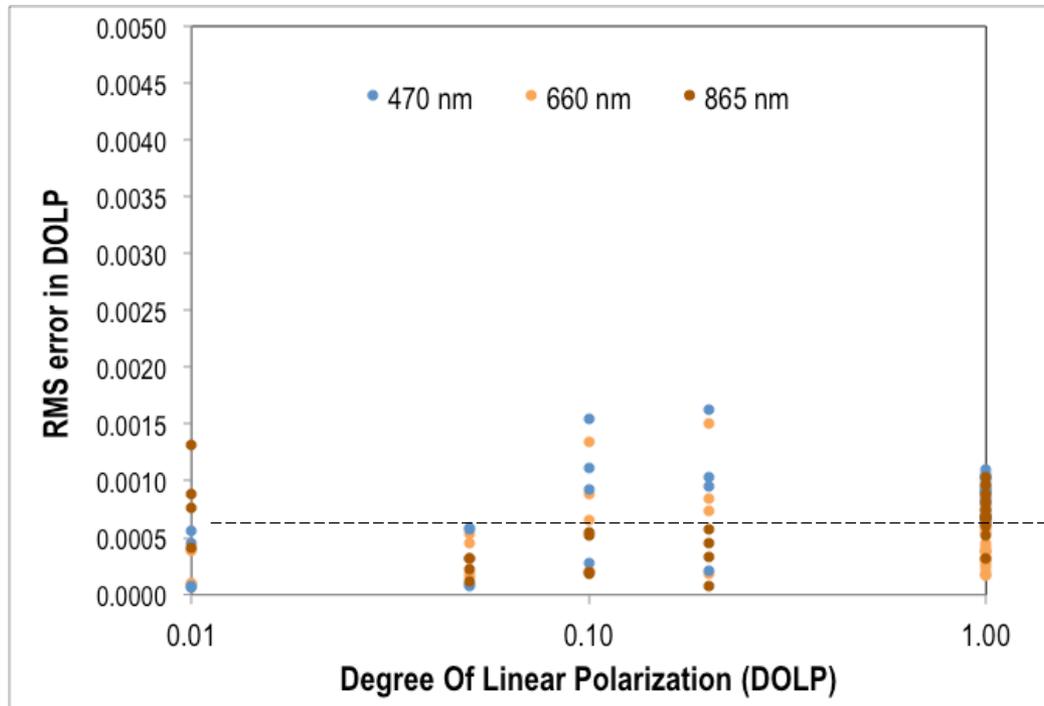
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- “Sweep” mode provides continuous angular coverage.
- Supernumerary and glory interference fringes are highly apparent in polarized light and enable sensitive retrievals of cloud droplet size distributions.

High polarimetric accuracy and sensitivity

8



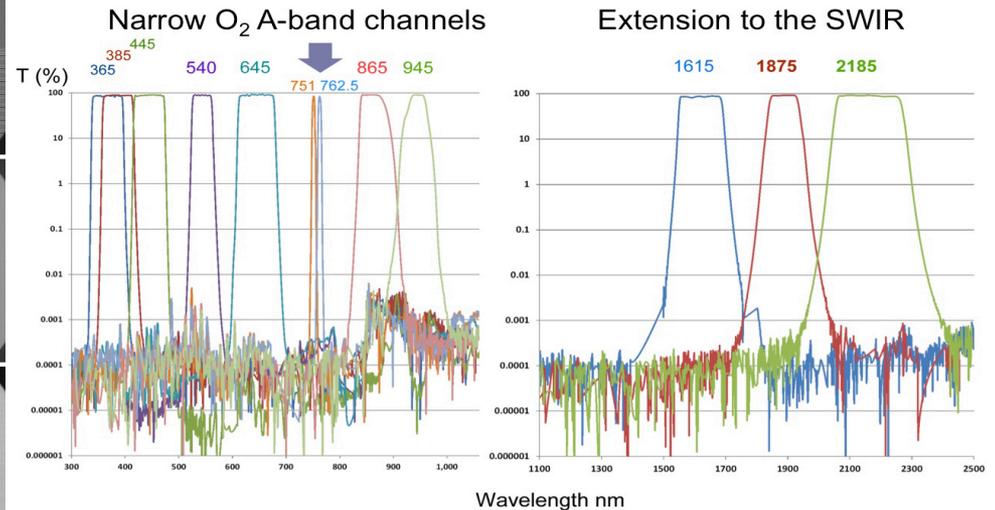
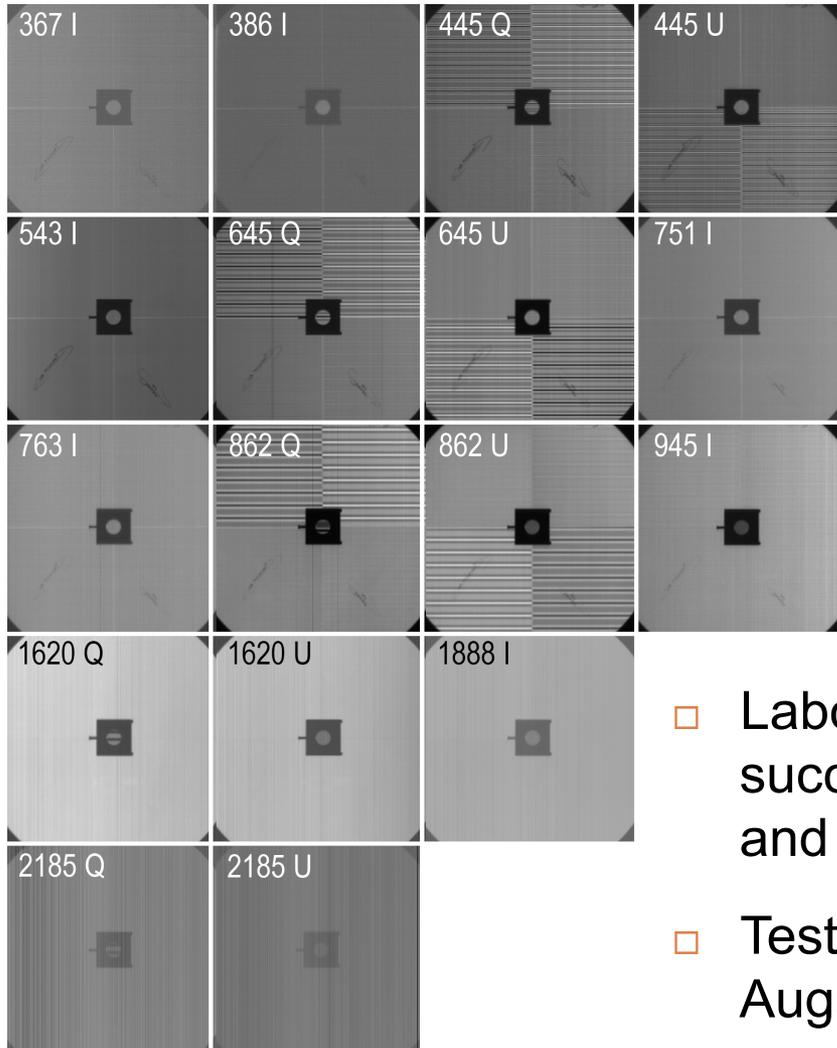
← ACE DOLP
uncertainty
requirement:
 ≤ 0.005

Median DOLP error
from calibrated
AirMSPI data = 0.0006

- Polarimetric calibration eliminates systematic errors.
- For MSPI, the limiting error source is random measurement noise, and is controlled by achieving high signal-to-noise ratio (SNR)

Extension of spectral coverage and range using AirMSPI-2

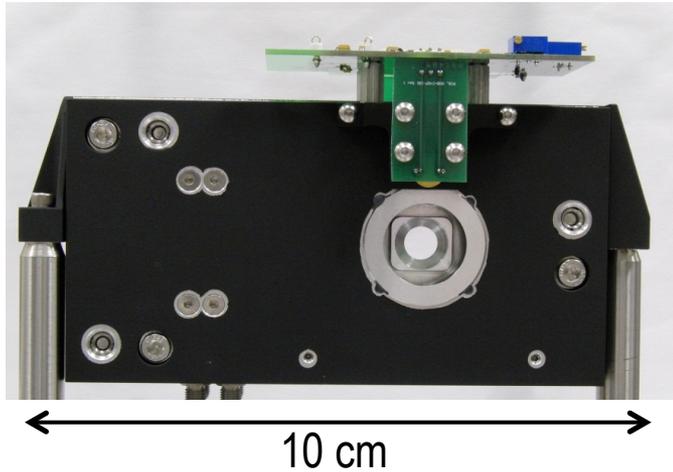
9



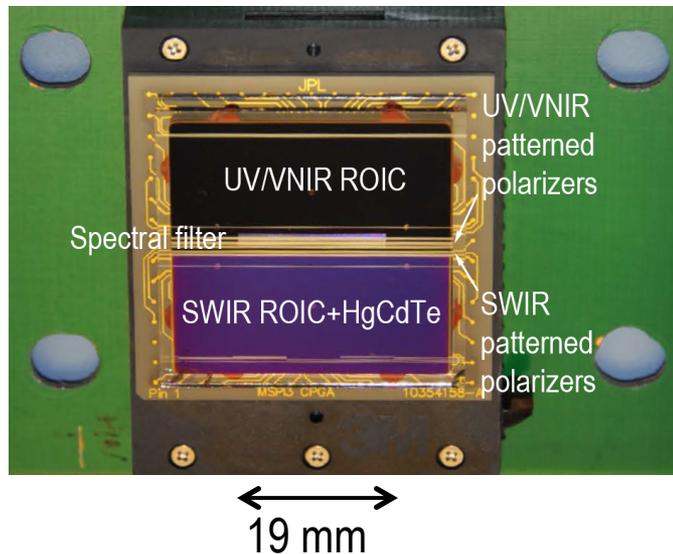
- Laboratory data from AirMSPI-2 demonstrates successful extension of the MSPI modulator and focal plane technology into the SWIR.
- Test flights on the ER-2 are planned for August.

MSPI technologies have been space-qualified to TRL 6

10



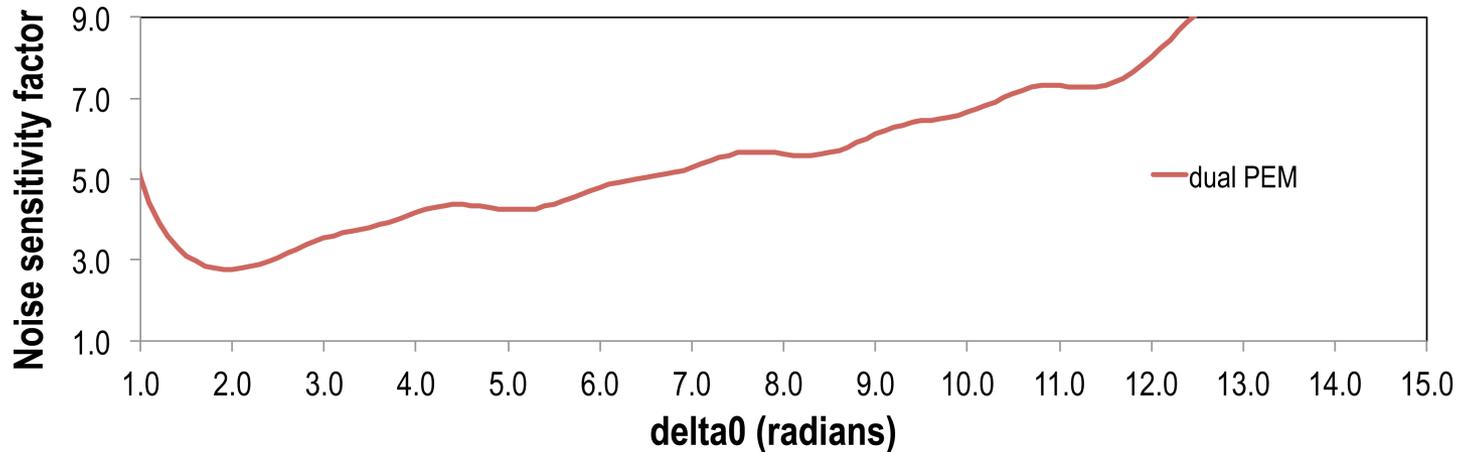
- Polarization modulator
 - Photoelastic modulators
 - Athermal, achromatic quarter waveplates (quartz:MgF₂:sapphire)



- Focal plane
 - High speed/low noise readout integrated circuit (ROIC)
 - Embedded Si CMOS photodiodes for UV/VNIR, hybridized HgCdTe for SWIR
 - Mosaicked spectropolarimetric filters

Limitation of dual-PEM approach

11



The noise sensitivity factor attenuates signal-to-noise ratio (SNR)

- Usable wavelength range is restricted to about 400 – 2100 nm
 - ▣ At high retardances (short wavelengths), the modulation patterns are highly peaked, which sacrifices efficiency and enhances noise sensitivity
 - ▣ At low retardances (long wavelengths), the modulation patterns have reduced amplitude, which also enhances noise sensitivity

Demodulation of the signal from a single PEM enables improved performance

12

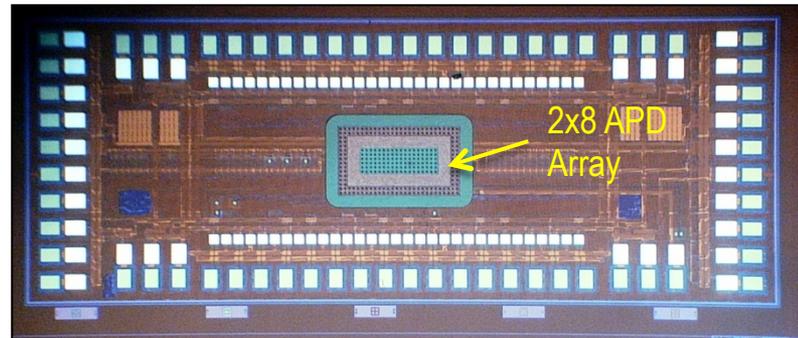
- Requirements
 - ▣ Sampling and demodulation of the 42 kHz waveforms from a single PEM
 - ▣ Low demodulation noise
 - ▣ Sensing from UV to SWIR and potentially MWIR
 - ▣ High speed readout without penalty of capacitor thermal (kTC) noise associated with repeated signal sampling

- How can this be done?
 - ▣ Temporal binning of charges collected during each 24 ms PEM cycle
 - ▣ Use of avalanche photodiode (APD) gain overwhelms kTC noise associated with the charge binning

The concept makes use of avalanche photodiodes (APDs)

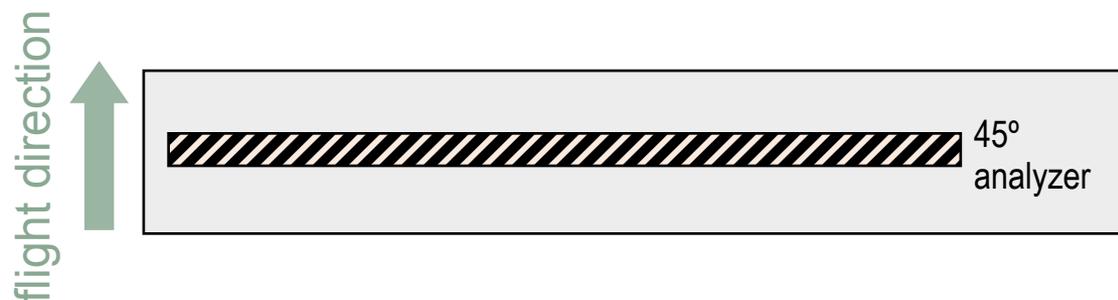
13

- Photoelectrons liberate secondary electrons, which in turn knock more electrons out of the semiconductor lattice
- The total number of electrons generated is the avalanche gain G , which ranges from ~ 10 -1000
 - ▣ kTC noise has a much less deleterious effect than conventional photodetectors because the ratio of switching noise to shot noise decreases by $G^{1/2}$
- A 2x8 HgCdTe APD array with 16 high-speed digital and analog outputs is being procured from DRS Technologies for proof of concept (350 nm – 3.5 mm)



Pushbroom imaging provides I , Q , and U using 1 line array for a given spectral band

14

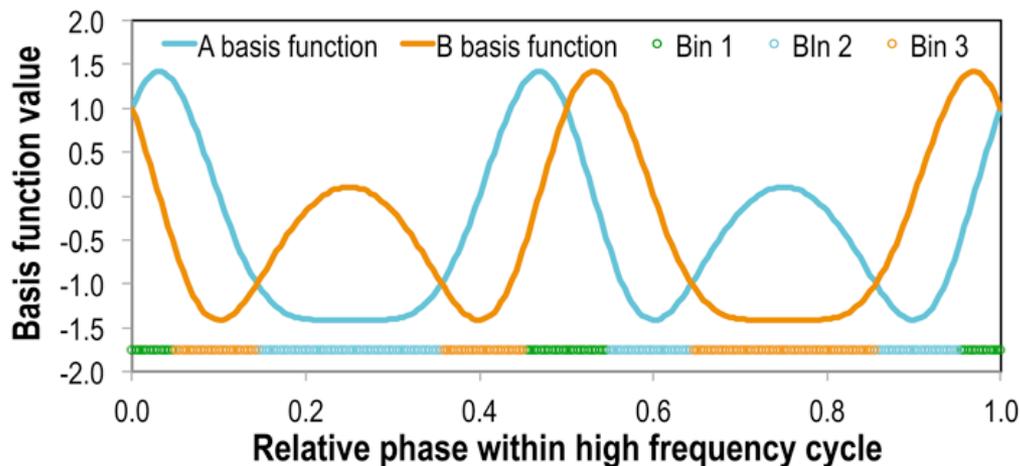


$d(t)$ = PEM time-varying retardance (sinusoidal)

$$S_{45} = \frac{1}{2} [I + Q \sin \delta + U \cos \delta] = \frac{1}{2} [I + A(\cos \delta + \sin \delta) + B(\cos \delta - \sin \delta)]$$

$$A = (Q + U) / 2$$

$$B = (U - Q) / 2$$



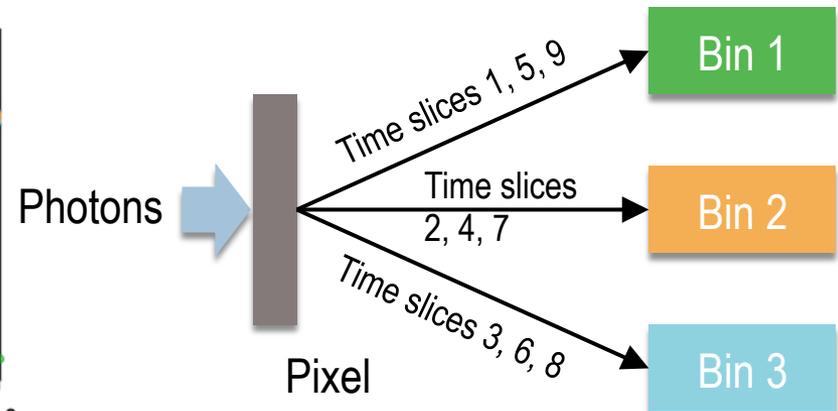
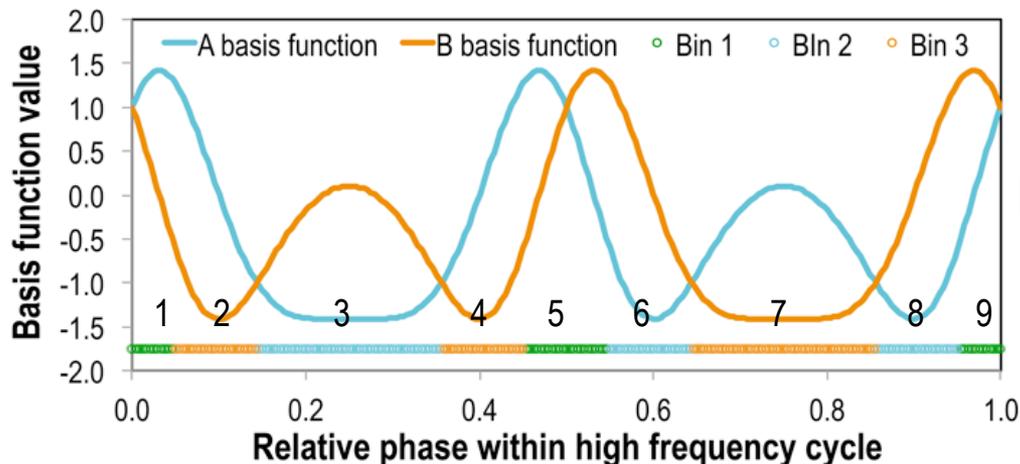
Modulation of A and B are mirror images of each other during a PEM high frequency cycle

Equalizes noise sensitivity of Q and U

Binning concept

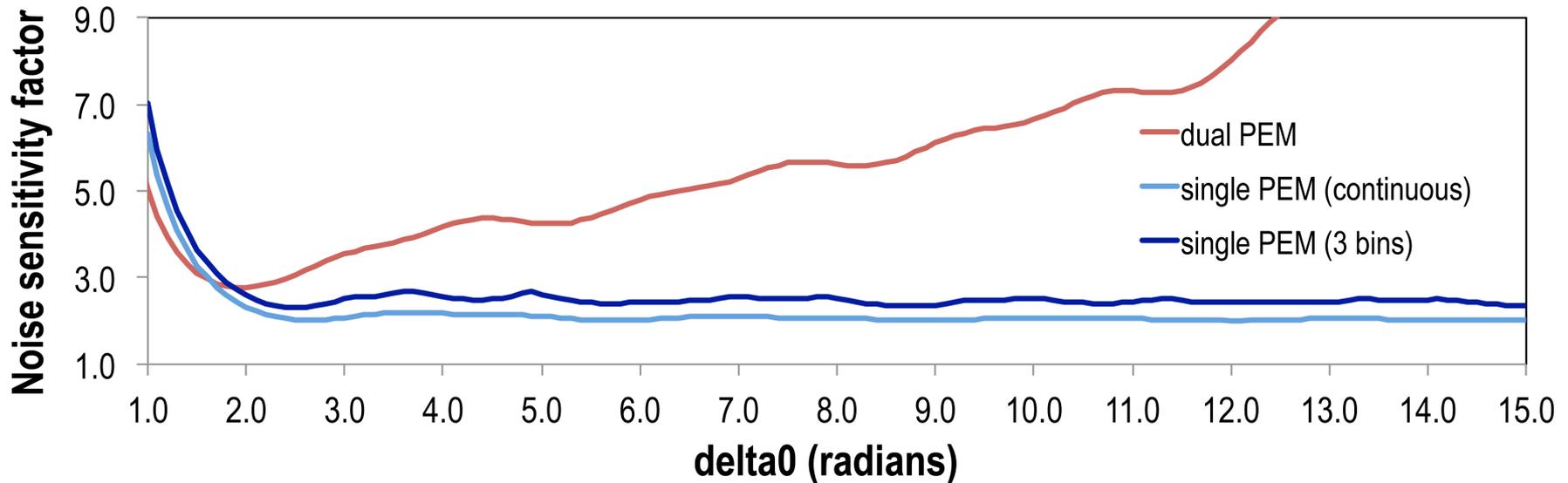
15

- Since we need to solve for just 3 unknowns (I , A , and B ; or equivalently, I , Q , and U) the required information content is achievable by distributing the continuously arriving signal into 3 discrete bins
- Optimizing the time sampling of the 3 bins provides lowest sensitivity to random noise



Comparison of dual PEM and single PEM noise sensitivities

16

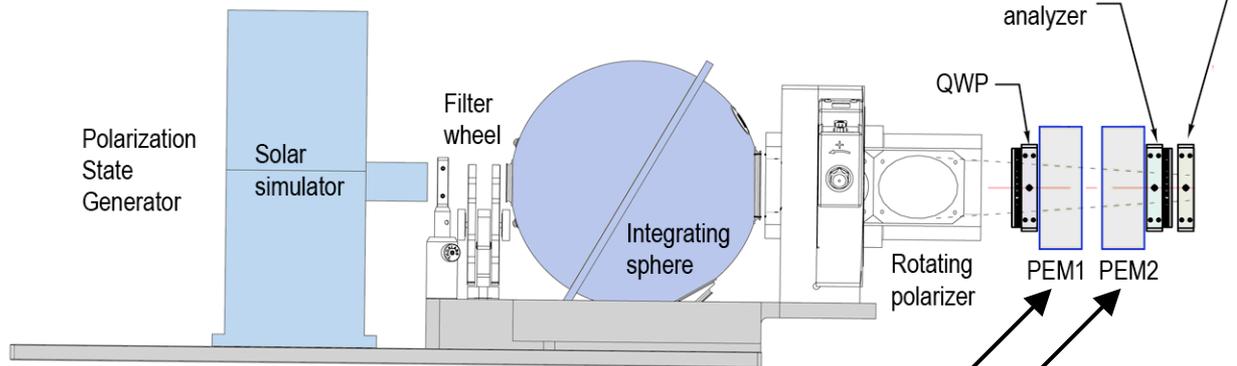


- Single PEM approach makes more efficient use of available light and is less sensitive to random noise over a much wider range of PEM retardance (i.e., wavelengths)

Planned proof of concept

17

Existing polarization state generator (PSG) from IIP-07, IIP-10



New optics

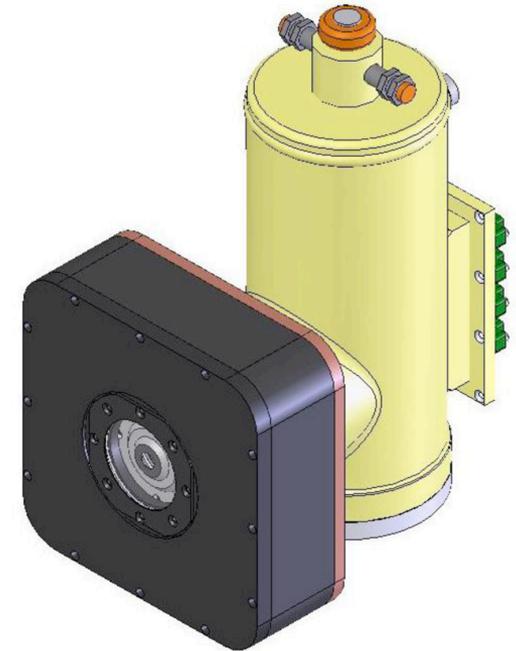
Collimating lens

Polarization analyzer

QWP

Rotating polarizer

PEM1 PEM2



Detector and dewar from DRS

Only a single PEM is required to retrieve I, Q, and U

After initial tests, a second PEM with fast axis at 45° to the first will be added to measure circular polarization V

Conclusions

- The MSPI dual PEM polarimetric imaging approach has been successfully demonstrated in airborne flight
- The next generation concept eliminates one PEM and uses avalanche photodiodes to further advance the measurement capability
 - Less sensitivity to random noise
 - Good noise performance over a wider range of wavelengths (UV-MWIR feasible)
 - I, Q, and U are all recovered from a single pixel
 - Inclusion of a second PEM with fast axis at 45° to first enables measurement of circular polarization, V
- A proof-of-concept system using a 2x8 APD array is being built