



# A Methane Lidar for Greenhouse Gas Measurements

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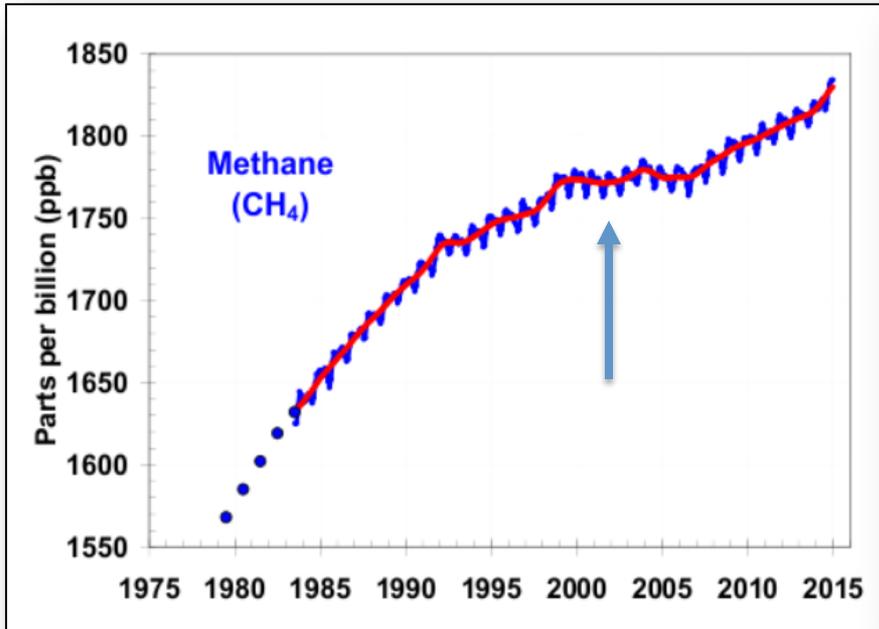
# Outline



- Motivation - Why measure Methane?
- GSFC Measurement Approach
- Technology and Challenges
- Current Status
- Airborne Campaign
- Summary

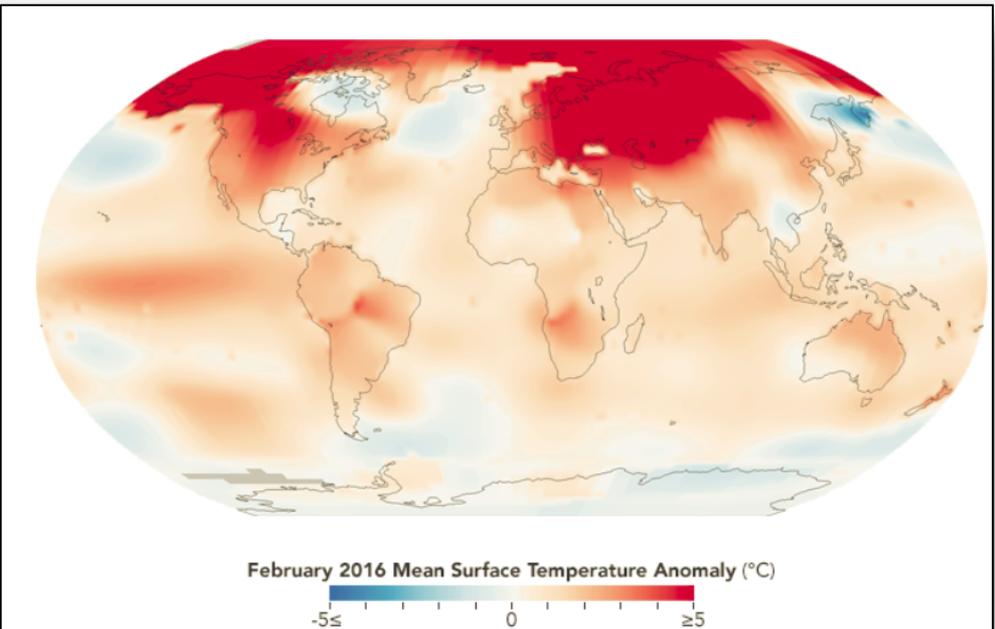


# Why measure Methane?



Source: [www.esrl.noaa.gov/gmd/aqgi/aqgi.html](http://www.esrl.noaa.gov/gmd/aqgi/aqgi.html)

Methane Trend since 1975



Source: [http://www.giss.nasa.gov/research/features/201603\\_gistemp/](http://www.giss.nasa.gov/research/features/201603_gistemp/)

February 2016 was the warmest February in 136 years of modern temperature records. That month deviated more from normal than any month on record.

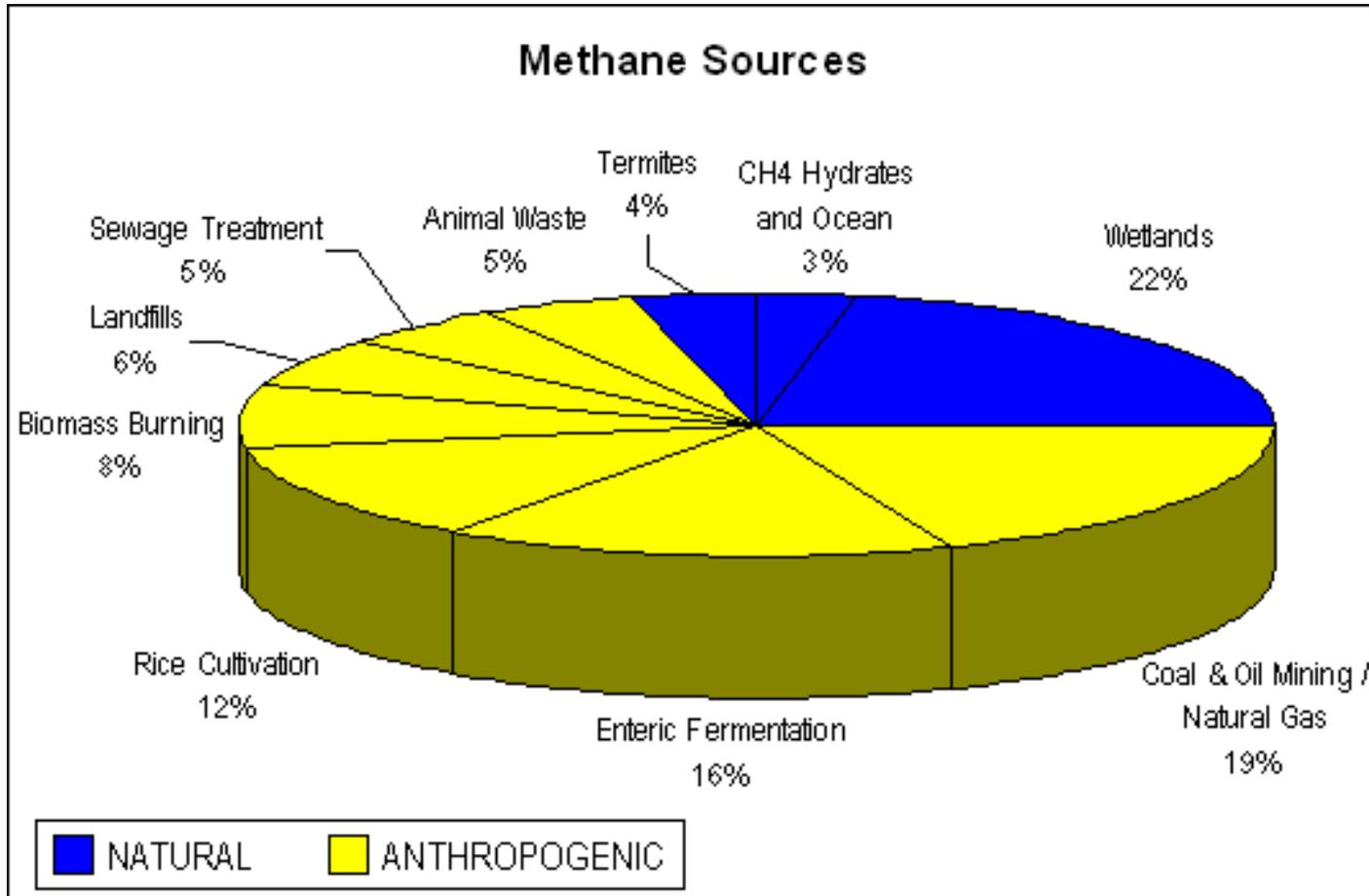


Gas	Estimated 1750 tropospheric concentration <sup>1</sup>	Recent tropospheric concentration <sup>2</sup>	GWP <sup>3</sup> (100-yr time horizon)	Atmospheric lifetime <sup>4</sup> (years)	Increased radiative forcing <sup>5</sup> (W/m <sup>2</sup> )
<b>Concentrations in parts per million (ppm)</b>					
Carbon dioxide (CO <sub>2</sub> )	278 <sup>6</sup>	397.2 <sup>7</sup>	1	~ 100-300 <sup>4</sup>	1.91
<b>Concentrations in parts per billion (ppb)</b>					
Methane (CH <sub>4</sub> )	722 <sup>8</sup>	1823 <sup>2</sup>	28	12 <sup>4</sup>	0.50
Nitrous oxide (N <sub>2</sub> O)	270 <sup>9</sup>	327 <sup>2</sup>	265	121 <sup>4</sup>	0.19
Tropospheric ozone (O <sub>3</sub> )	237 <sup>1</sup>	337 <sup>2</sup>	n.a. <sup>3</sup>	hours-days	0.40

Source: DoE <http://cdiac.esd.ornl.gov/> and IPCC Chapter 8

CH<sub>4</sub> is removed from the atmosphere by a single process, oxidation by the hydroxyl radical (OH), but the effect of an increase in atmospheric concentration of CH<sub>4</sub> is to reduce the OH concentration, which, in turn, reduces destruction of additional methane, effectively lengthening its atmospheric lifetime.

# Sources of Methane



Source: <http://icp.giss.nasa.gov/education/methane/intro/cycle.html>

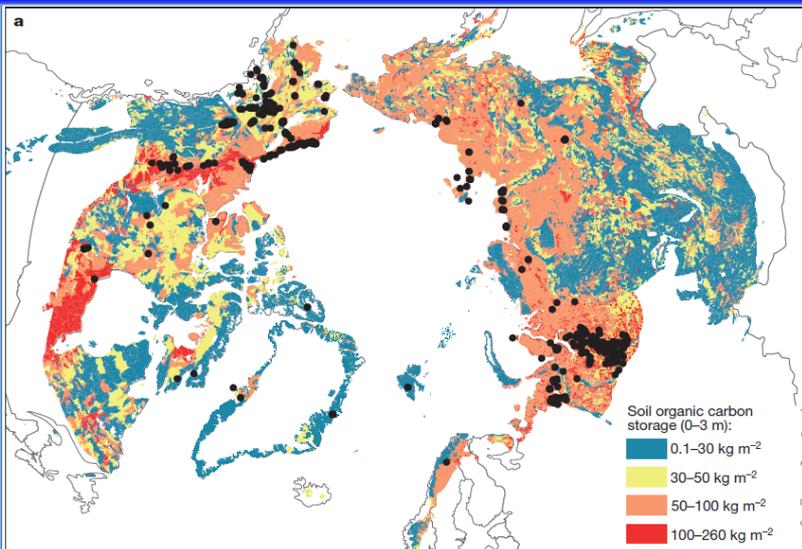


Katey Walter Anthony Univ. of Alaska Fairbanks <https://www.youtube.com/watch?v=YegdEOSQotE>

FEATURE 20 May 2015

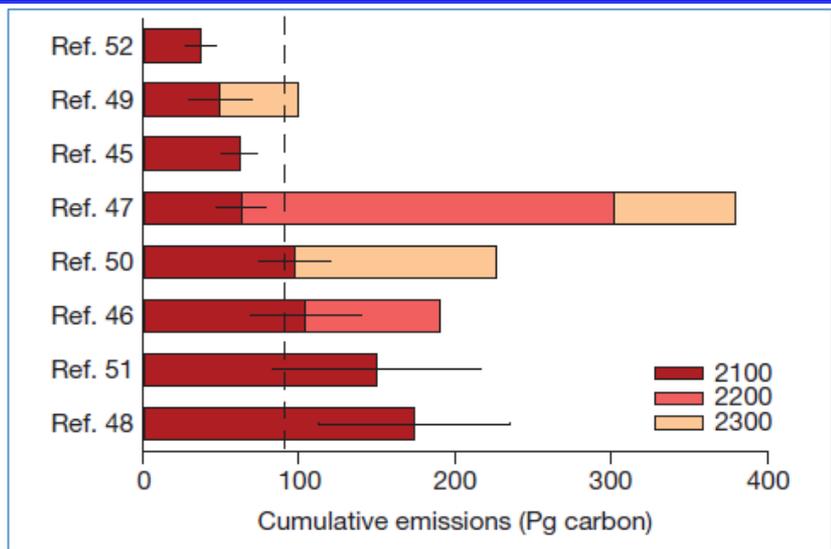
**New  
Scientist**

## Methane apocalypse? Defusing the Arctic's time bomb



Soil organic carbon maps. a, Soil organic carbon pool (kg cm<sup>2</sup>) contained in the 0–3m depth interval of the northern permafrost zone

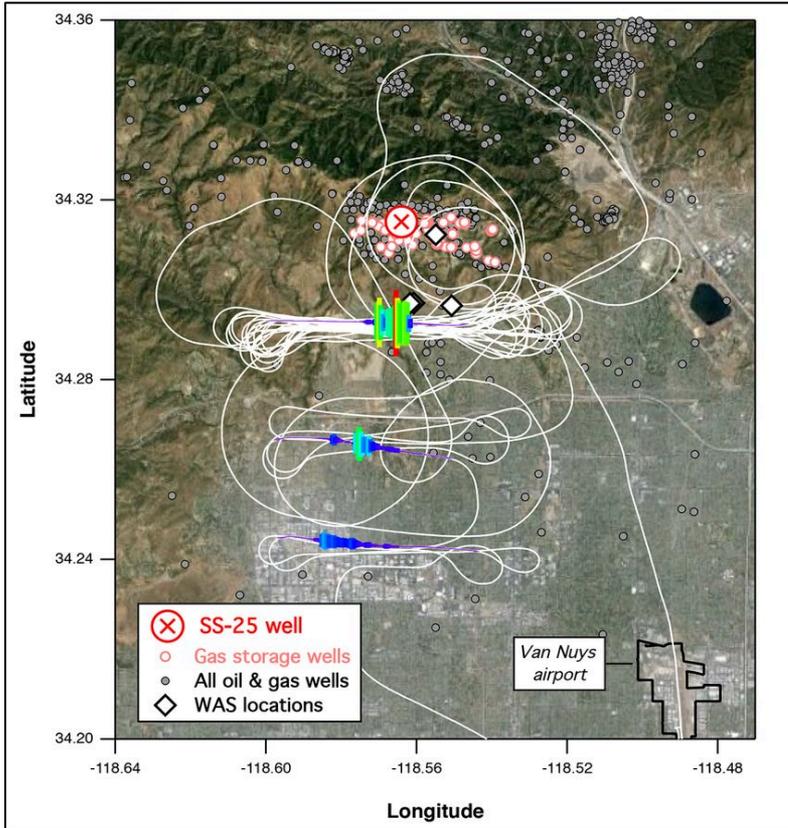
Source: E. A. G. Schuur, et.al., NATURE, VOL 520, 9 APRIL 2015, 174



Model estimates of potential cumulative carbon release from thawing permafrost by 2100, 2200, and 2300.

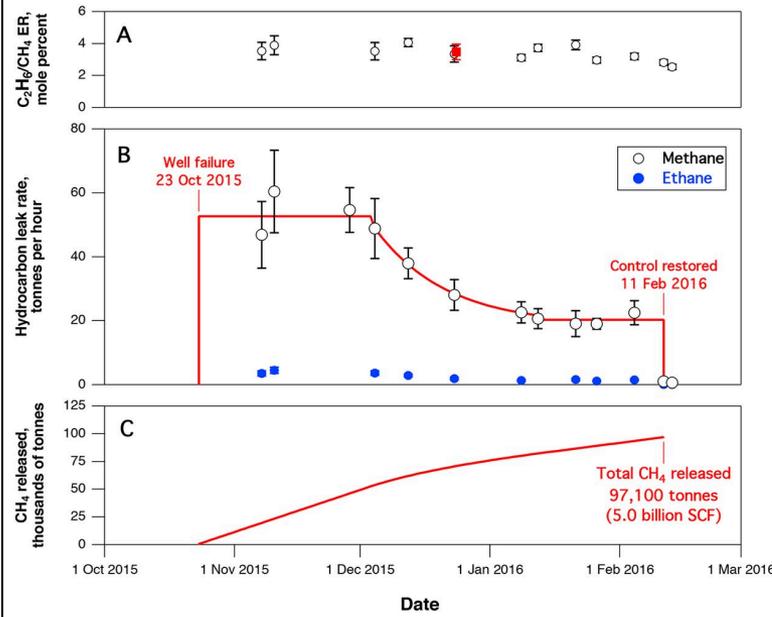
- *“Large quantities of organic carbon are stored in frozen soils (permafrost) within Arctic and sub-Arctic regions. A warming climate can induce environmental changes that accelerate the microbial breakdown of organic carbon and the release of the greenhouse gases carbon dioxide and methane. This feedback can accelerate climate change, but the magnitude and timing of greenhouse gas emission from these regions and their impact on climate change remain uncertain...” E. A. G. Schuur, et.al., NATURE, VOL 520, 9 APRIL 2015, 174*
- *“Here, we report year-round CH<sub>4</sub> emissions from Alaskan Arctic tundra eddy flux sites and regional fluxes derived from aircraft data. We find that emissions during the cold season (September to May) account for ≥50% of the annual CH<sub>4</sub> flux, with the highest emissions from noninundated upland tundra.” Donatella Zona et.al., “Cold season emissions dominate the Arctic tundra methane budget”. PNAS, January 5, 2016, vol. 113, no. 1, 40–45*

# Porter Ranch, CA CH<sub>4</sub> Leak



## REPORT

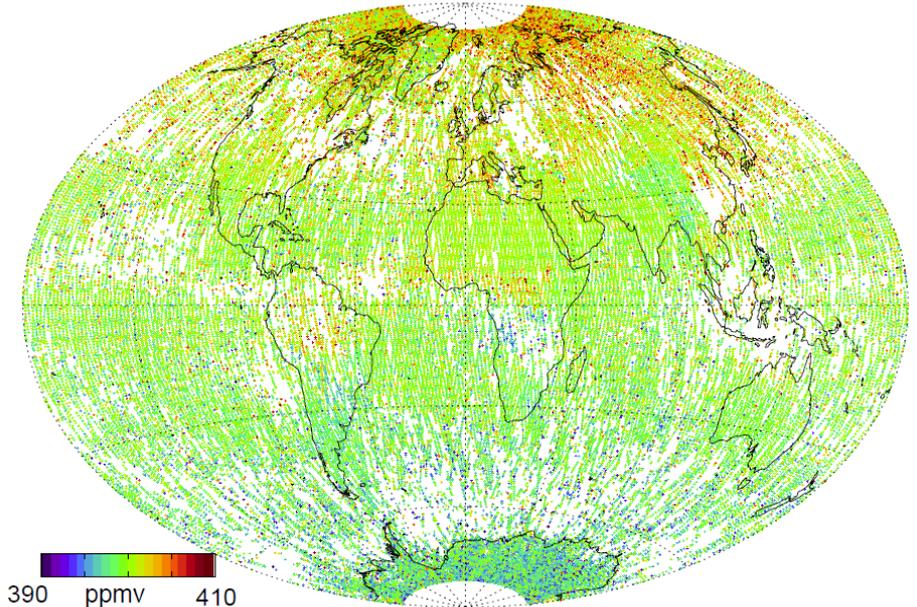
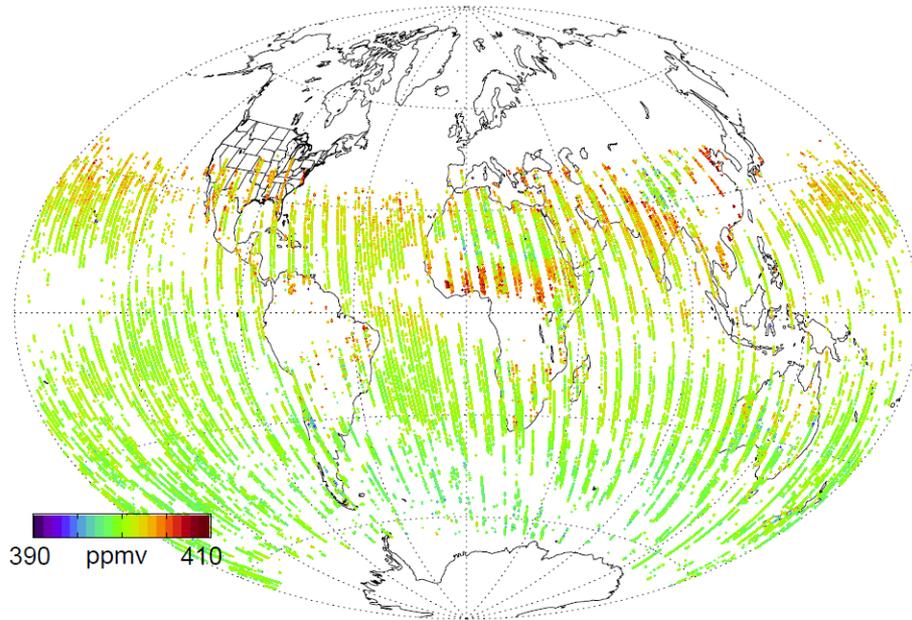
# Methane emissions from the 2015 Aliso Canyon blowout in Los Angeles, CA



*“The 23 October 2015 blowout of a well connected to the Aliso Canyon underground storage facility in California resulted in a massive release of natural gas. Analysis of methane (CH<sub>4</sub>) and ethane (C<sub>2</sub>H<sub>6</sub>) data from dozens of plume transects from 13 research aircraft flights between 7 Nov 2015 and 13 Feb 2016 shows atmospheric leak rates of up to 60 metric tonnes of CH<sub>4</sub> and 4.5 metric tonnes of C<sub>2</sub>H<sub>6</sub> per hour. At its peak this blowout effectively doubled the CH<sub>4</sub> emission rate of the entire Los Angeles Basin, and in total released 97,100 metric tonnes of methane to the atmosphere.” S. Conley et.al. Science 25 Feb 2016, pp. DOI: 10.1126/science.aaf2348*



# Why use a laser?



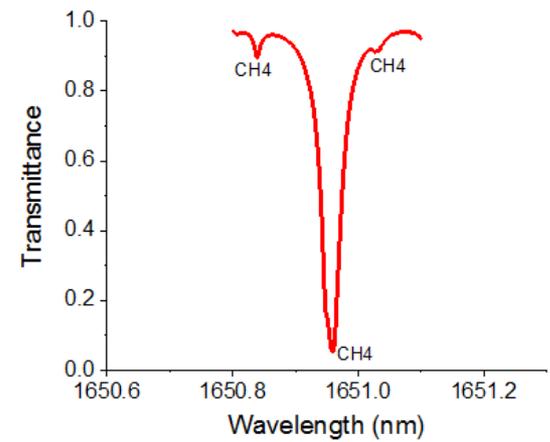
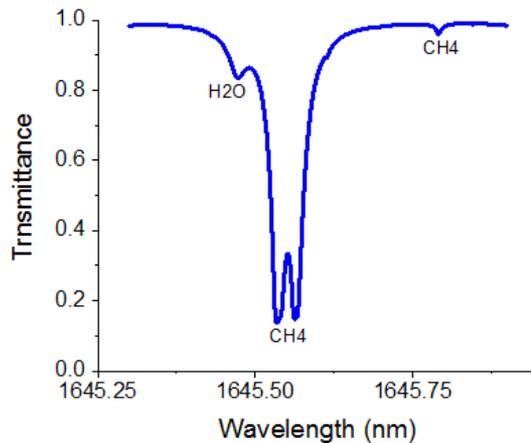
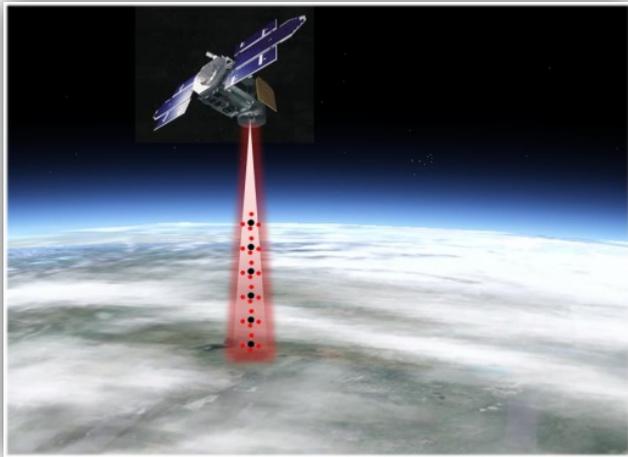
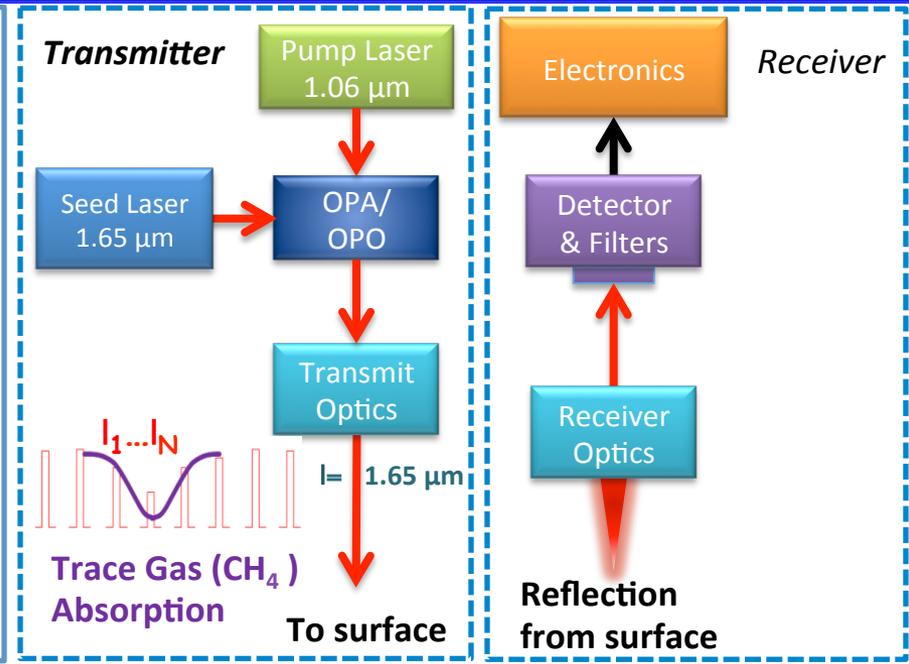
Comparison of actual OCO-2 coverage (left) vs. simulated ASCENDS coverage for December 16-31 2015. The sparse sampling OCO-2 coverage at high latitudes is a major drawback of passive remote sensing missions. (Simulation provided by Dr. Stephan R. Kawa, GSFC).



# GSFC CH<sub>4</sub> IPDA Lidar

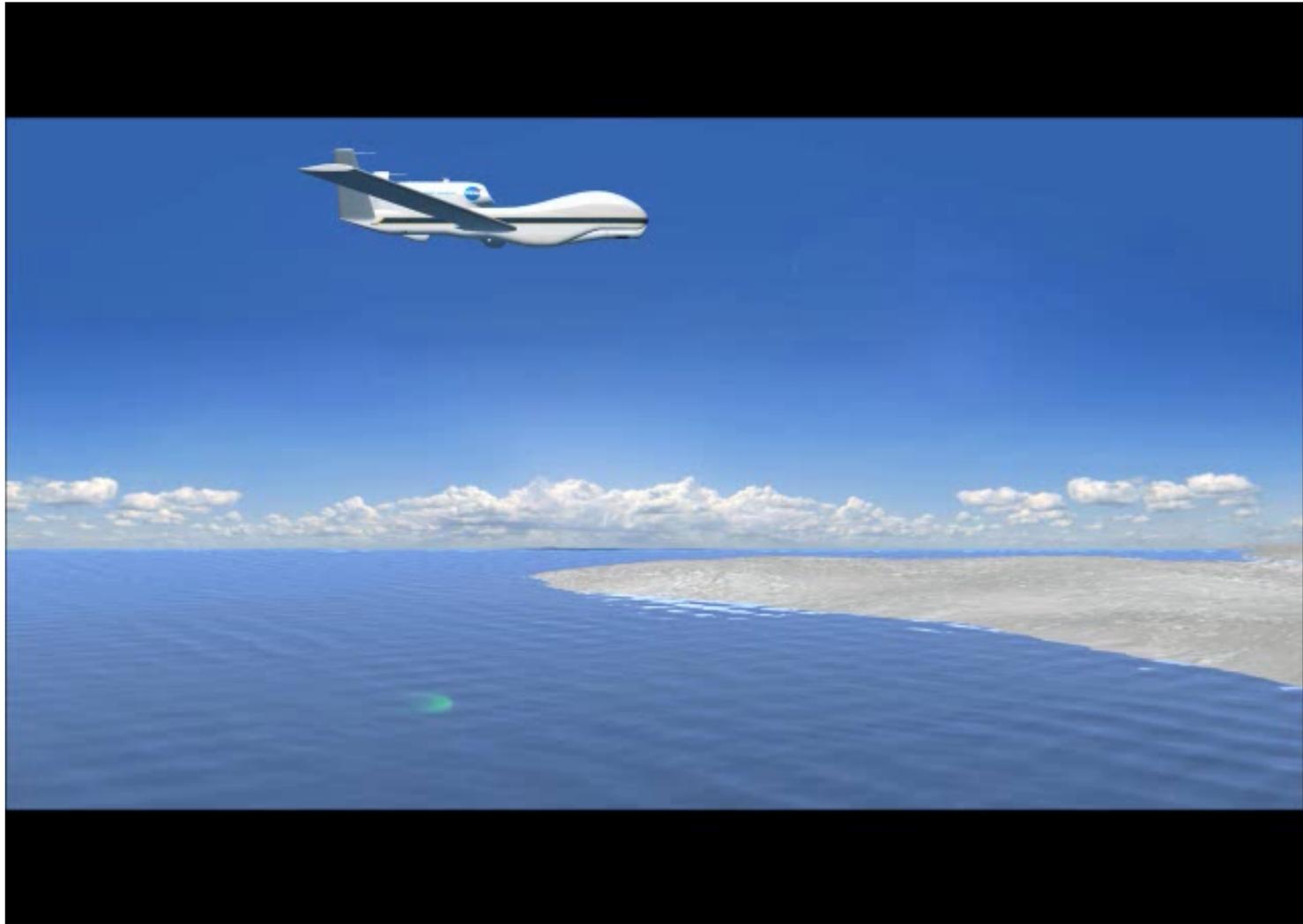


- Transmitter (Laser) technology
  - Current (optimum) Wavelength for CH<sub>4</sub> Earth Detection: ~1.64-1.66 μm
  - Optical Parametric Oscillators (OPO) and Optical Parametric Amplifiers (OPA) are the “baseline” solutions for the transmitter.
  - Other options (Er:YAG and Er:YGG) now possible.
- Receiver (Detector) Technology
  - DRS e-APD



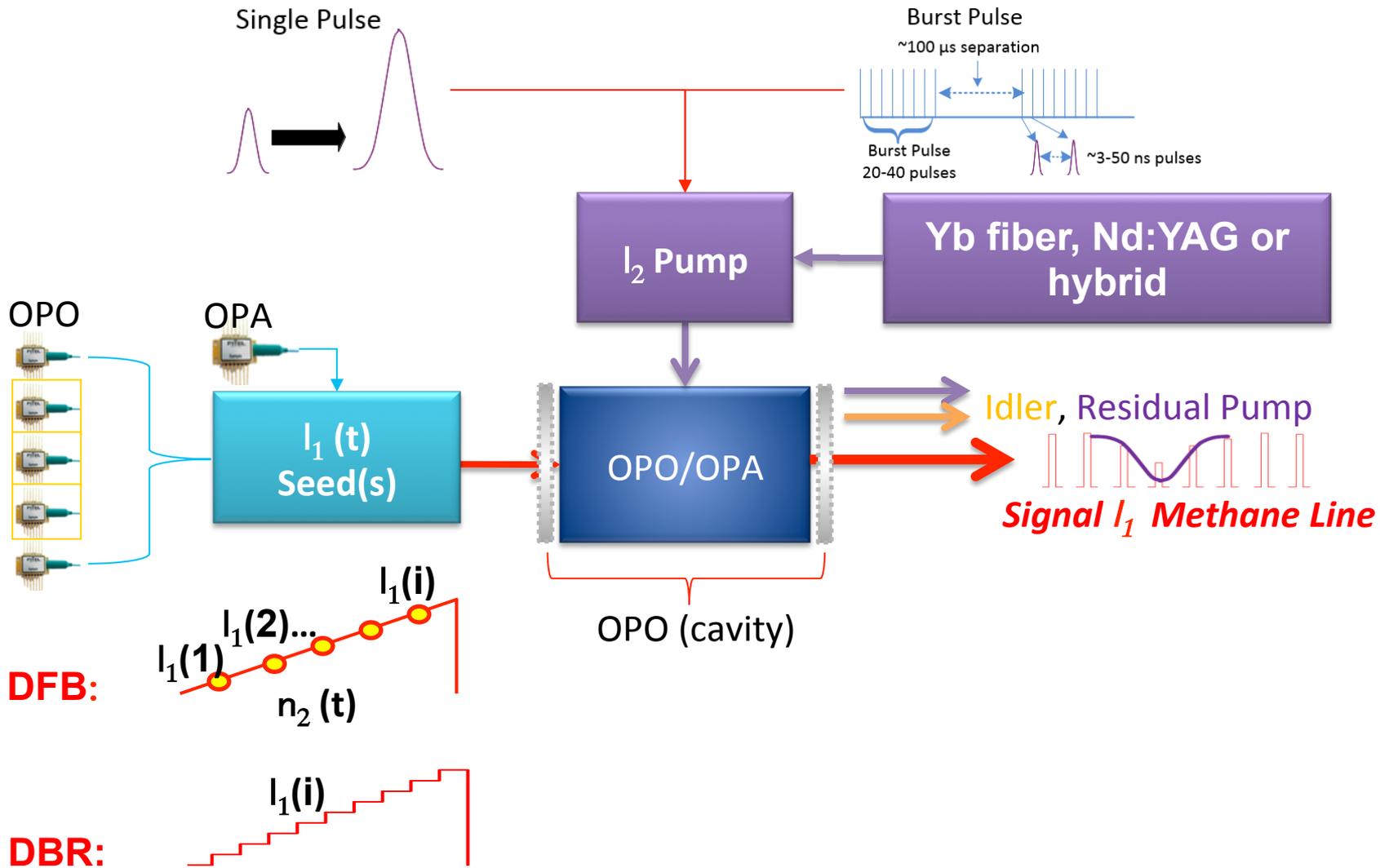


# GSFC CH<sub>4</sub> Lidar with Integrated Path Differential Absorption Lidar (IPDA)





# CH<sub>4</sub> Laser Transmitter: OPO-OPA





# Current GSFC Power scaling options OPA/OPO

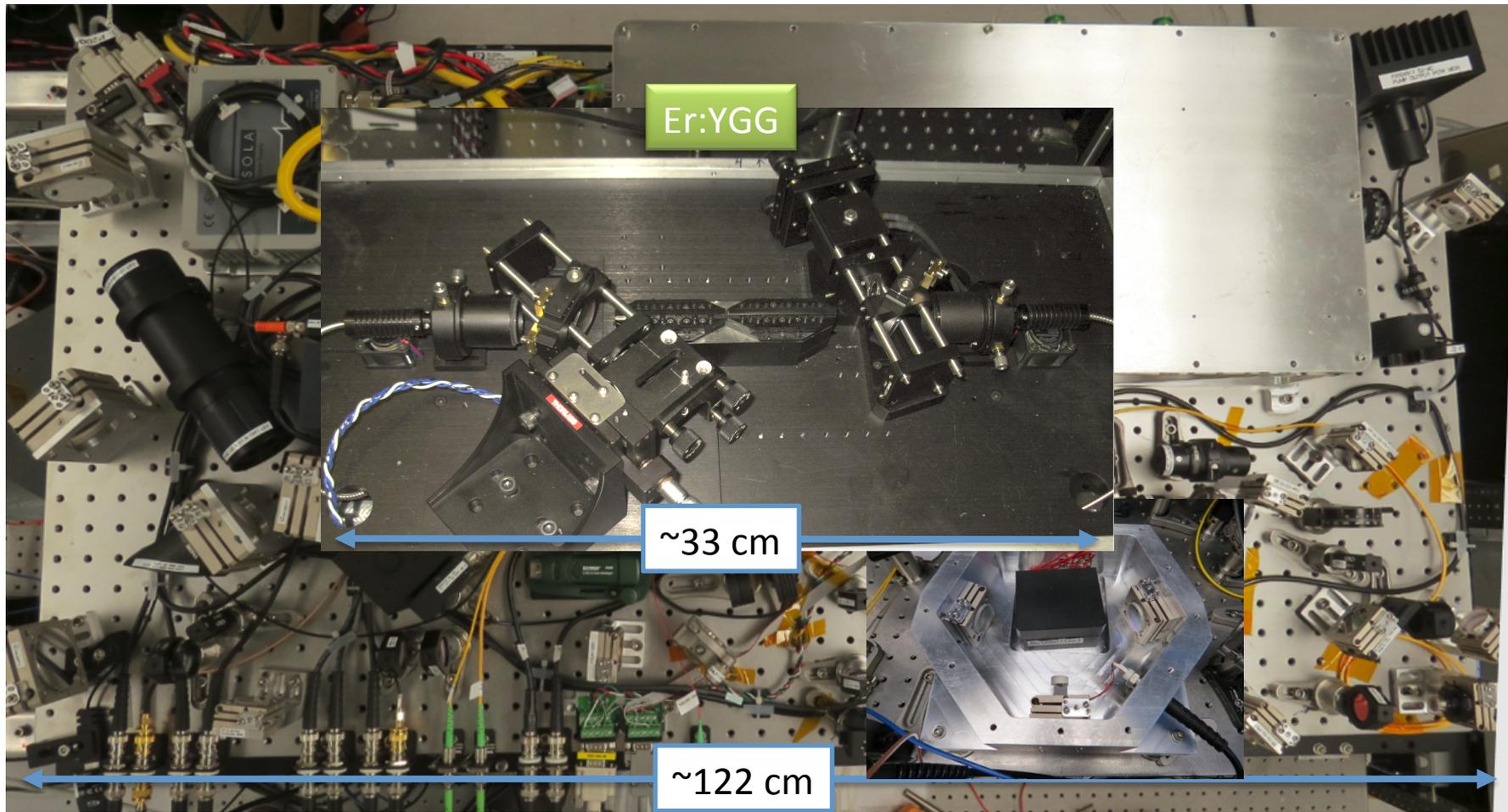


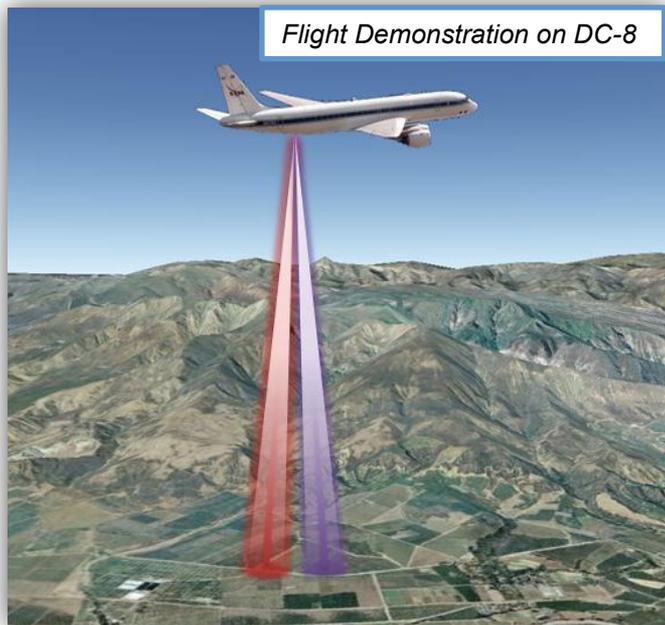
Approach	#1. OPA with smaller burst pulses	#2. OPA with large pump pulse	#3. OPO with large pump pulse
Pump laser	<ol style="list-style-type: none"> <li>1. Burst mode laser. Need to achieve higher energy and pulse uniformity. Hybrid shown to work.</li> <li>2. Burst mode fiber MOPA with Waveguide Amplifier shows promise</li> </ol>	<ol style="list-style-type: none"> <li>1. High power Yb fiber laser (1030 nm).</li> <li>2. Planar Wave Amplifier with commercial laser as Master Oscillator.</li> <li>3. Custom Nd:YAG laser</li> </ol>	<ol style="list-style-type: none"> <li>1. Custom Nd:YAG laser (1064 nm)</li> <li>2. High power Yb fiber laser (1030 nm).</li> </ol>
Seed laser	Existing DFB lasers are OK but would prefer a DBR laser and higher power	High seed power <u>needed</u> Would prefer a DBR	Existing DFB laser is OK would prefer a DBR laser and higher power
Parametric stage	Single OPA stage possible but currently at low energy.	Need multiple OPA stages to achieve high power	Need for cavity locking & step tuning

Er:YAG and Er:YGG

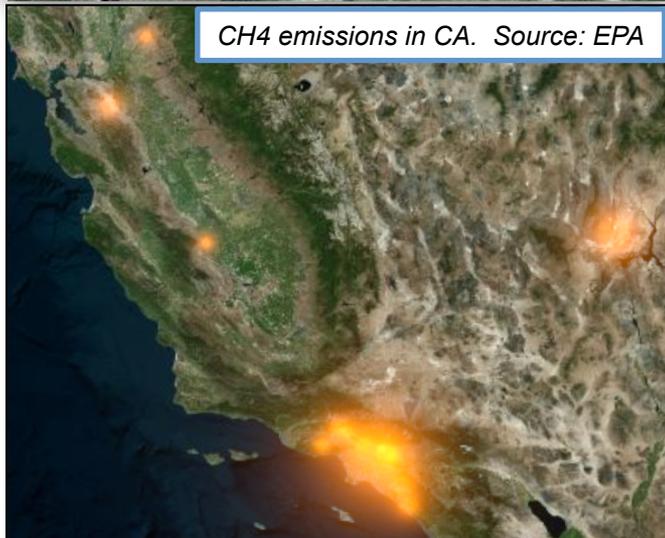


- Why consider other transmitter options?
  - OPAs and OPOs are parametric conversion techniques. They are complex and difficult to implement are sensitive to vibration.
  - Size/mass of airborne instrument needs to be reduced.
  - Erbium Yttrium-aluminum-garnet (YAG) 1.645  $\mu\text{m}$
  - Erbium Yttrium-gallium-garnet (YGG) 1.651  $\mu\text{m}$
- Potential for “simpler” and more efficient transmitter technology.
- Tuning remains an issue.





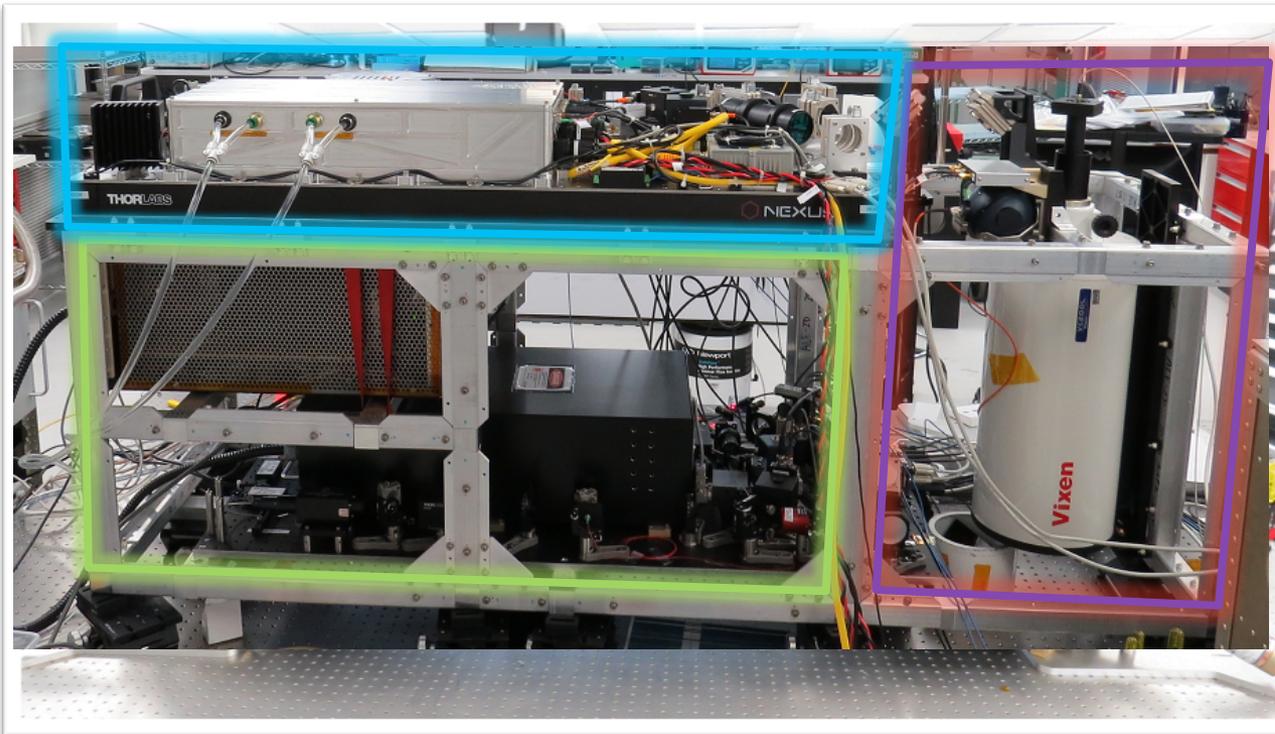
Flight Demonstration on DC-8



CH<sub>4</sub> emissions in CA. Source: EPA

- **Flight Test Methane LIDAR Instruments:**
  - GSFC Methane Sounder (20-l OPA and 5-l OPO)
  - GSFC Picarro
  - COSS-HSC Optec Solutions
  - In-situ CO<sub>2</sub> (LaRC G. Diskin)
- **Conduct several test flights from NASA's Armstrong Science Aircraft Integration Facility (SAIF) in Palmdale, CA:**
  - 1 Engineering flight
  - 2 science flights
  - Approximately 12 hours of flight time in mostly in CA/NV
- **Compare OPO-OPA performance**
- **Assess detector performance**
- **Assess CH<sub>4</sub> LIDAR measurements over Western US**
- **Evaluate derivation of XCH<sub>4</sub> from LIDAR observations and compare with in-situ and calibrations sites whenever possible.**

# CH<sub>4</sub> Airborne Instrument

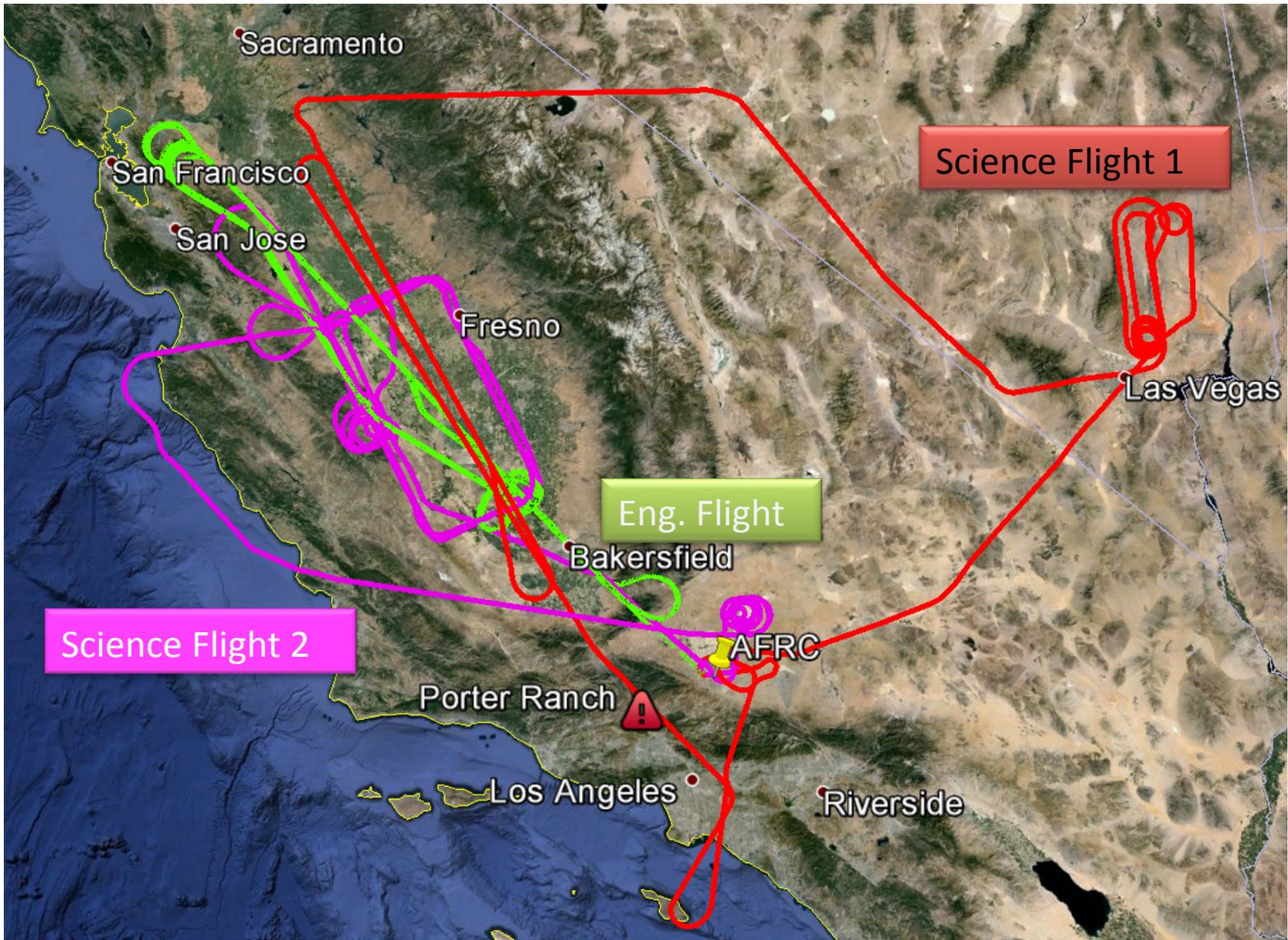


Parameter	Value (OPA/OPO)
Center $\lambda$	1650.9 nm
Number of $\lambda$	20/5
Pulse Width	~700/80 ns
Energy/pulse	~30/250 $\mu$ J
Bin width	4 ns
Divergence	~150 $\mu$ rad
Receiver diam.	20 cm
Field of view	300 $\mu$ rad
Receiver BP	0.8 nm (FWHM)
Averaging time	1/16 s *
Detector Resp.	~1-1.5 x 10 <sup>9</sup> V/W

\*Data analysis uses 1s averages



# Flight Tracks



# Flight Conditions



Clouds



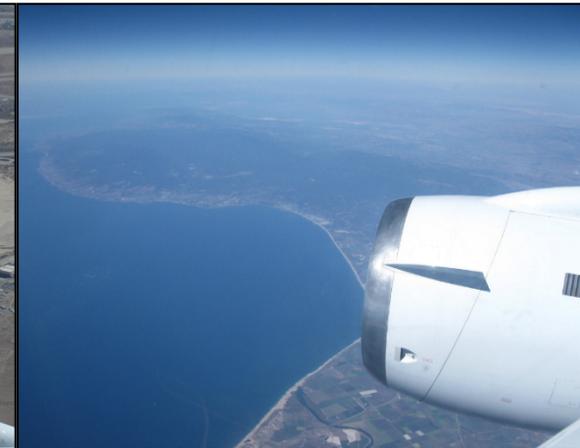
Cultivated/Urban/Landfill



Varying Topography



Desert/Spirals



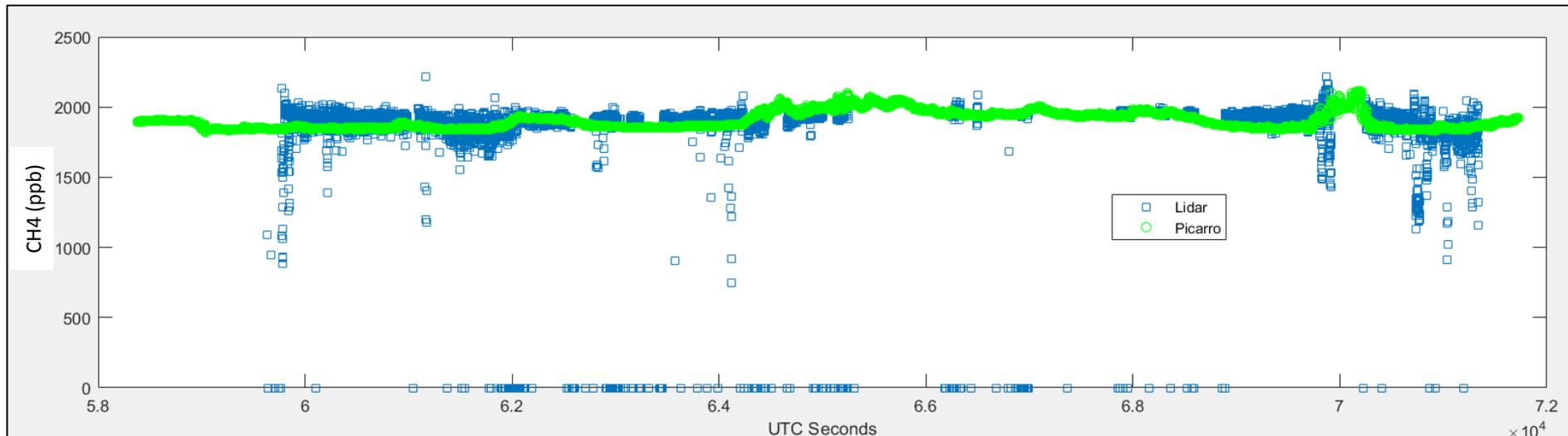
Ocean



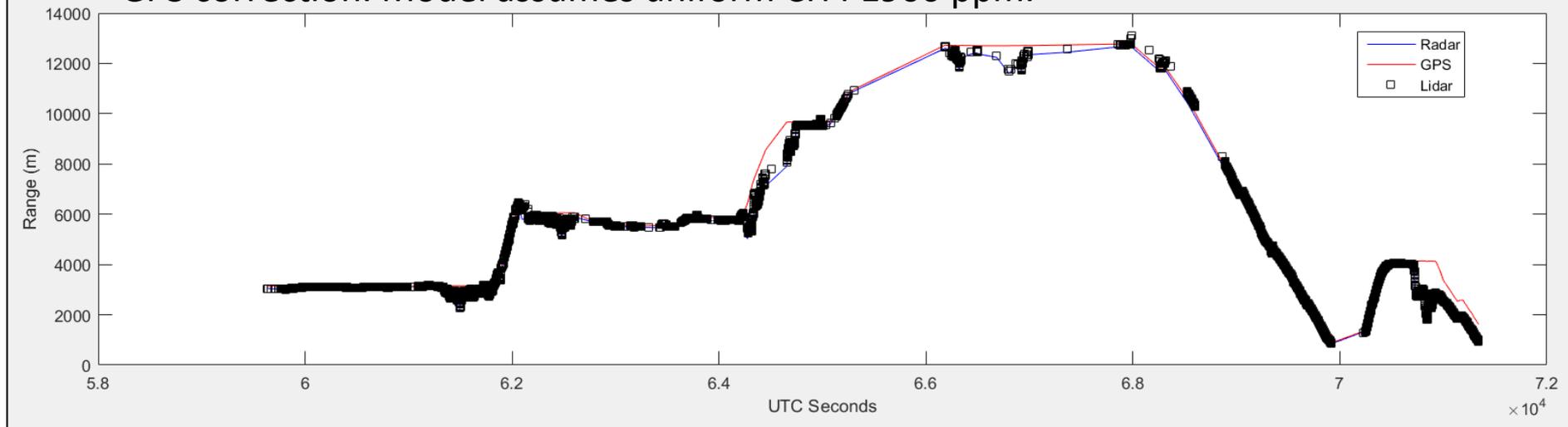
Evening/Night



# Engineering Flight CH<sub>4</sub> mixing Ratio

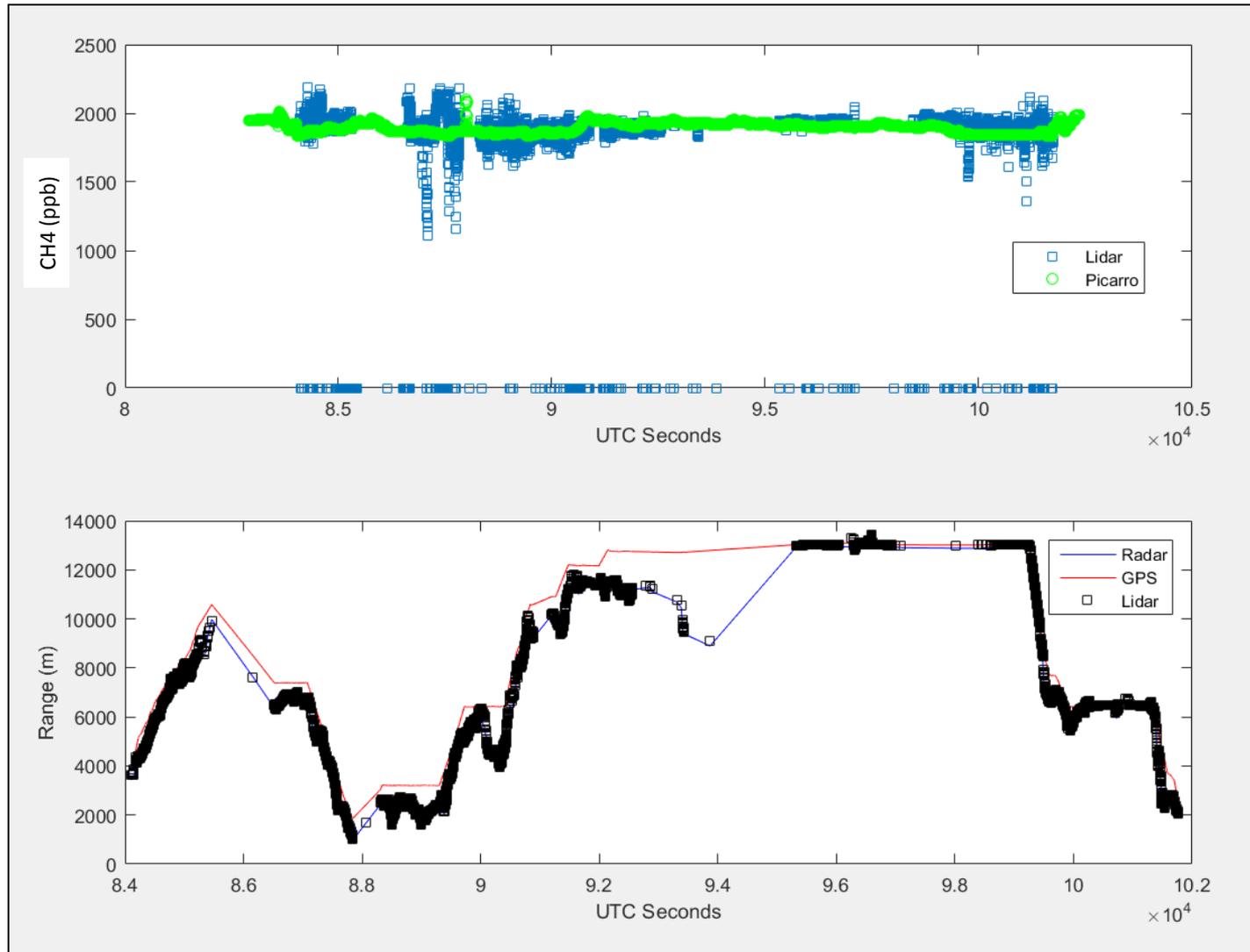


Data is filtered for Ground Returns. 1 s averaging. No DRS non-linearity correction. No GPS correction. Model assumes uniform CH<sub>4</sub> 1900 ppm.

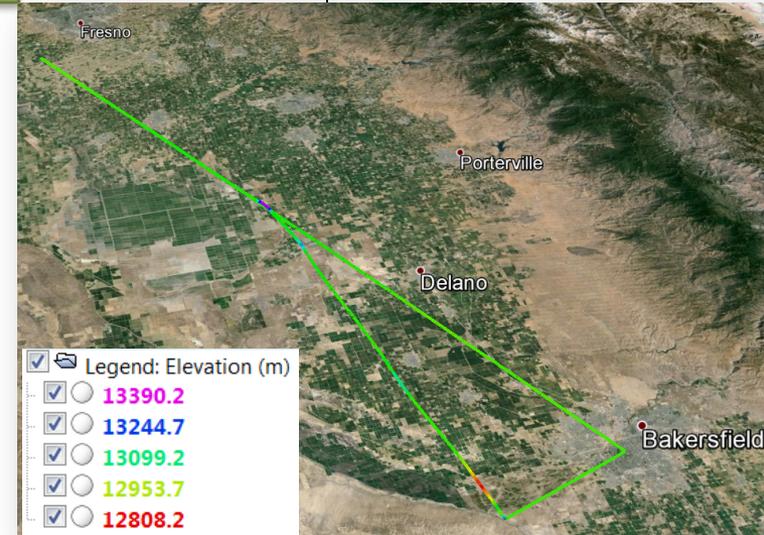
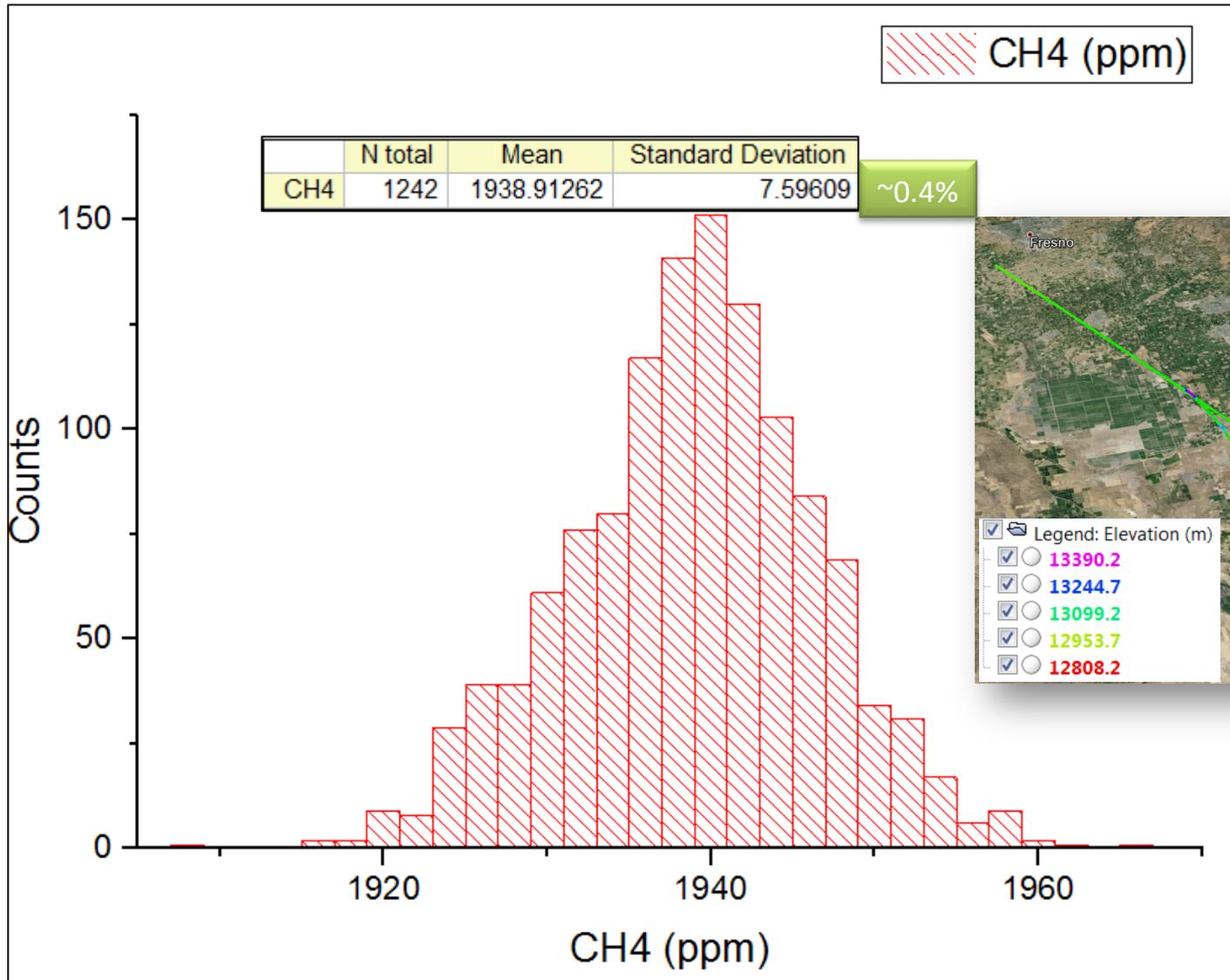




# Science Flight 1 CH<sub>4</sub> mixing Ratio



# OPA Precision 1s averaging

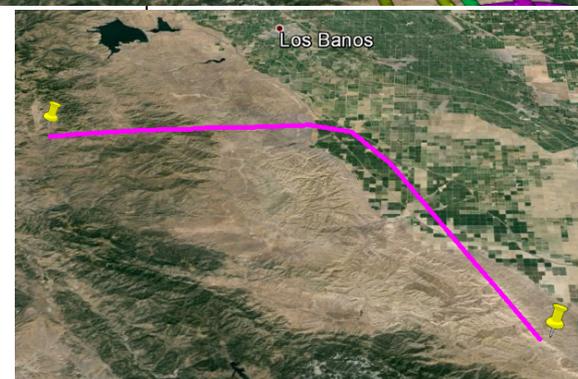
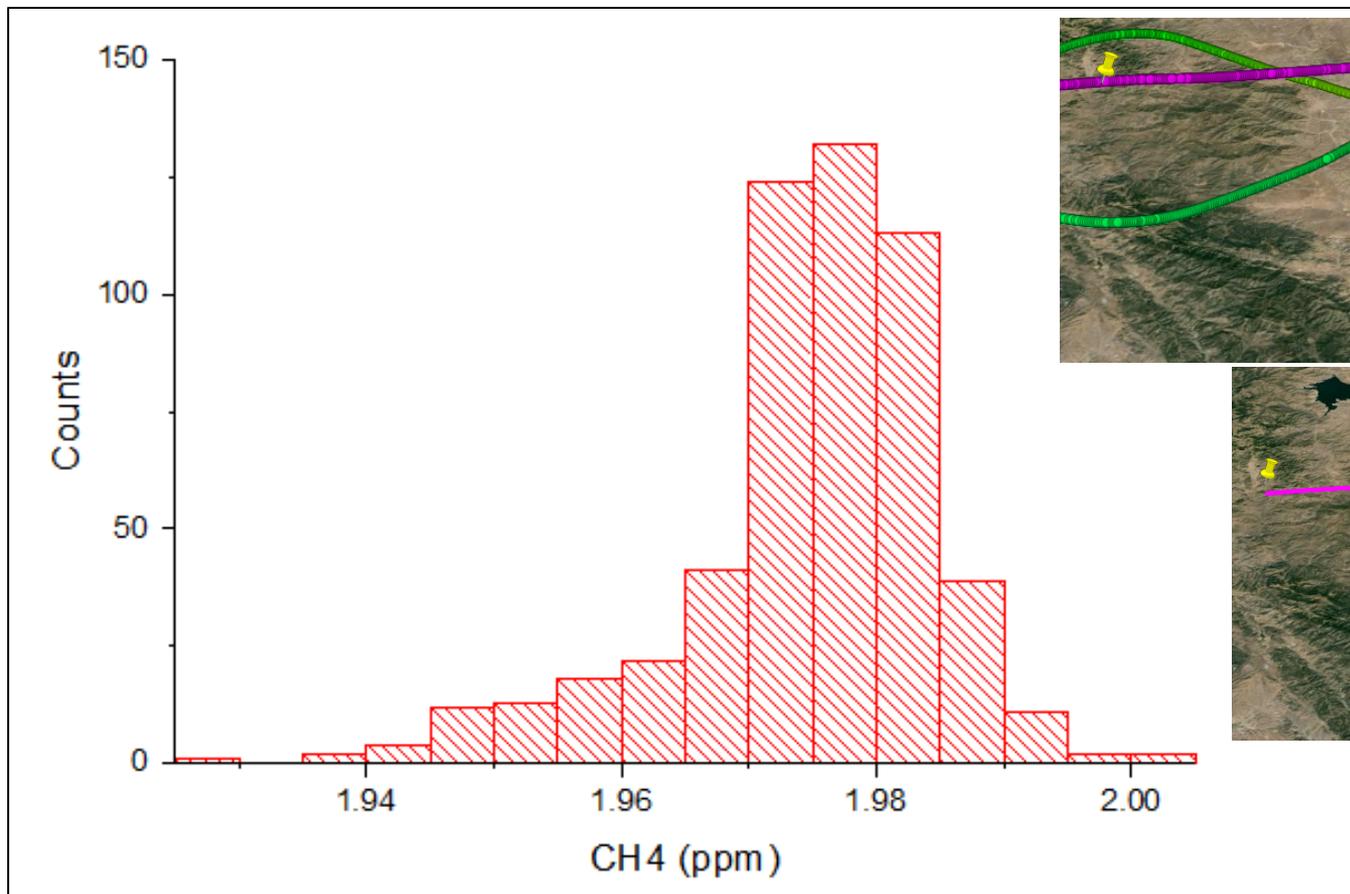


# OPO Precision 1s averaging\*



Descriptive Statistics			
	N total	Mean	Standard Deviation
KNCH4	536	1.97477	0.01008

~0.5%



\*Using K. Numata's retrievals



# Airborne Demonstration Summary



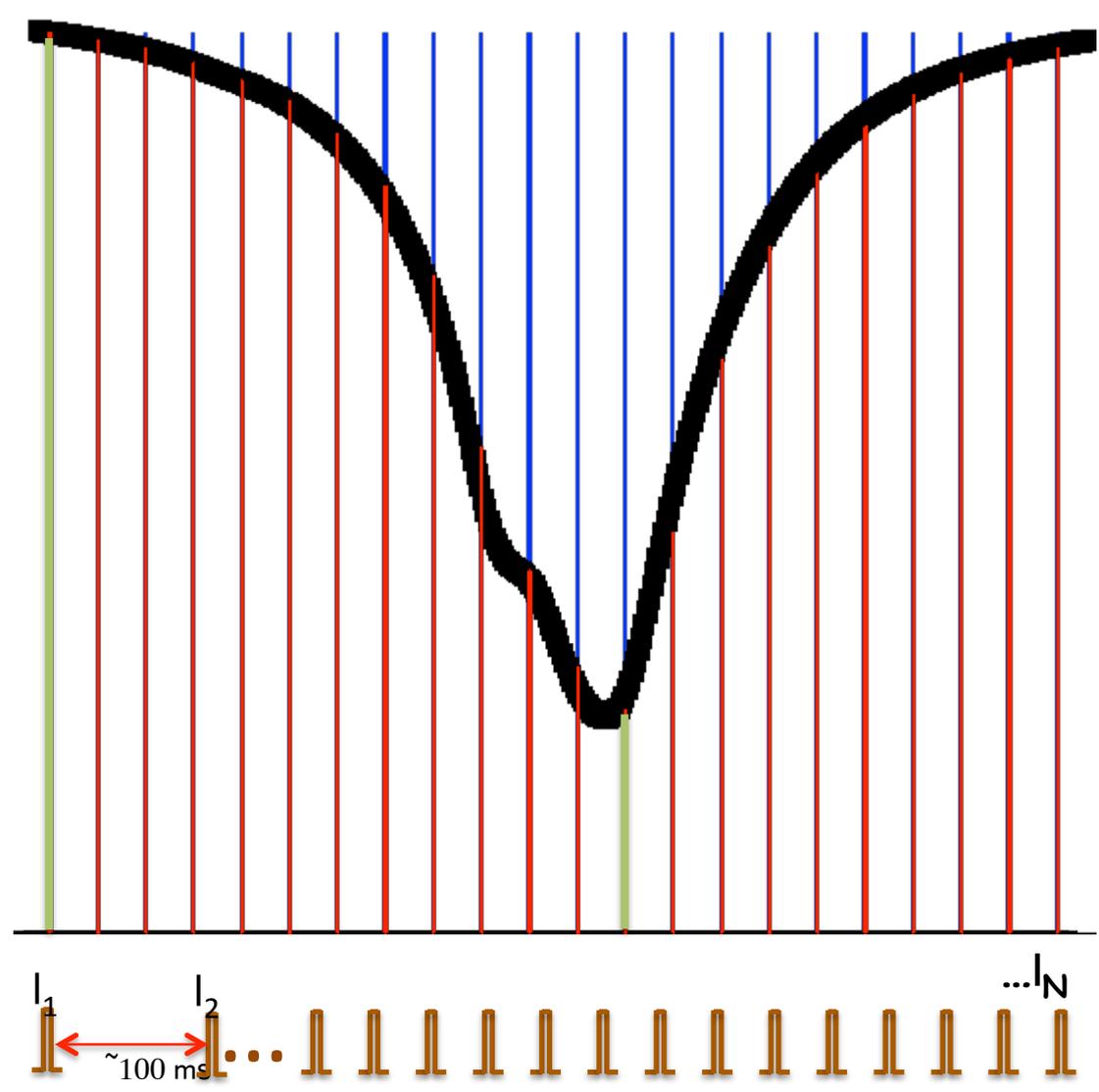
- ✓ DRS e-ADP works very well at 1651 nm.
- ✓ Responsivity is close to what we have for CO<sub>2</sub> but there is still some uncertainty on the actual value.
- ✓ DRS is linear over at least two to three orders of magnitude.
- ✓ Signal strength (link margin) during flight is roughly what we expected. Return signal estimation is complicated by burst pulse shape of the OPA and the high OPO energy and clouds/topography.
- ✓ Both OPA and OPO worked better than expected.
- ✓ Best precision for: OPA ~ 0.4% and OPO ~ 0.5%.
- ✓ 20 wavelengths (OPA) produced better fits than 5 (OPO) but OPA was extremely sensitive to LMA fiber and cabin temperature fluctuations and does not have adequate energy for space unless external amplifiers are used and/or the present OPA laser is improved.
- ✓ OPO cavity is sensitive to temperature and pressure but still worked well.
- ✓ OPO has the required (high) energy but the OPO flight was complicated by our inability to effectively attenuate the laser.
- ✓ OPO energy was more stable than OPA during flight.



# Summary

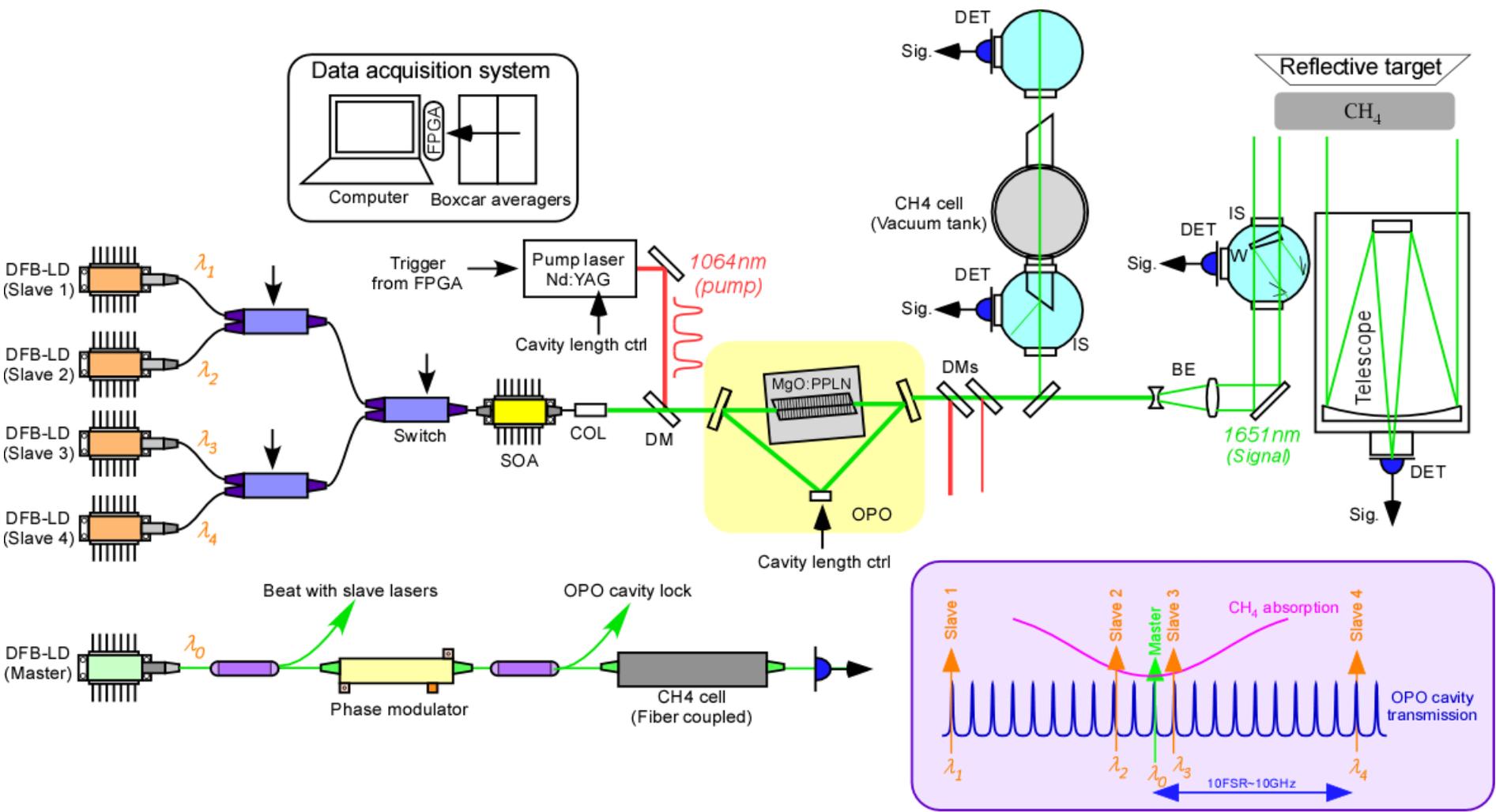


- ✓ Many different approaches and options for power scaling investigated
- ✓ Leveraged IRAD & ACT programs.
- ✓ Demonstrated two viable architectures for a CH<sub>4</sub> transmitter using OPA and multi-wavelength OPO
- ✓ Demonstrated power scaling to 250-290 μJ.
- ✓ Demonstrated and validated CH<sub>4</sub> airborne measurements using the two lidar transmitters.
- ✓ Other laser transmitter options being investigated
- **We would like to thank ESTO and GSFC IRAD for their support.**



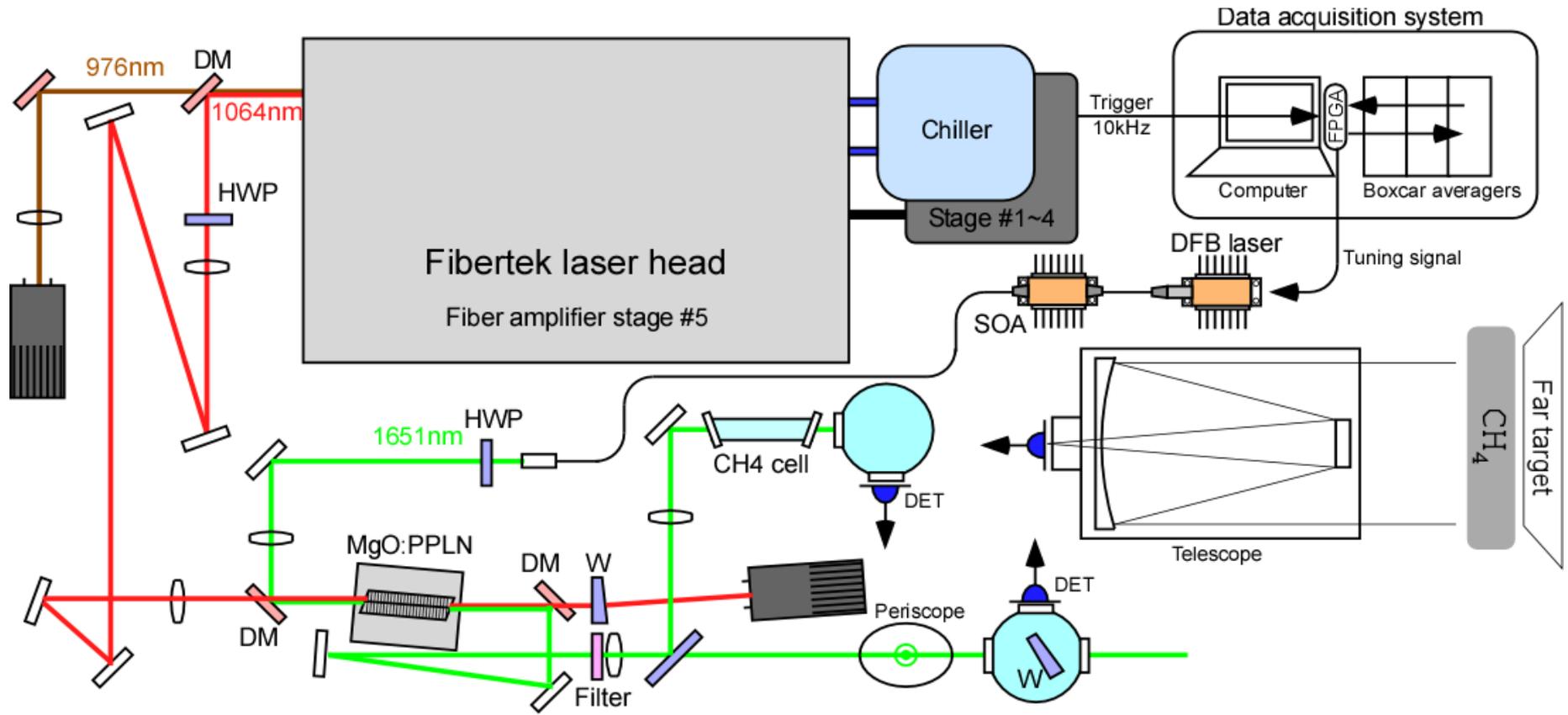


# Setup for 5-wavelength OPO



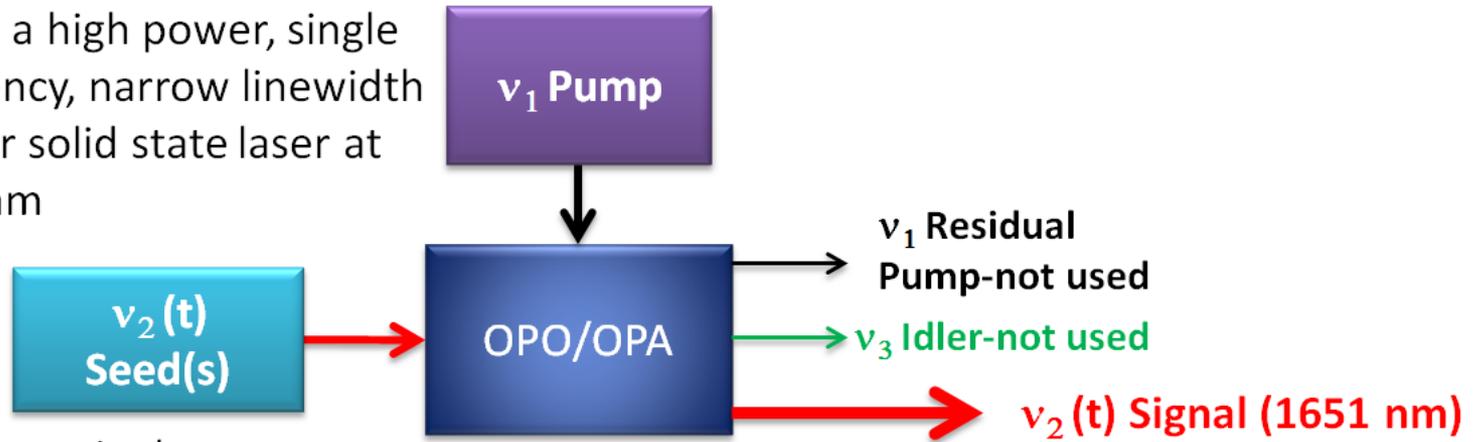


# OPA Open-path measurement setup





**Pump:** a high power, single frequency, narrow linewidth fiber or solid state laser at 1064 nm

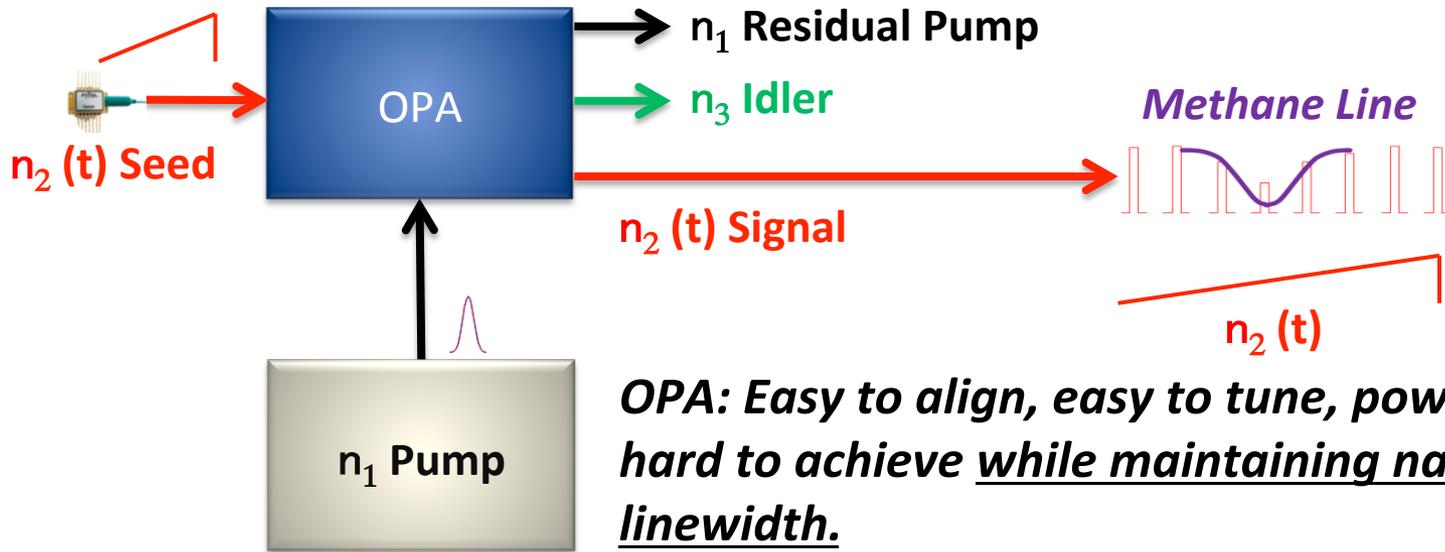


**Seed:** a low power, single frequency diode laser at 1651 nm.

Optical Parametric Oscillator (OPO) or Optical Parametric Amplifier (OPA). A non-linear crystal that amplifies the seed laser to the energy needed for space (250-300  $\mu$ J) **without** degrading the spectral characteristics

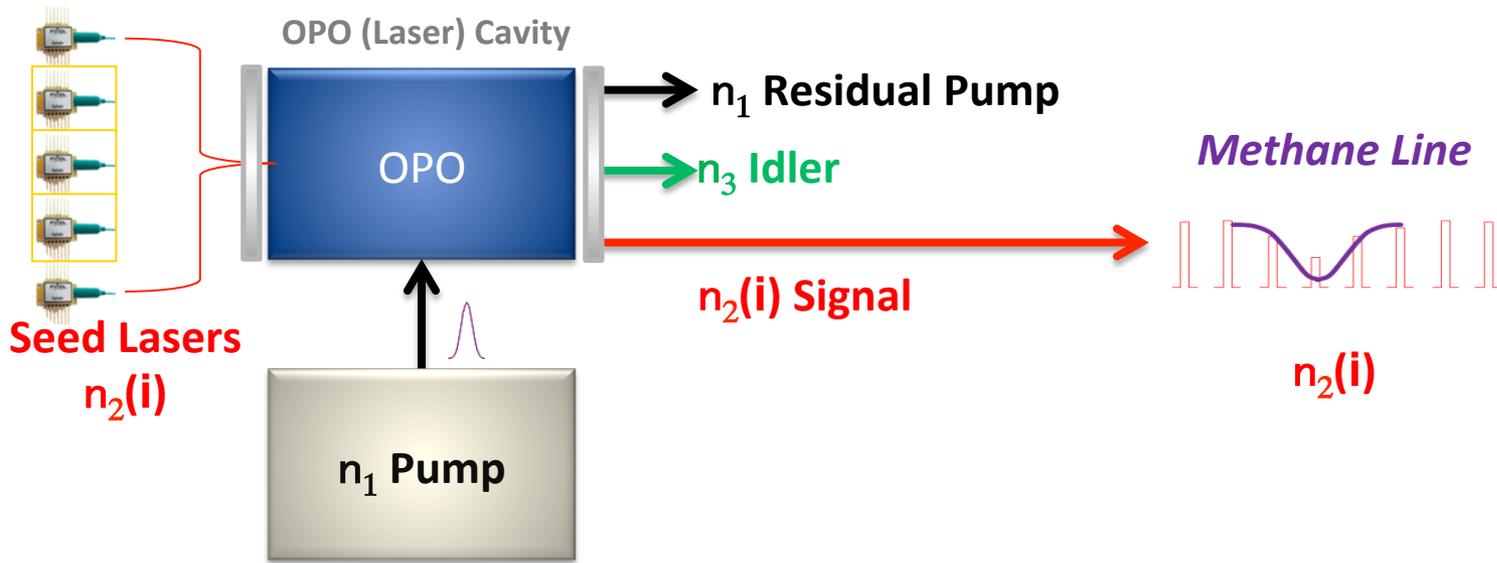


Used OPO/OPAs to measure CH<sub>4</sub> at near and mid IR, CO<sub>2</sub>, H<sub>2</sub>O and CO



***OPA: Easy to align, easy to tune, power scaling hard to achieve while maintaining narrow linewidth.***

***OPA samples the CH<sub>4</sub> line at several wavelengths using a single, continuously tuned seed laser***



***OPO: Complicated to align and tune; power scaling easier to achieve while maintaining narrow linewidth.***  
***OPO samples the CH<sub>4</sub> line at several discrete wavelengths using multiple seed lasers.***  
***All lasers must be locked.***



# Current GSFC Power scaling options OPA/OPO



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Er:YAG and Er:YGG