

---

# From Incubator to Orbit

---

## Ten Years of Infrared Sensor Technology Development at NASA Langley

Marty Mlynczak, David Johnson, Richard Cageao, M. Nurul Abedin, David Kratz  
NASA Langley Research Center



---

ESTO Technology Forum, June 22 2010

---

# Outline

---

- **Review of Earth's Energy Budget and Infrared Greenhouse Effect**
- **Measurement of IR spectra at top (and bottom) of the atmosphere**
- **Measurement of IR spectra within the atmosphere**
- **On to orbit – to the CLARREO mission**
- **Summary**



---

## Acknowledgements

---

- **Earth Science Technology Office (G. Komar)**
- **Radiation Sciences Program (H. Maring)**
- **NASA Langley Research Center (SD, SED, RTD)**
- **Corporate, Academic, and Government Partners**
  - **Space Dynamics Laboratory, Utah State University**
  - **Harvard Smithsonian Center for Astrophysics**
  - **DRS Technologies, Cypress, CA**
  - **Department of Energy**
  - **University of Wisconsin**
  - **AER, Inc.**
  - **SSAI, Hampton VA**
  - **ITT, Fort Wayne, IN**
  - **Raytheon Vision Systems, Santa Barbara, CA**



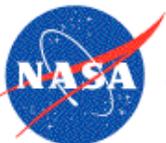
---

# Projects

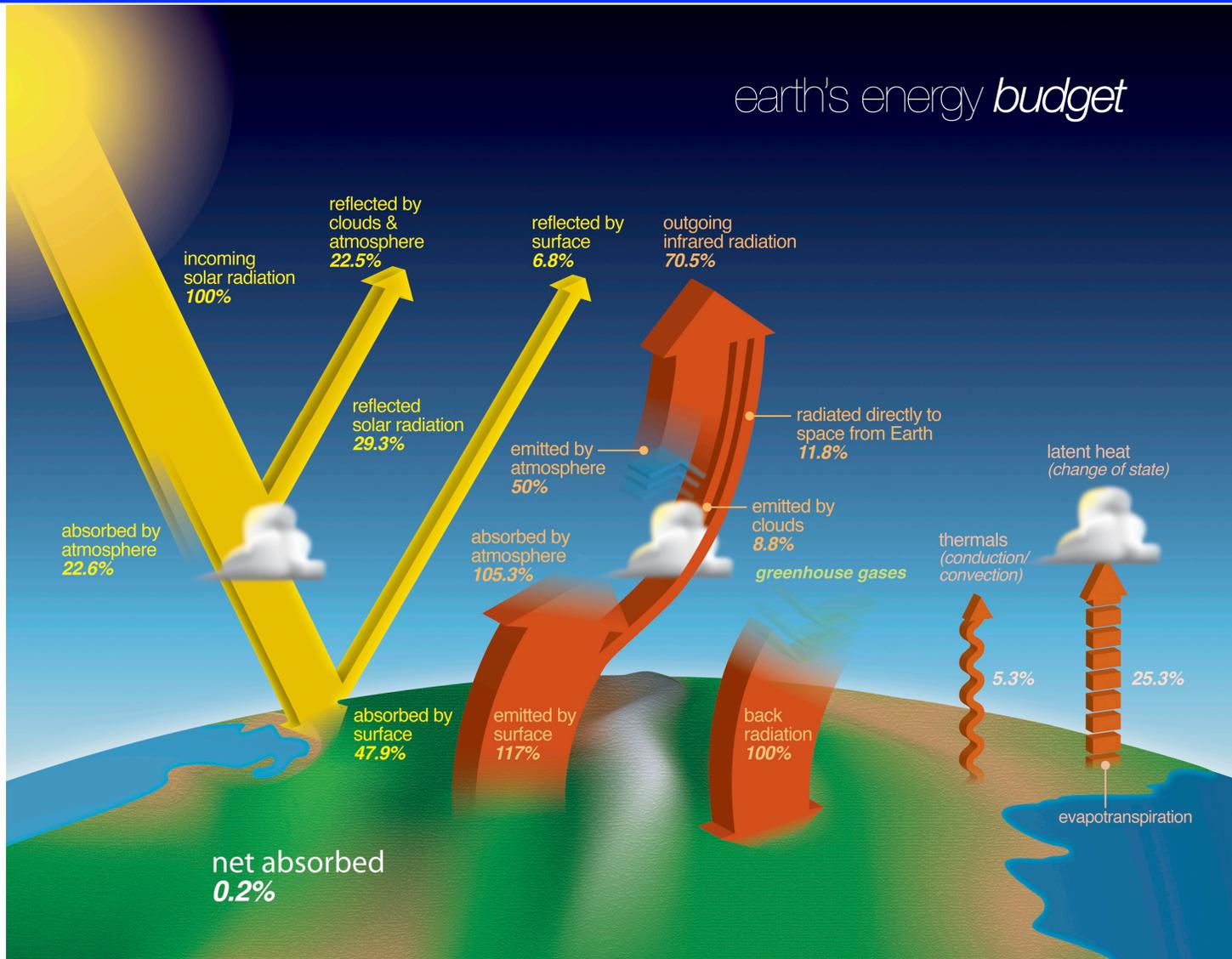
---

- **IIP 2001: FIRST**
  - Far-Infrared Spectroscopy of the Troposphere
- **IIP 2004: INFLAME**
  - In-situ Net Flux within the Atmosphere of the Earth
- **IIP 2007: CORSAIR**
  - Calibrated Observations of Radiance Spectra from the Atmosphere in the Infrared
- **ACT 2008: FIREBIB**
  - Far-Infrared Blocked Impurity Band Detectors
- **Radiation Sciences 2008: FORGE**
  - Far-Infrared Observations of the Radiative Greenhouse Effect

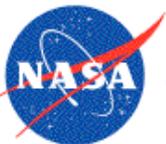
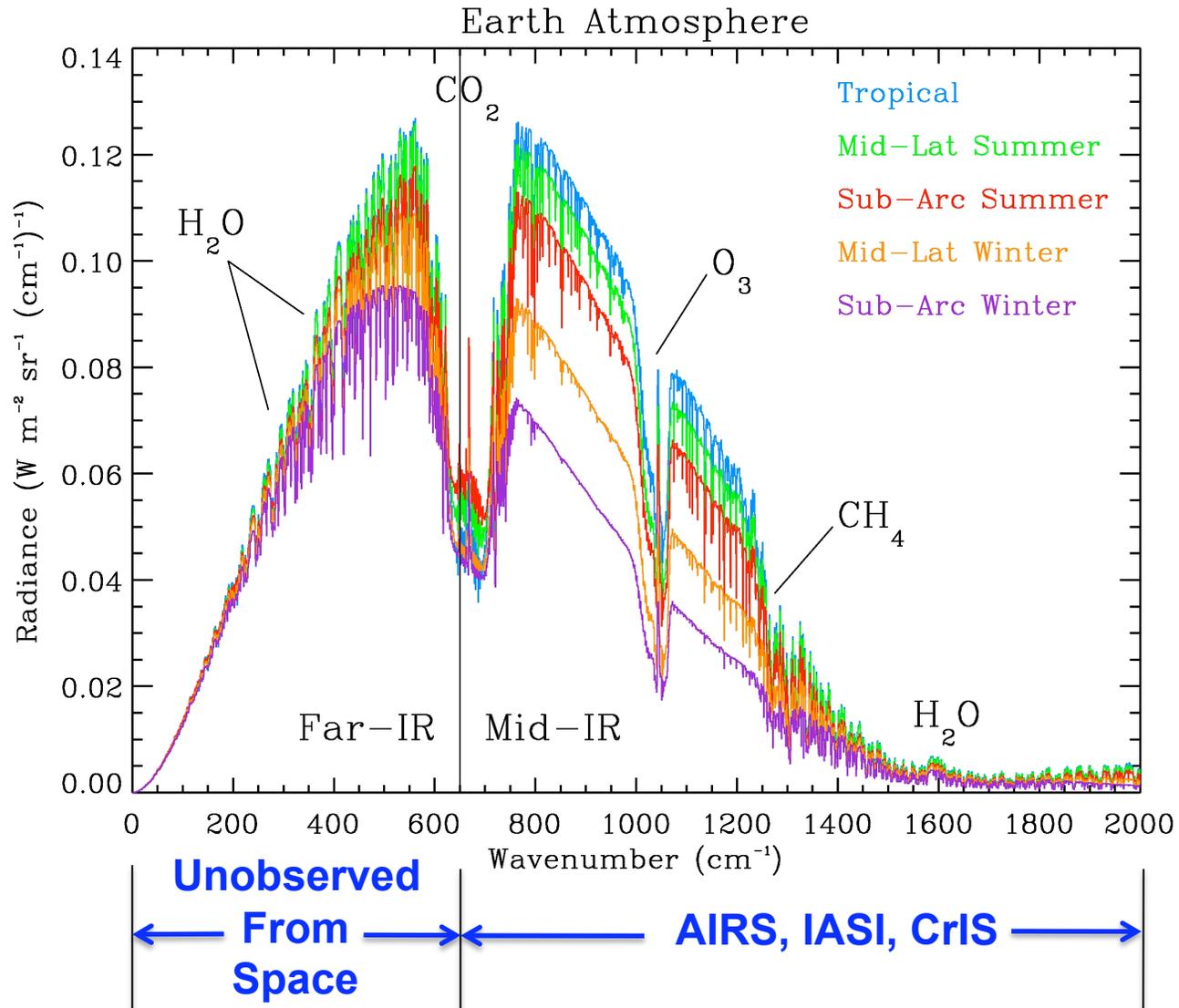
***Goal: To significantly advance ability to observe the infrared spectrum from space, within the atmosphere, and from the ground***



# Earth's Energy Budget



# Earth's Outgoing Longwave Radiation



---

# FIRST Instrument

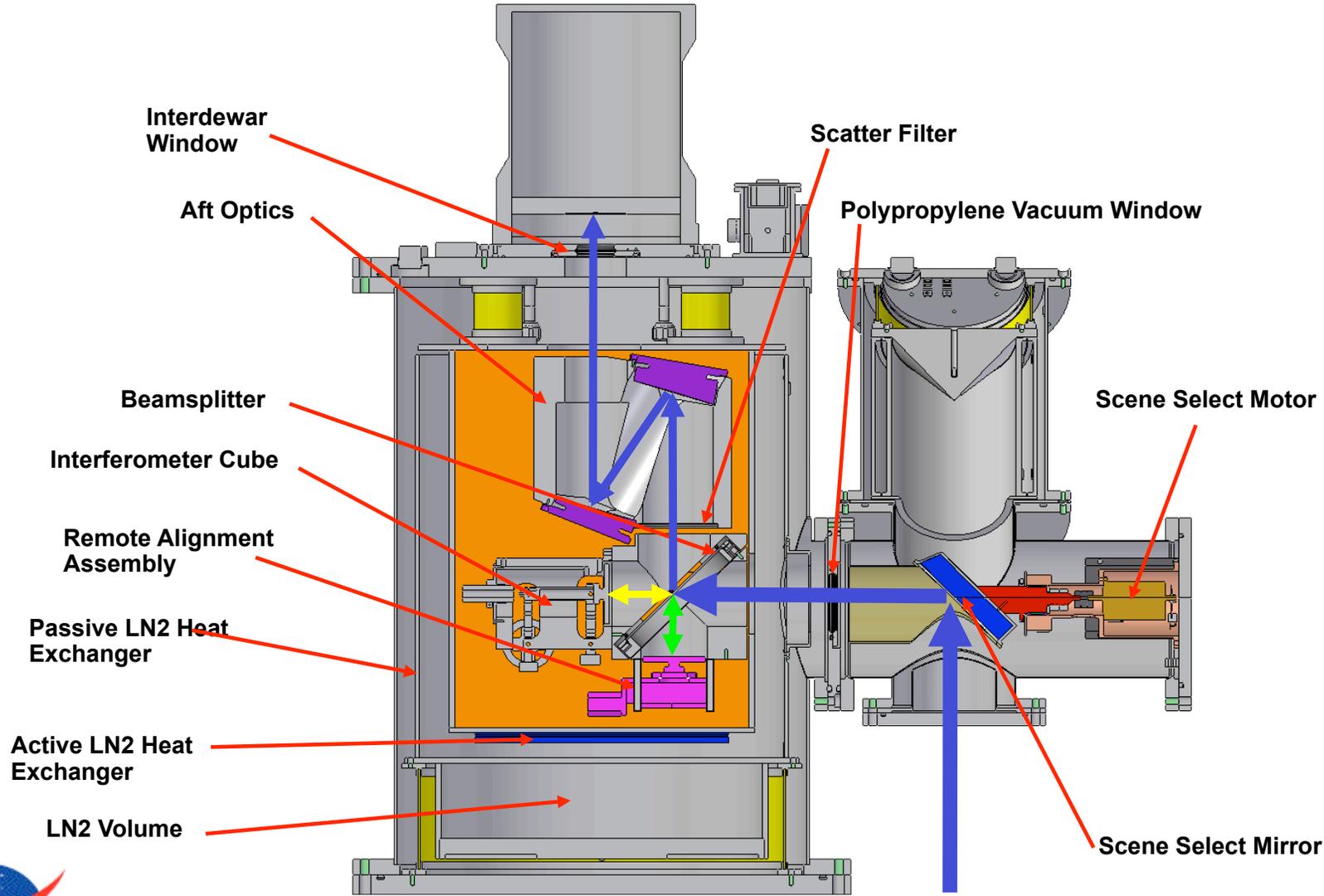
---

## FIRST - Far-Infrared Spectroscopy of the Troposphere

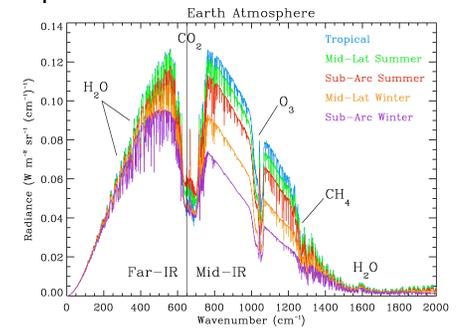
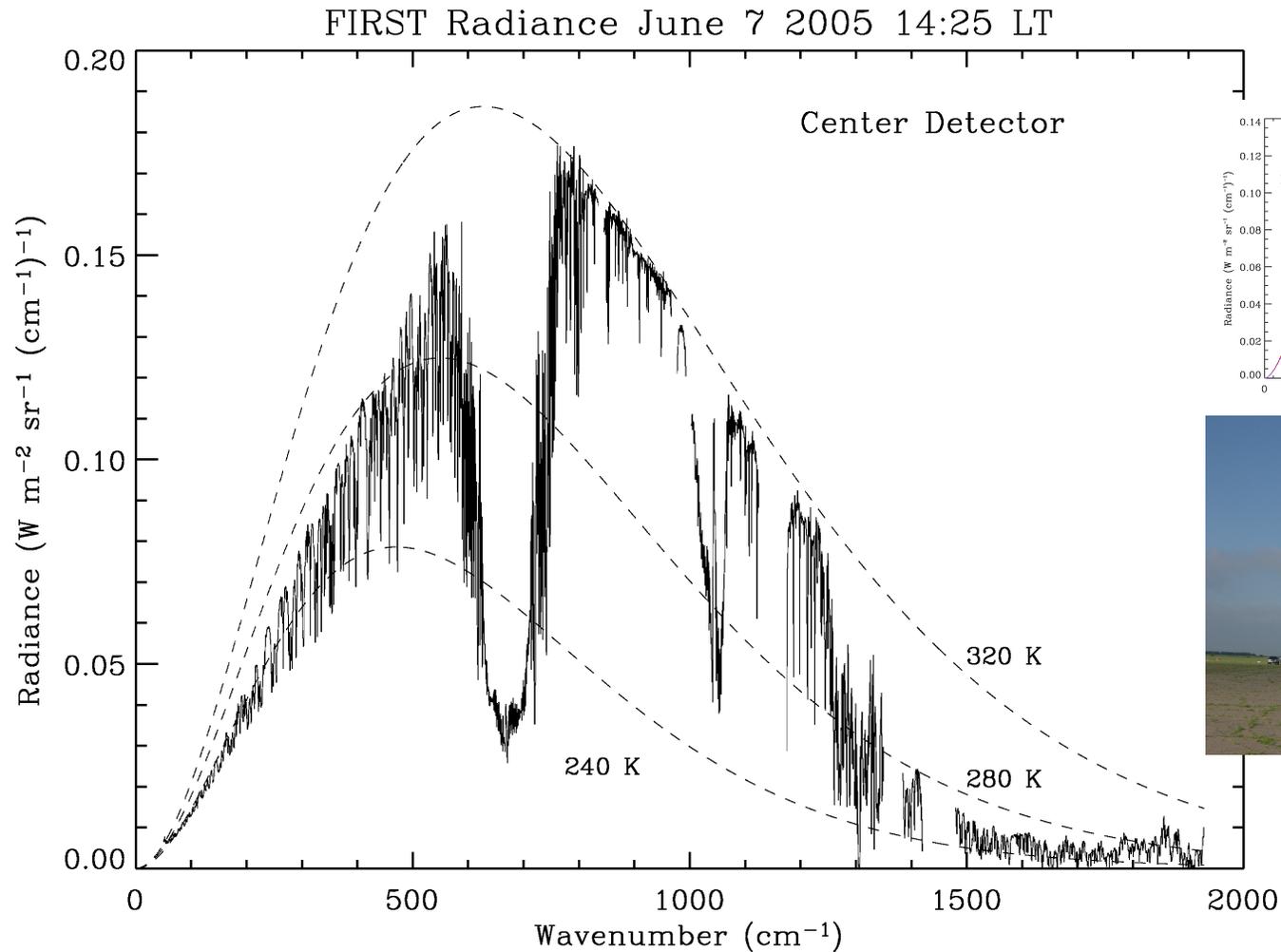
- Developed under NASA Instrument Incubator Program of ESTO
- Michelson Interferometer
  - *6 to 100  $\mu\text{m}$  on a single focal plane*
  - *0.625  $\text{cm}^{-1}$  unapodized (0.8 cm OPD)*
  - *Germanium on polypropylene beamsplitter*
  - *Bolometer detectors @ 4 K*
- Demonstrated on a high-altitude balloon flight June 7 2005
- Second balloon flight September 18 2006
- Ground-based capability demonstrated March 2007
- Selected by NASA for FORGE Campaign in Chile in 2008
- FORGE Campaign August – October 2009



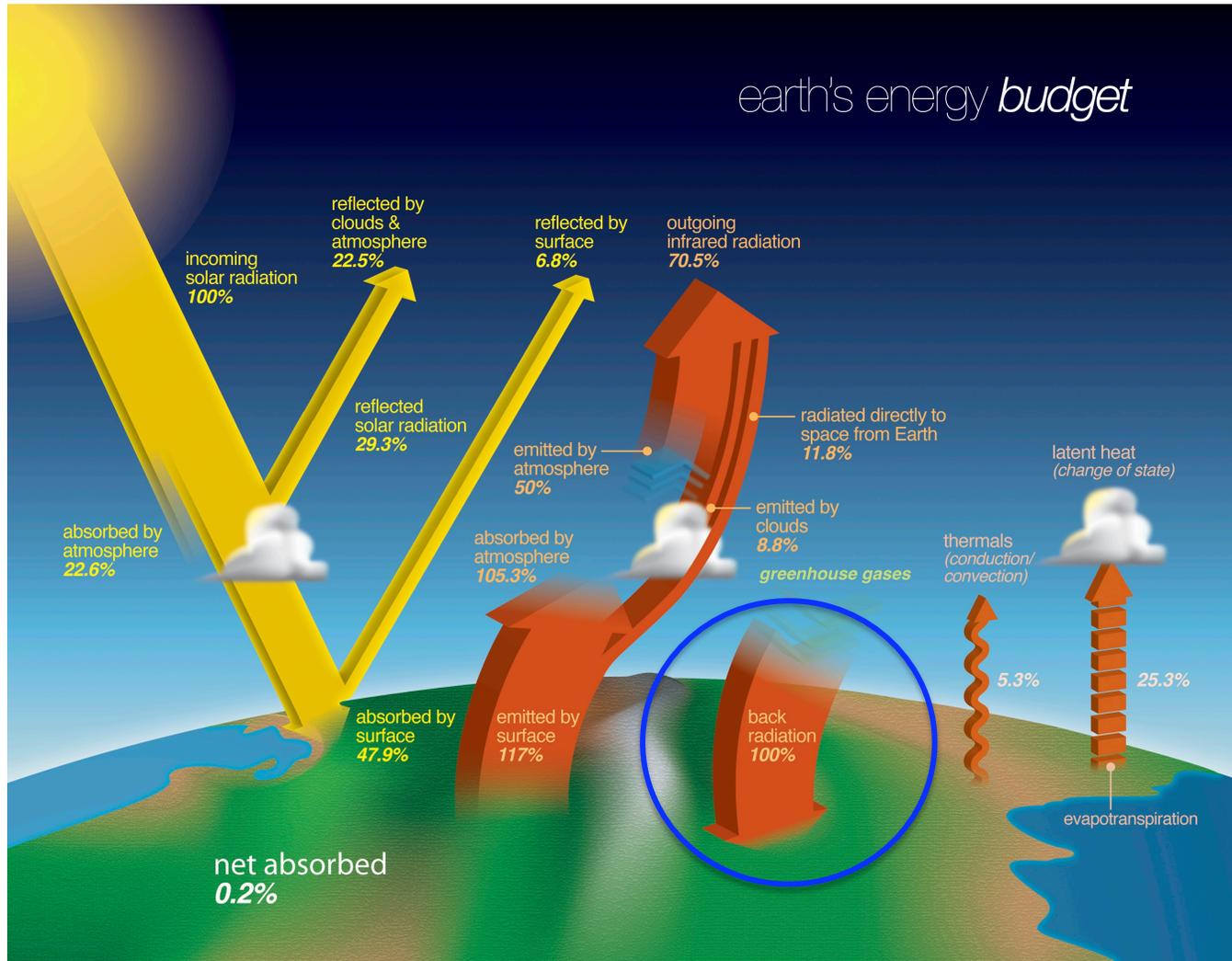
# FIRST Spectrometer Overview



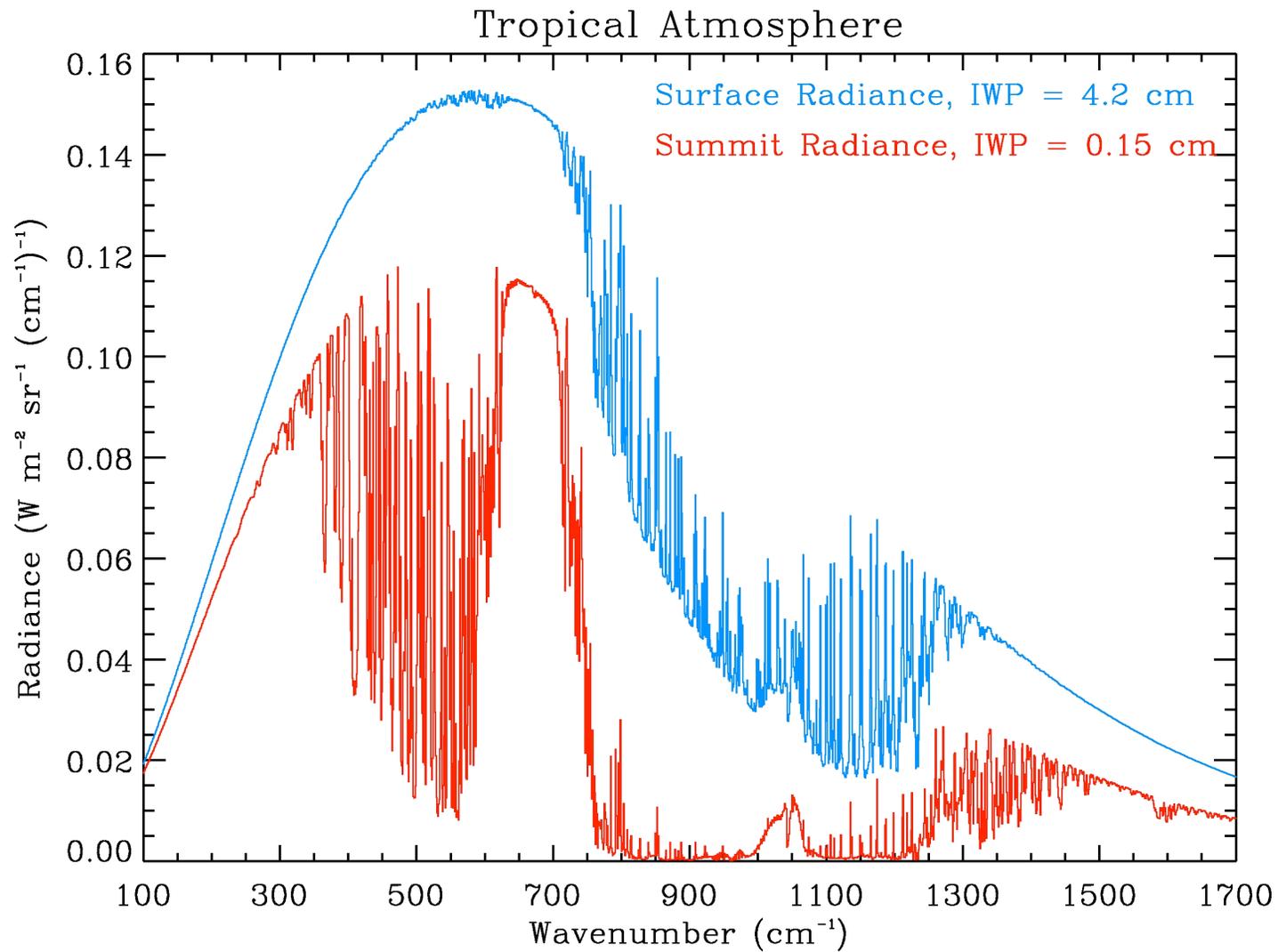
# FIRST Thermal Infrared Spectrum - TOA



# Earth's Energy Budget



# Earth's Downwelling Infrared Radiation – at Surface



---

# The FORGE Project – Scientific Objective

---

- Use new observational capability provided by FIRST to directly observe and quantify entire infrared greenhouse effect – *at Earth's surface*
- Conduct “radiative closure” experiment
  - Measure infrared radiation emitted by the atmosphere – *at Earth's surface looking up*
  - Simultaneously measure temperature and water vapor profiles to provide inputs for theory
- Conduct experiment at high, dry location – Cerro Toco, Atacama Desert, Chile
  - Altitude: 17, 600 feet above sea level. Mean pressure ~ 0.5 Atmosphere
  - *Precipitable Water Vapor < 1 millimeter*
- Part of larger “RHUBC-II” Campaign run by Dept. of Energy
  - Teams from U. Wisconsin, AER Inc., Italy, Germany, PNNL, Los Alamos NL, U. Denver, NASA Langley

*Safety of Team during deployment at altitude was paramount*



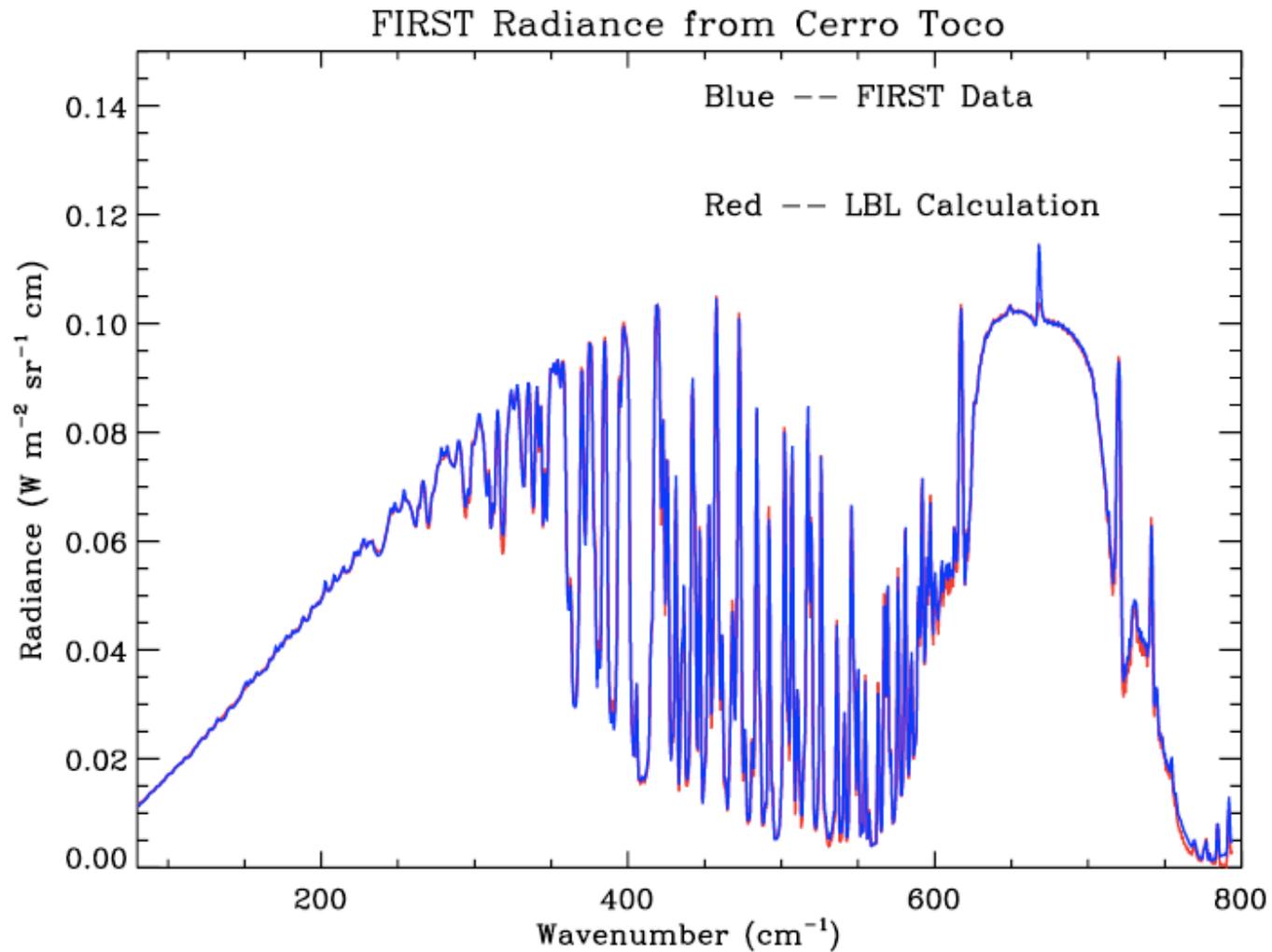
---

# Operations at 17,600 Feet

---



# FIRST Far-IR Spectrum – 09/05/09

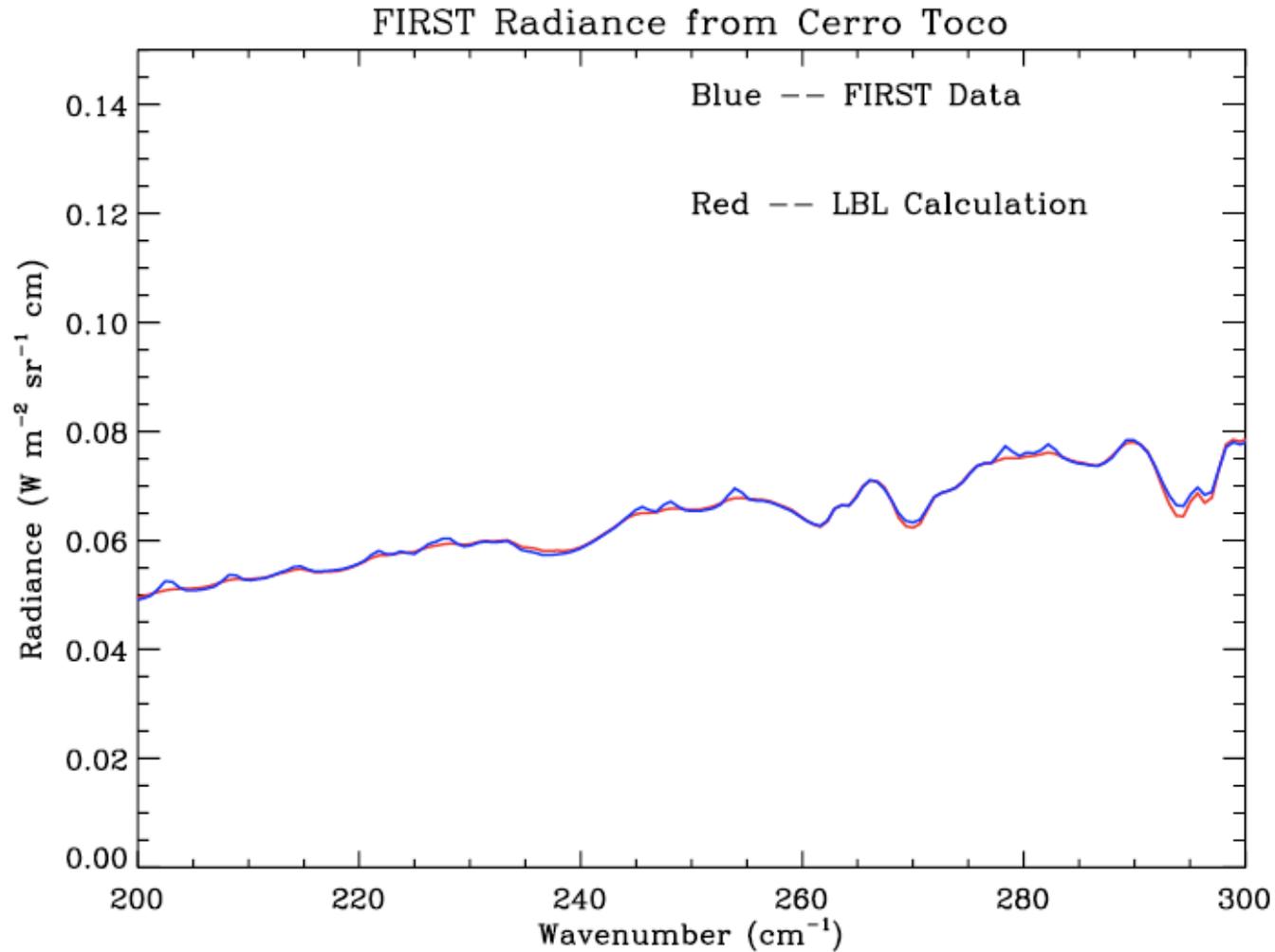


PWV = 0.75 millimeter (“wet” day)

---

# September 5 2009 – PWV = 0.75 mm

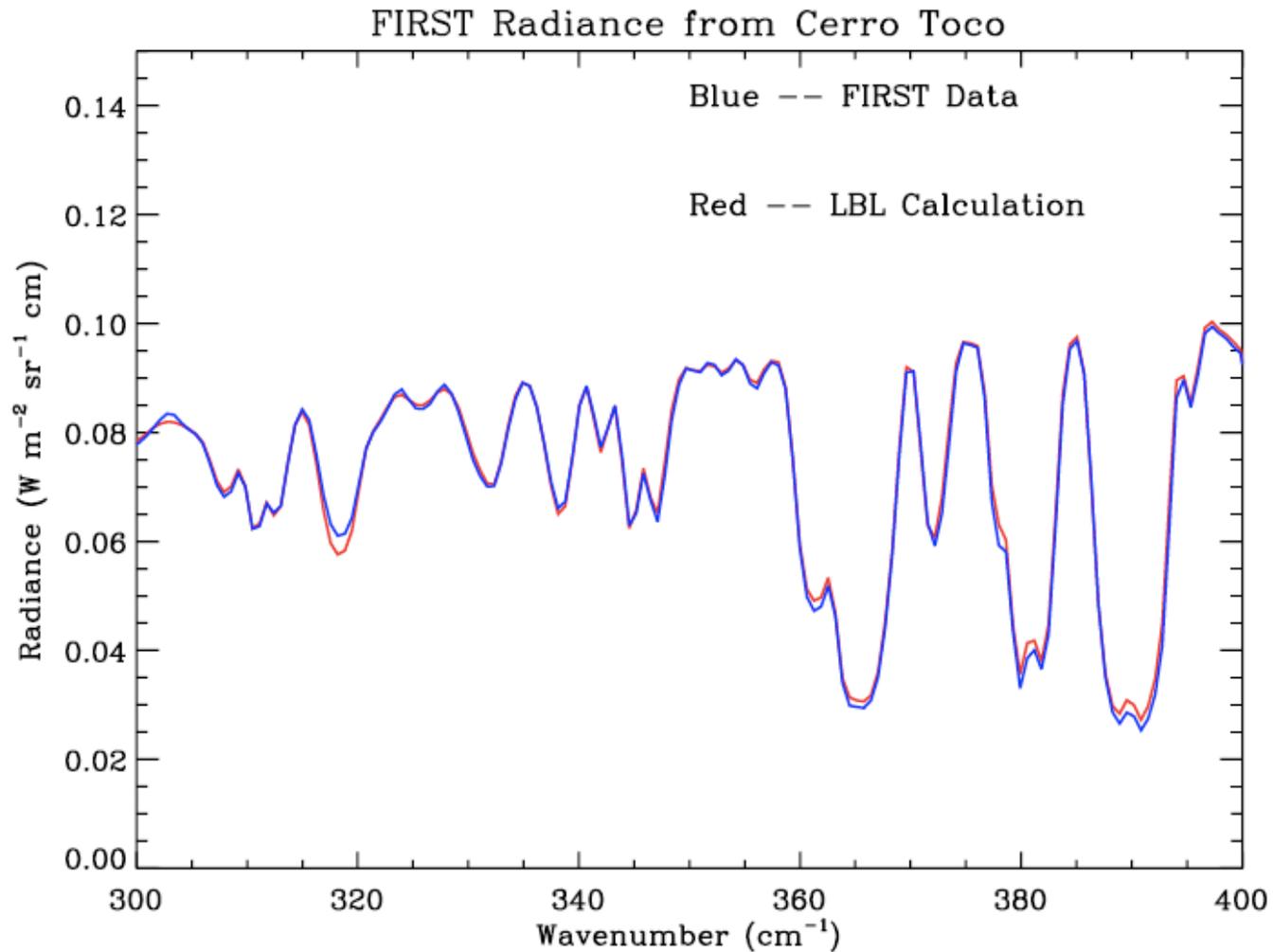
---



---

# September 5 2009 – PWV = 0.75 mm

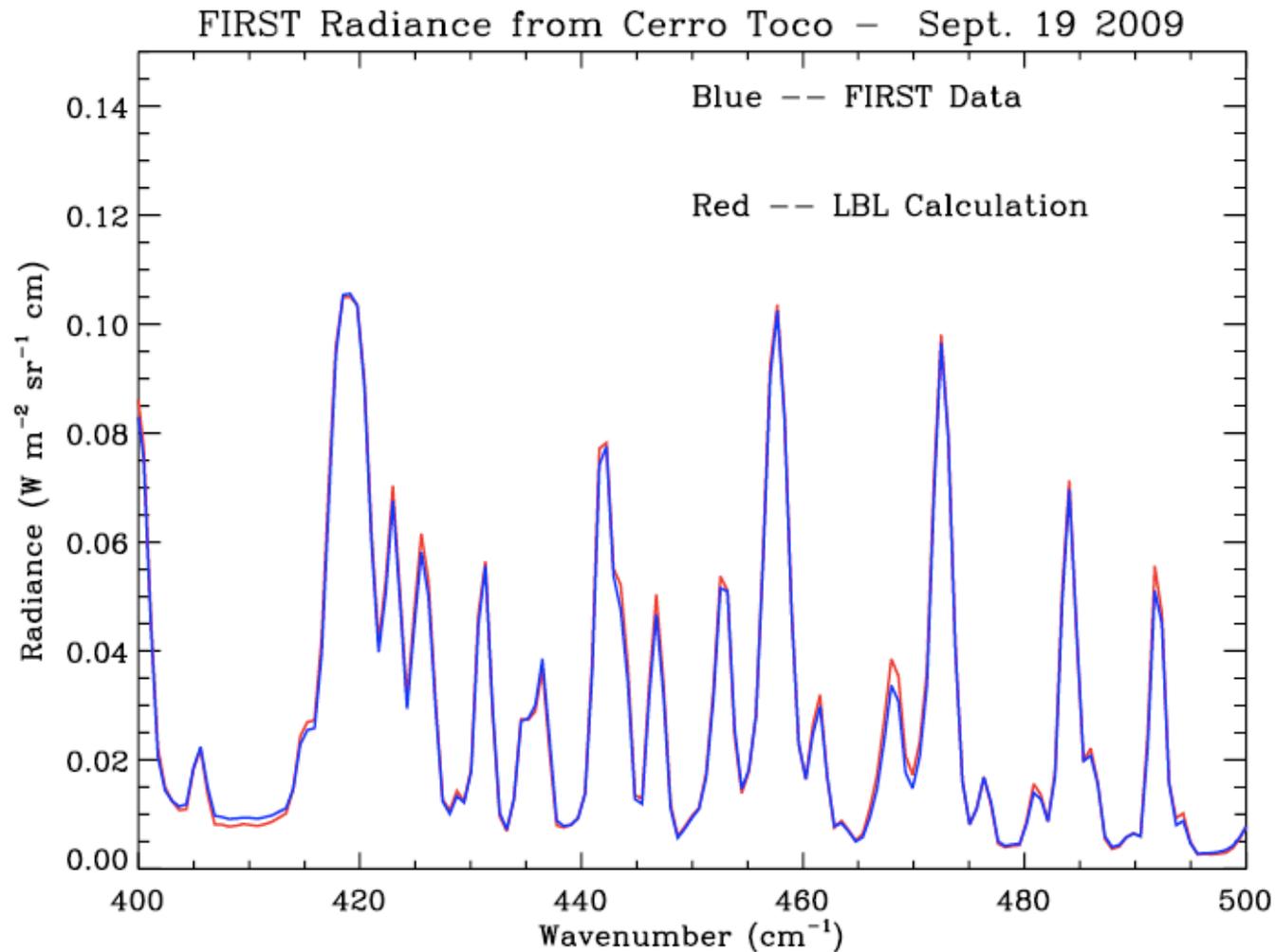
---



---

# September 19 2009 – PWV = 0.4 mm

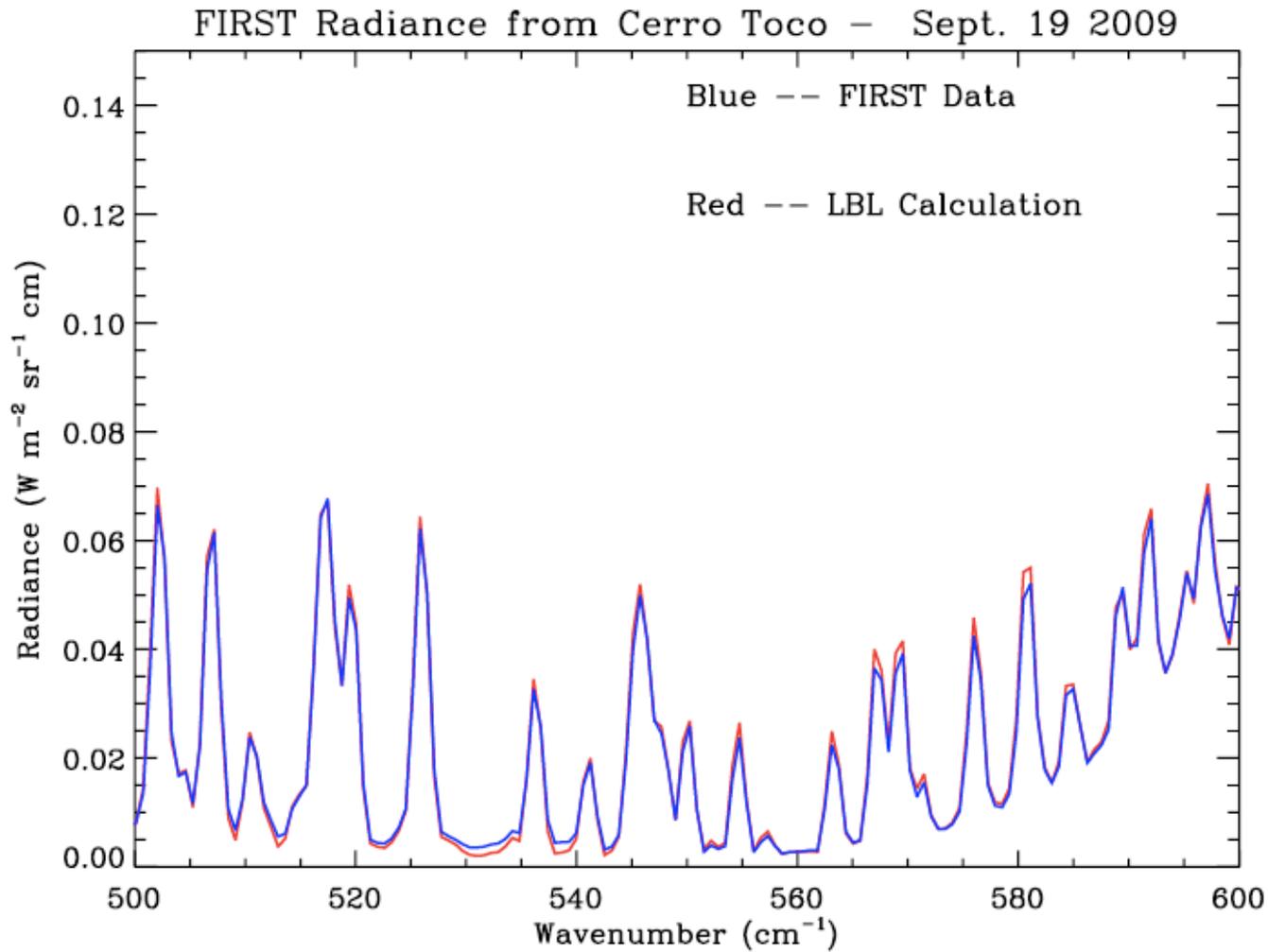
---



---

# September 19 2009 – PWV = 0.4 mm

---



---

# FIRST Summary

---

- Successfully developed unique instrument to measure, at high spectral resolution, the complete infrared spectrum of the Earth and atmosphere
- Transitioned from “Incubator” to operational science instrument
- Demonstrated technologies in:
  - FTS
  - Beamsplitter
  - Focal Plane
- *Onto INFLAME and measurement of IR radiation within the atmosphere*



# INFLAME – Fundamental Aspects

---

Every atmospheric model has essentially 3 equations:

- Momentum: ( $F = ma$ )
- Continuity: (Conservation of mass)
- Energy: (First Law of Thermodynamics)

Energy equation requires knowledge of rate at which atmosphere heats and cools:

- Radiation
- Latent process (water condensation/evaporation)
- Conduction

**INFLAME Goal: Measure the rates of heating & cooling of the atmosphere by visible and infrared radiation**

# In-Situ Net Flux within the Atmosphere of the Earth

---

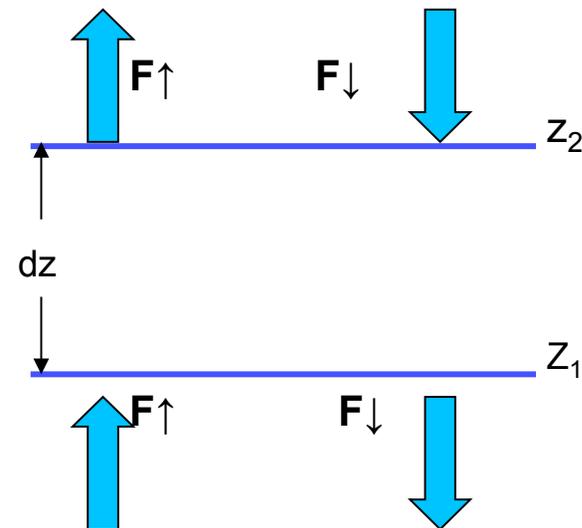
## Atmospheric Heating and Cooling Rates

$$F \uparrow (z) = \int_0^1 I(z) \mu d\Omega$$

$$F \downarrow (z) = \int_{-1}^0 I(z) \mu d\Omega$$

$$F_{net}(z) = F \uparrow (z) - F \downarrow (z)$$

$$\frac{\partial T}{\partial t} = \frac{1}{\rho C_p} \frac{\partial F_{net}(z)}{\partial z}$$



Require an instrument to **directly** measure the net flux

---

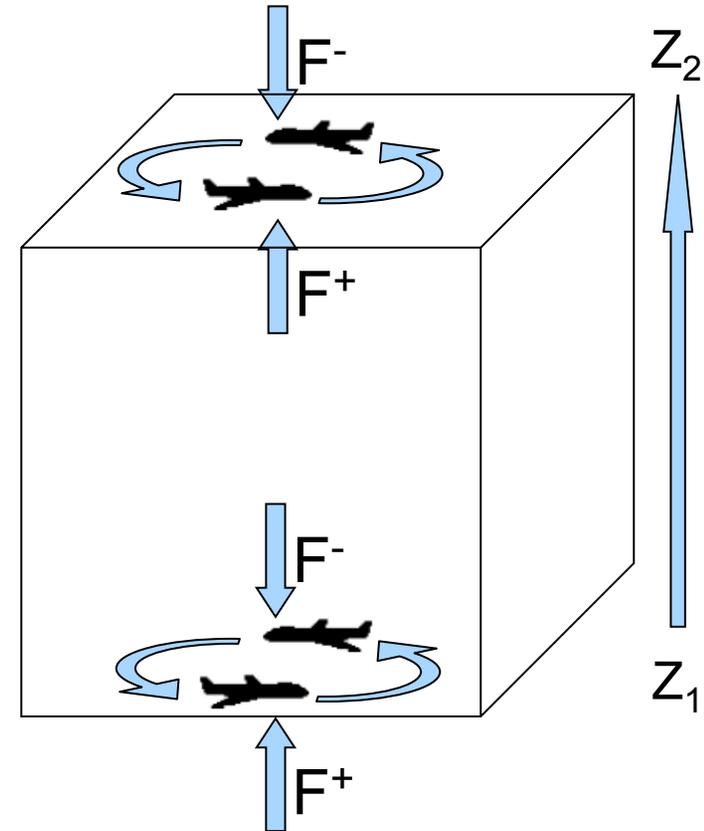
# Measuring Net Flux Divergence



- From an airborne platform:
  - Measure Net Flux at  $Z_1$
  - Measure Net Flux at  $Z_2$

- Approximate Heating Rate by:

$$\frac{\partial T}{\partial t} = \frac{1}{\rho C_p} \frac{F_{net}(z_2) - F_{net}(z_1)}{(z_2 - z_1)}$$

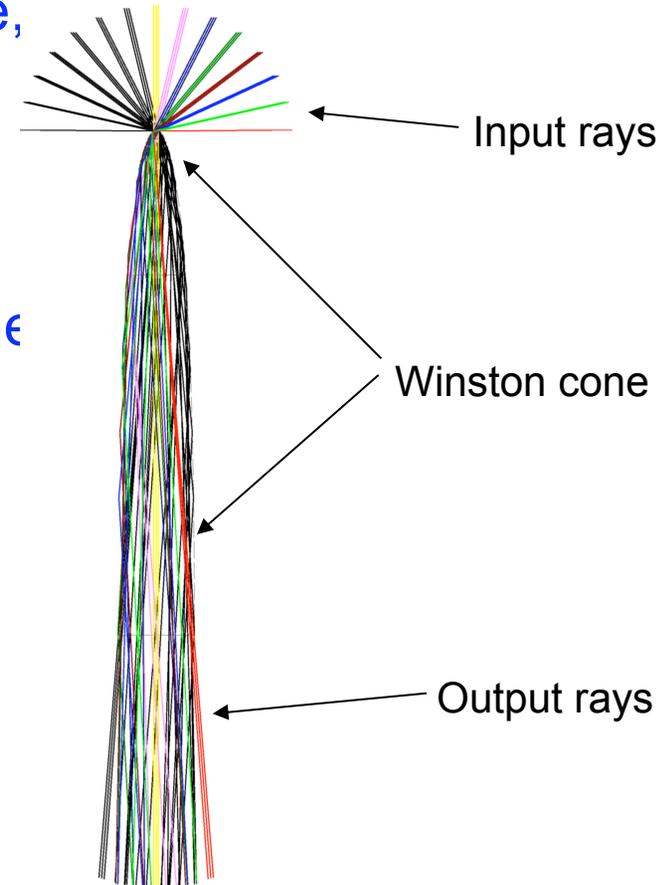




Langley Research Center

# INFLAME Approach: F+ and F-

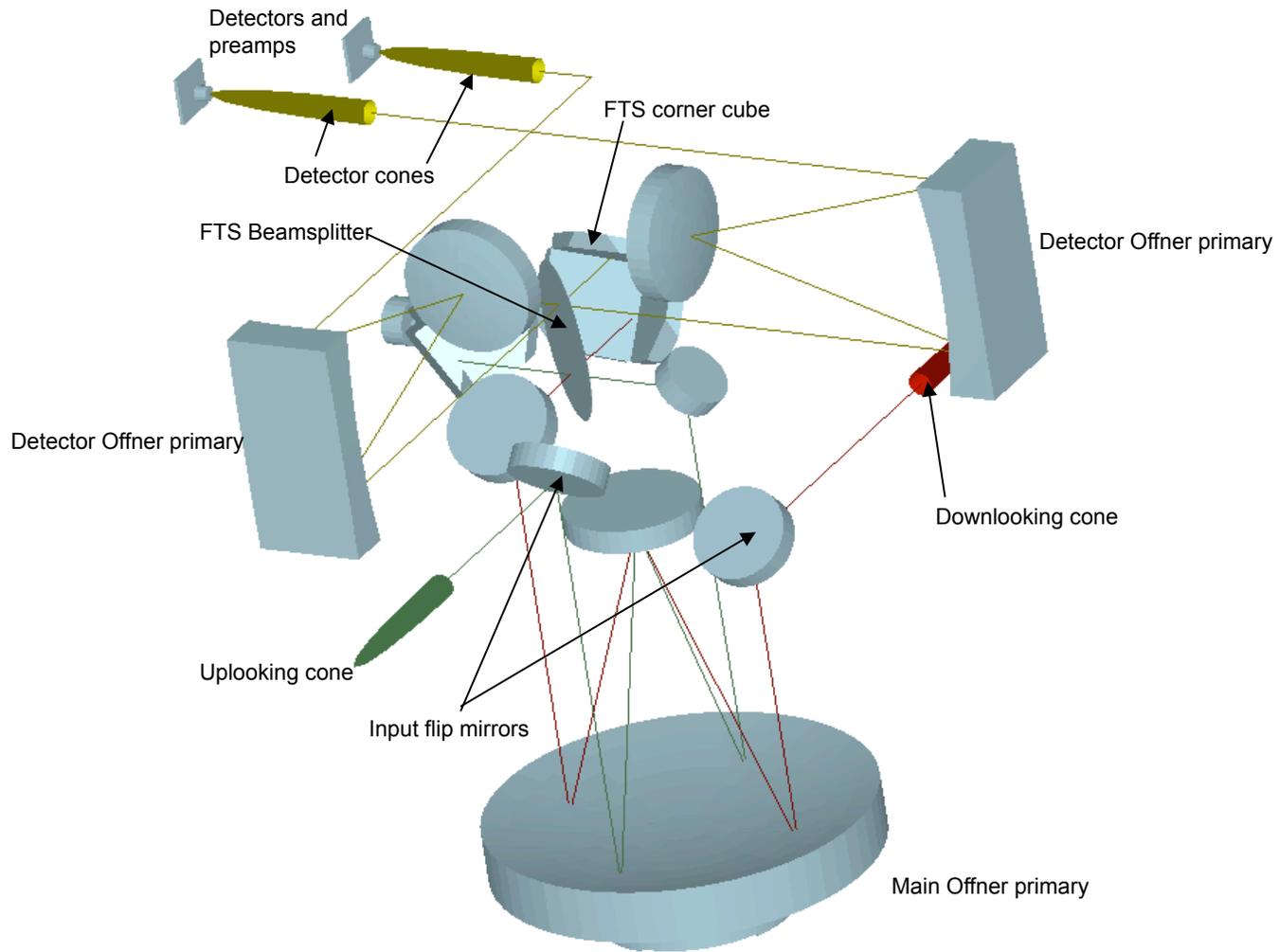
- Most instruments measure radiance, not flux.
- Measuring flux requires collecting light over a full hemisphere.
- We use a non-imaging Winston cone to collect radiation and collimate it into an f/6.8 beam.
  - Input aperture is 1 mm diameter.
  - Output aperture is 13.6 mm diameter.





Langley Research Center

# INFLAME Optical Layout

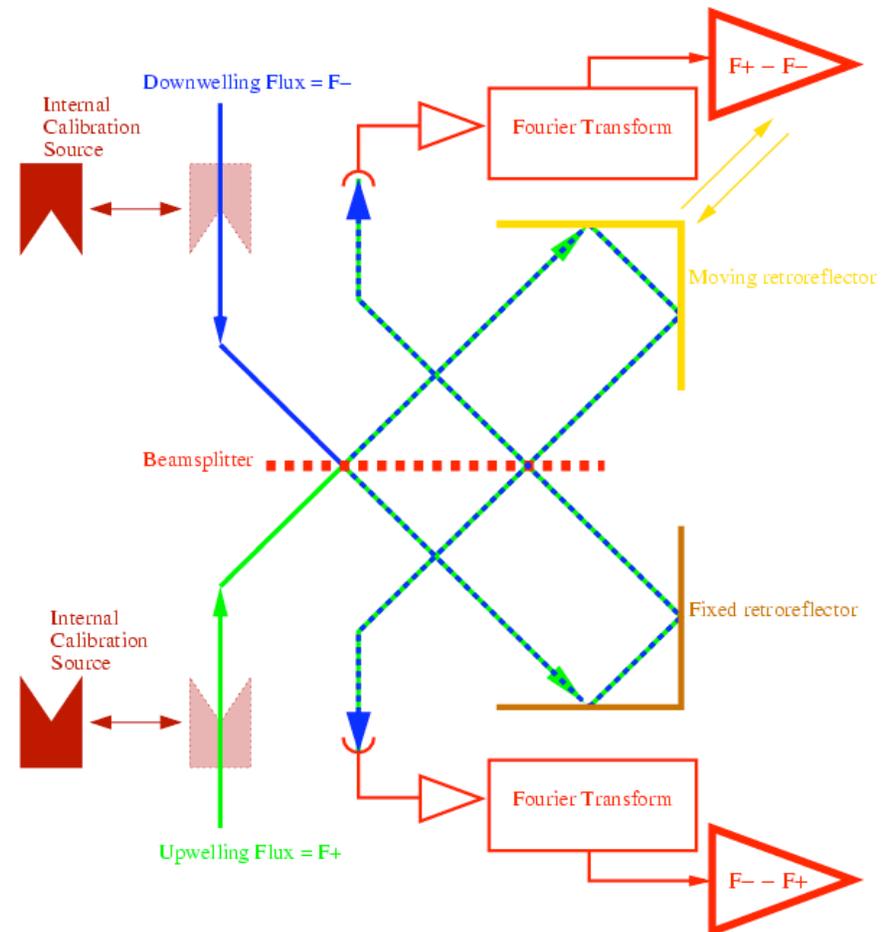




# INFLAME Approach: Fz

- Direct uplooking and downlooking apertures to the two inputs of a 4-port Fourier transform spectrometer (FTS);
- Scan FTS to produce complimentary interferograms at two outputs;
- Fourier transform interferograms to estimate the spectrum of the net flux

$$F_z = F_+ - F_-$$





Langley Research Center

# INFLAME Measurement Platform

Learjet



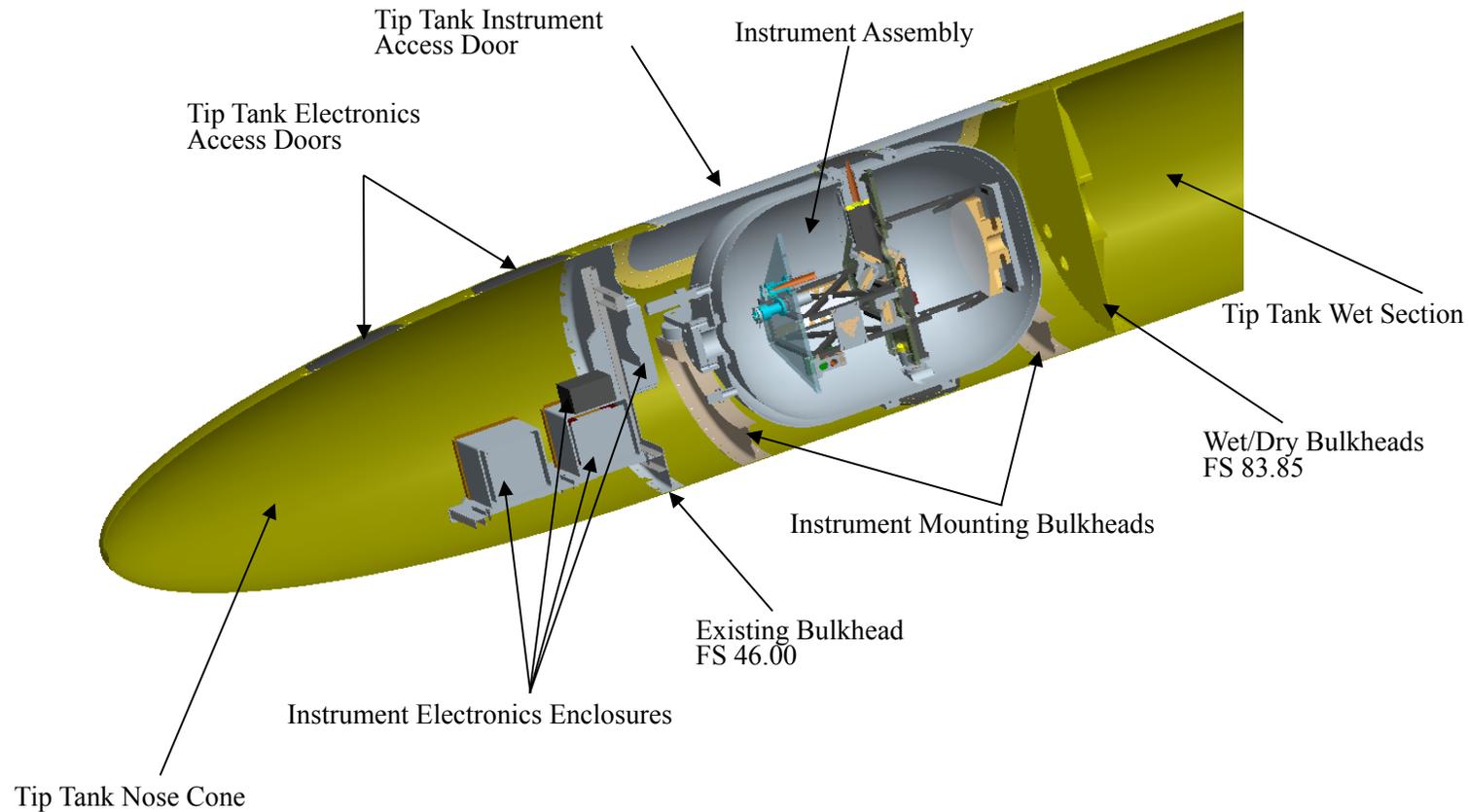
**INFLAME**  
mounted in  
wingtip fuel  
tanks





Langley Research Center

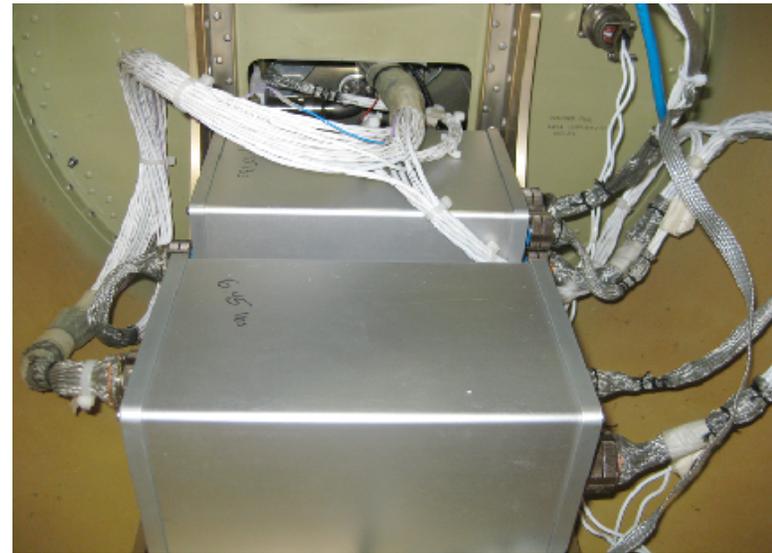
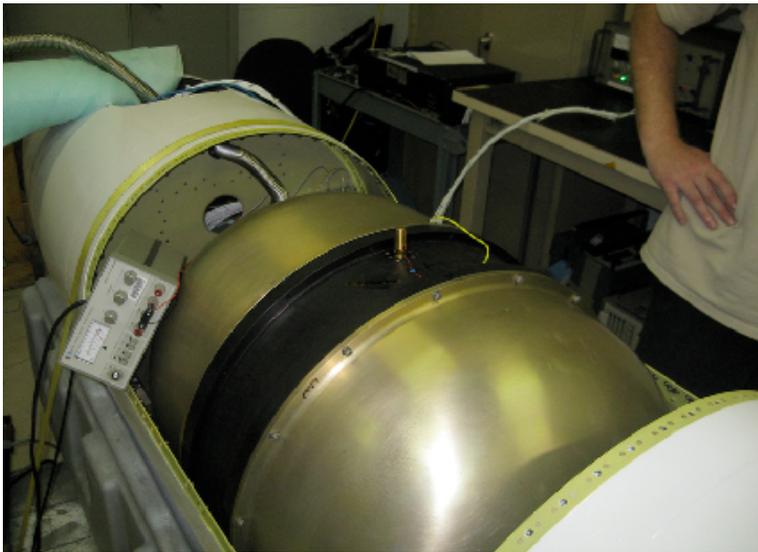
# Fuel Tank Integration





Langley Research Center

# LW Compatibility: 7/29/2009





Langley Research Center

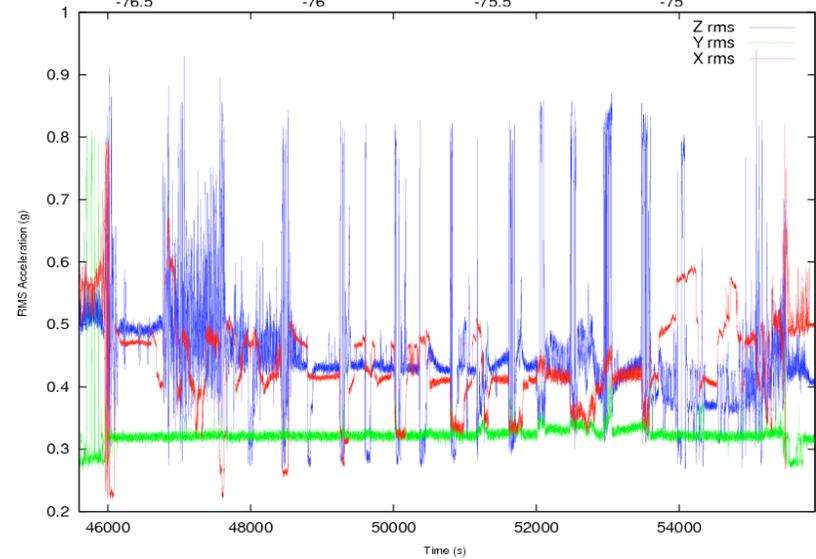
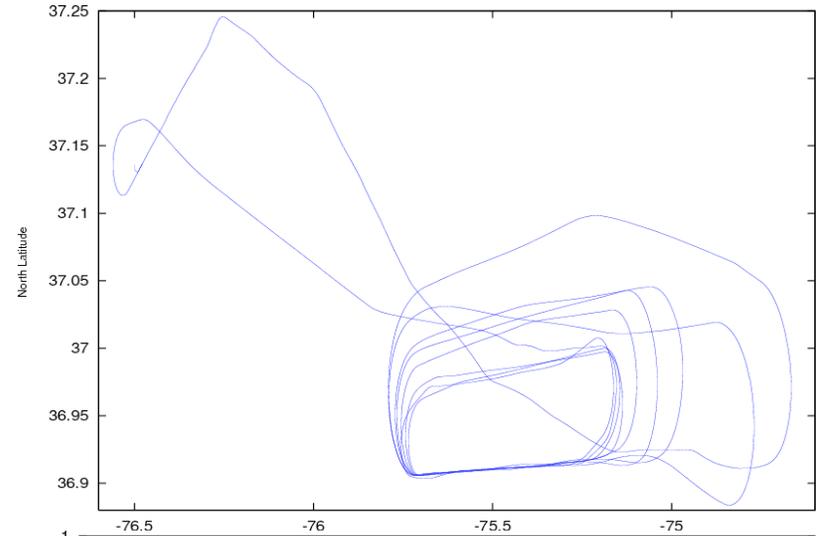
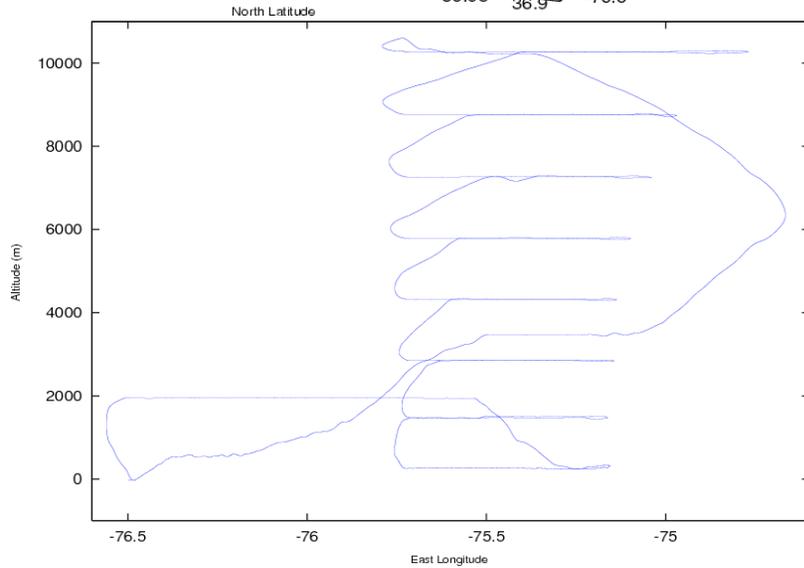
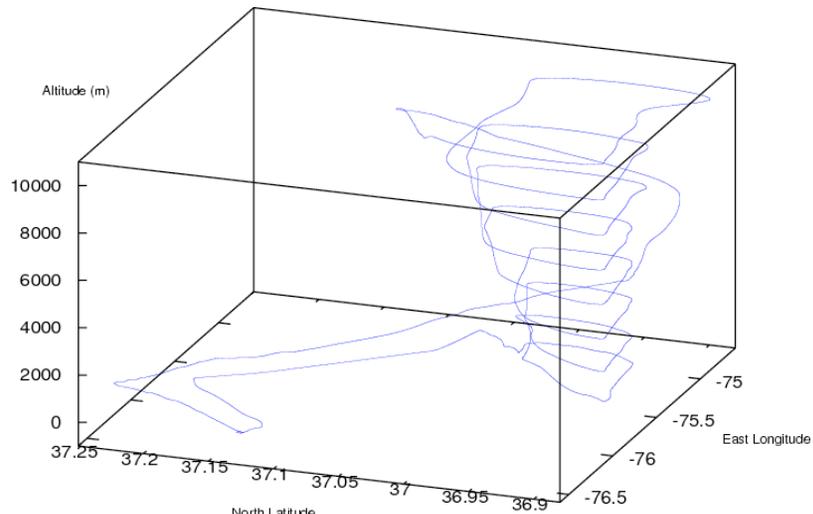
# LW Flight Install: 1/4/2010





Langley Research Center

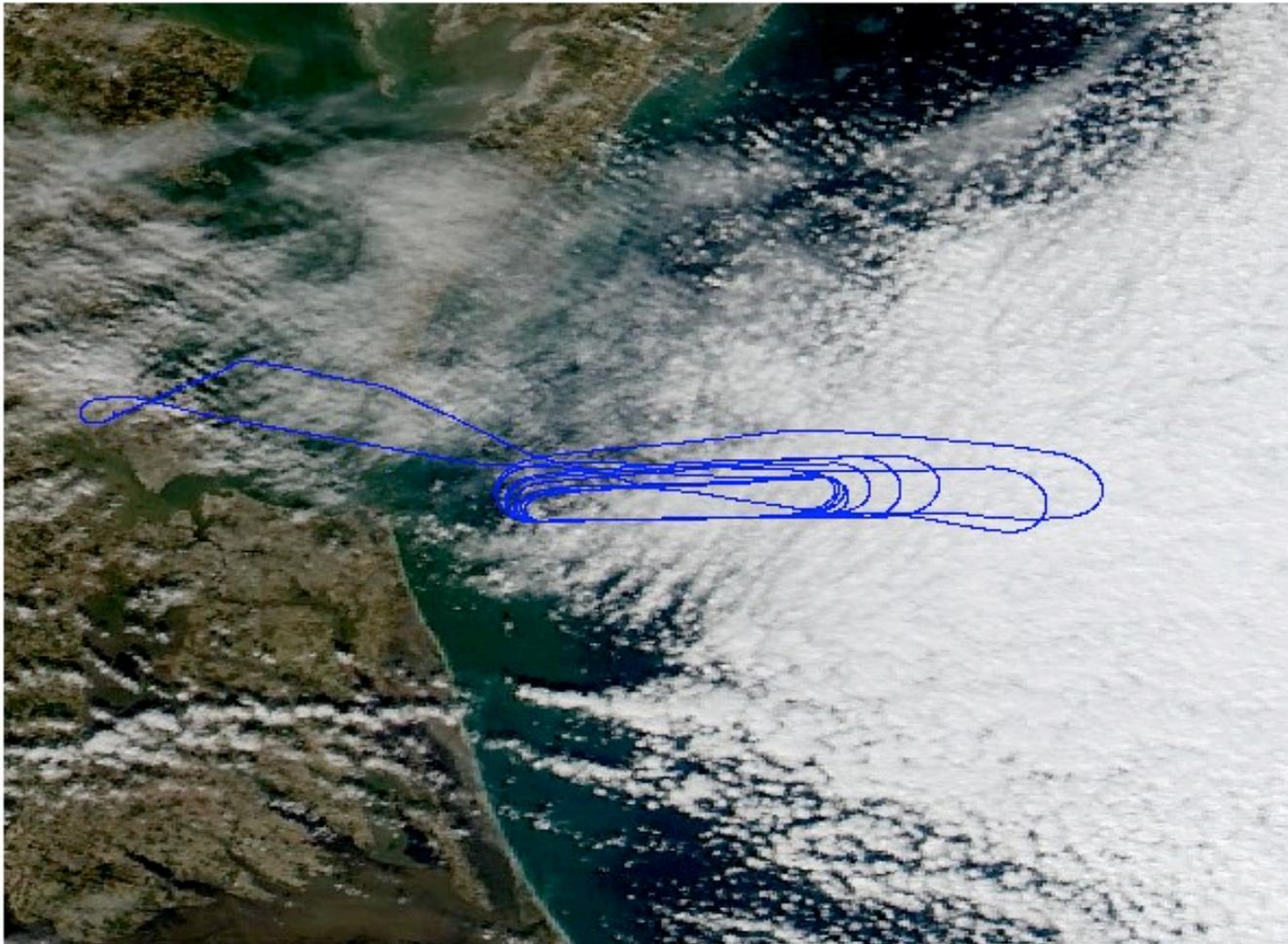
# Flight Track and Accelerometer





Langley Research Center

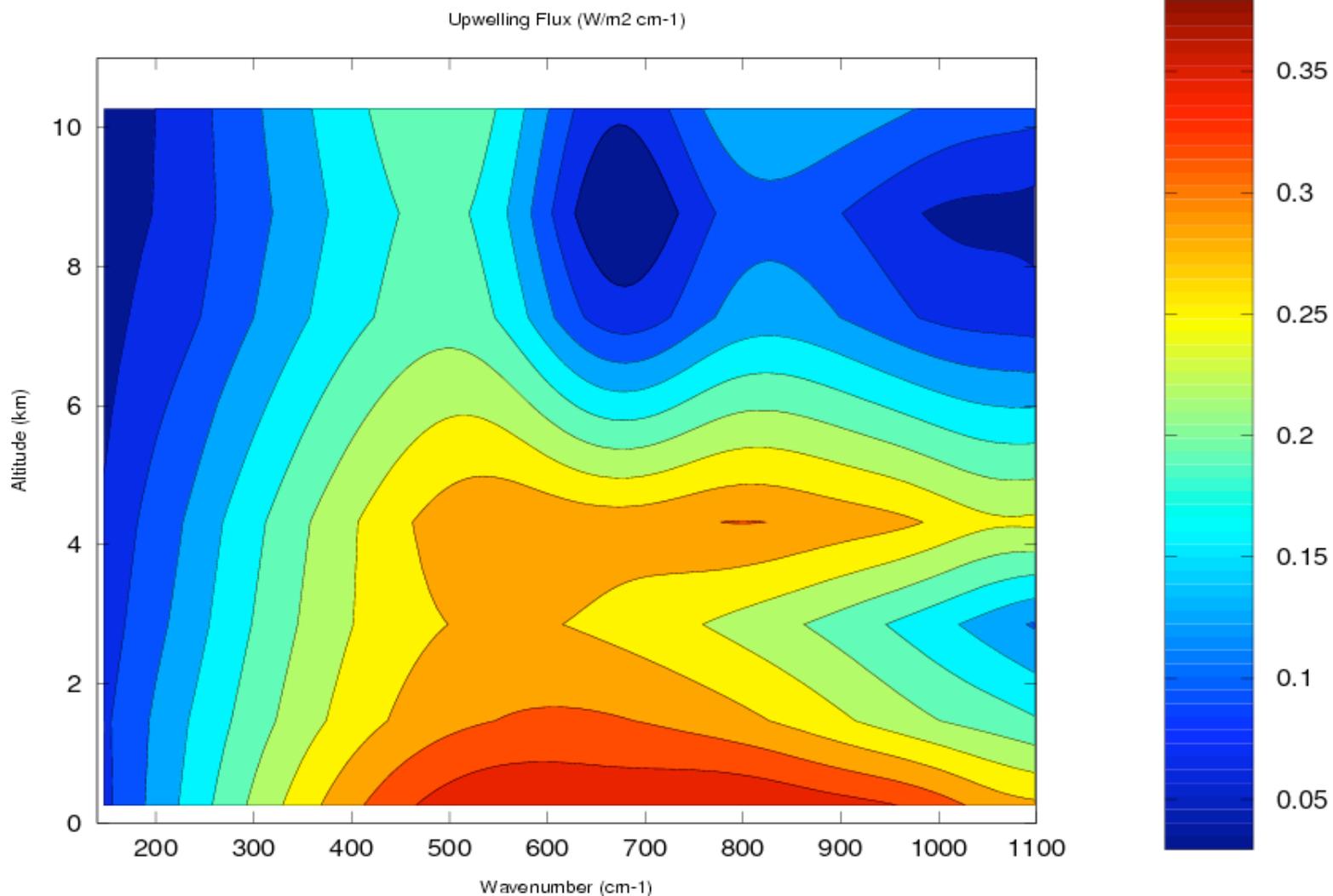
# MODIS Visible Imagery





Langley Research Center

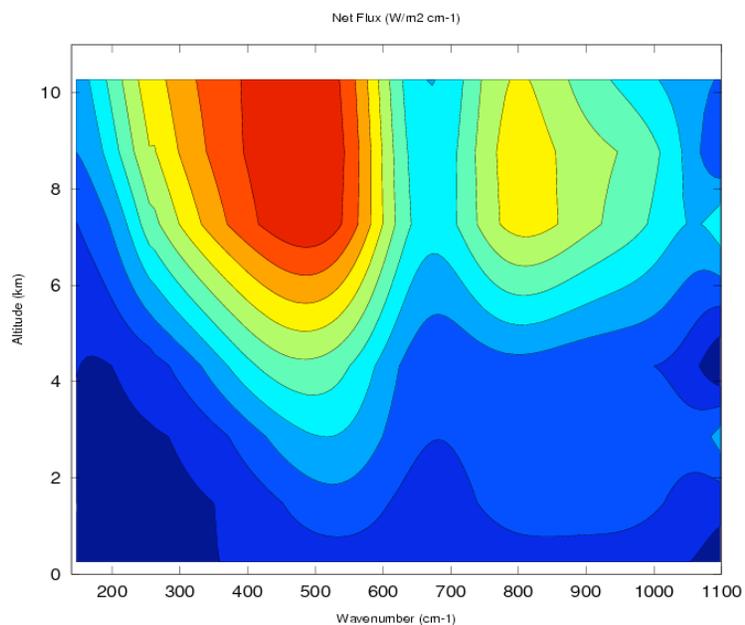
# INFLAME Data Product





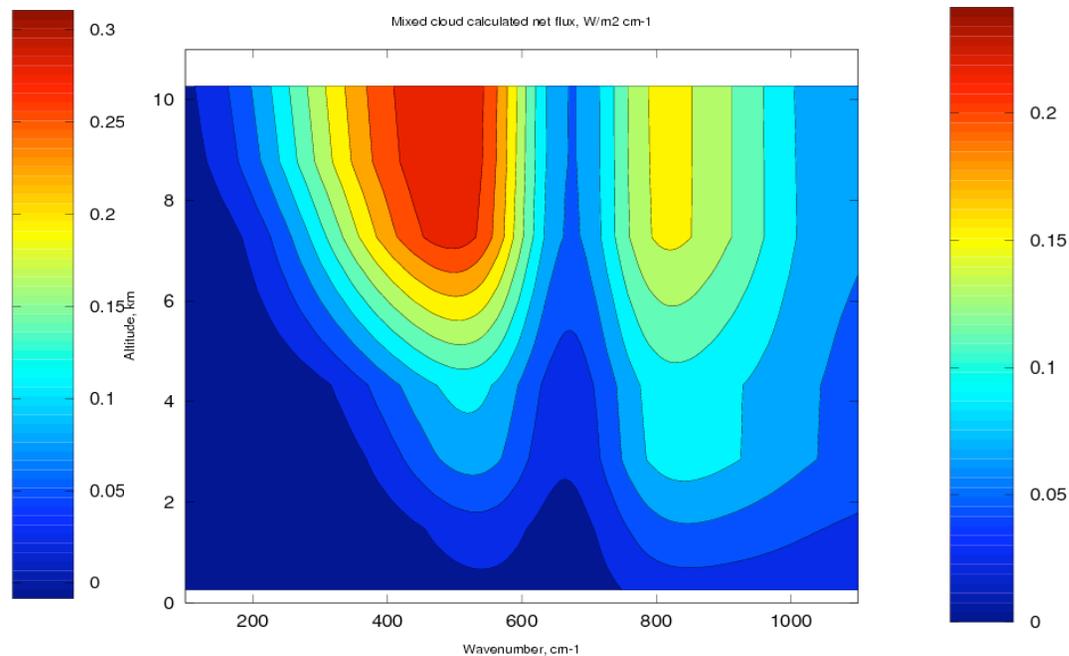
Langley Research Center

# INFLAME LW Net Flux



**Measured LW Net Flux**

**1/5/2010**



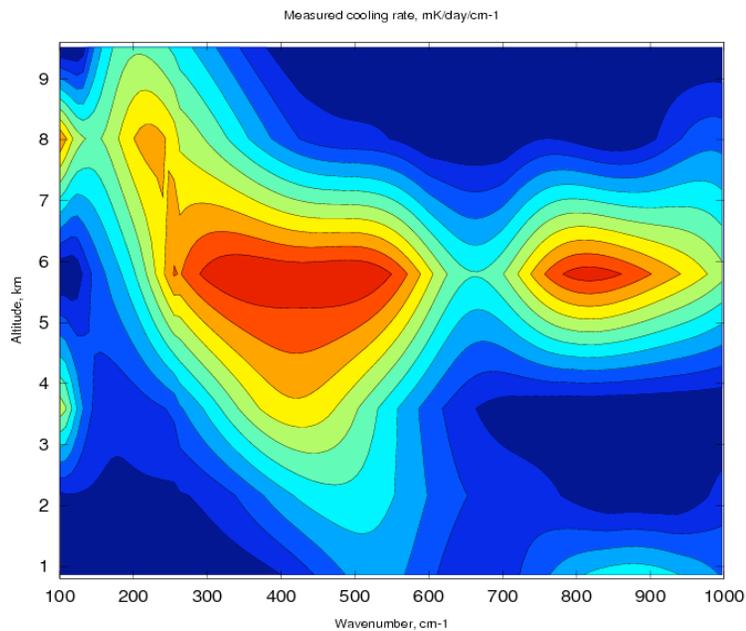
**Calculated LW Net Flux**

**1/5/2010**



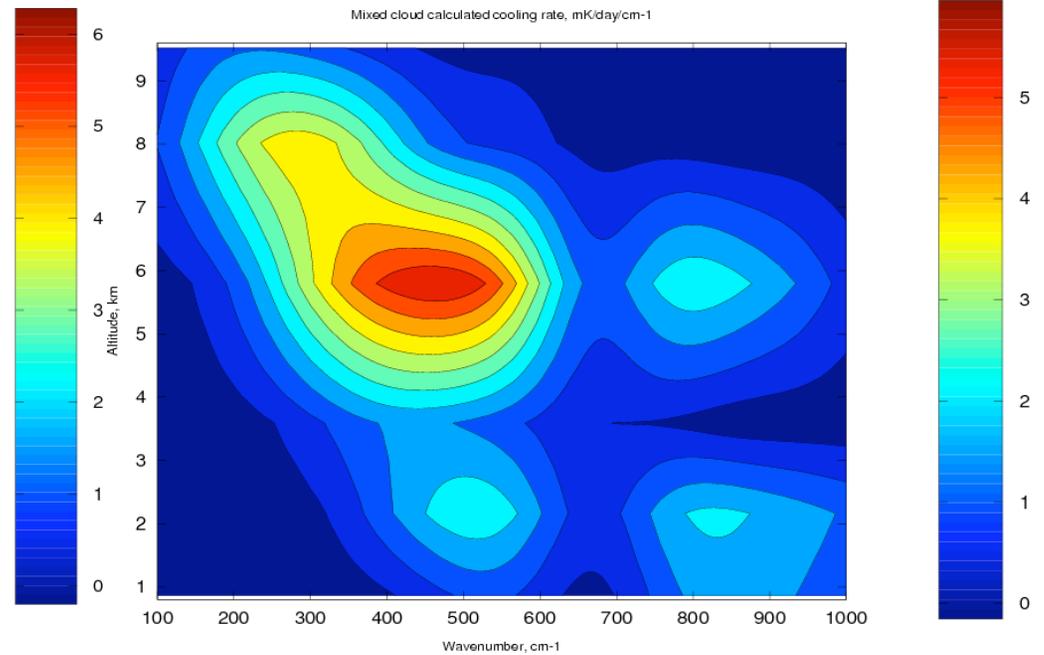
Langley Research Center

# INFLAME Derived LW Cooling Rates



Measured LW Cooling mK/Day/cm<sup>-1</sup>

1/5/2010 flight.



Calculated LW Cooling mK/Day/cm<sup>-1</sup>

1/5/2010 flight



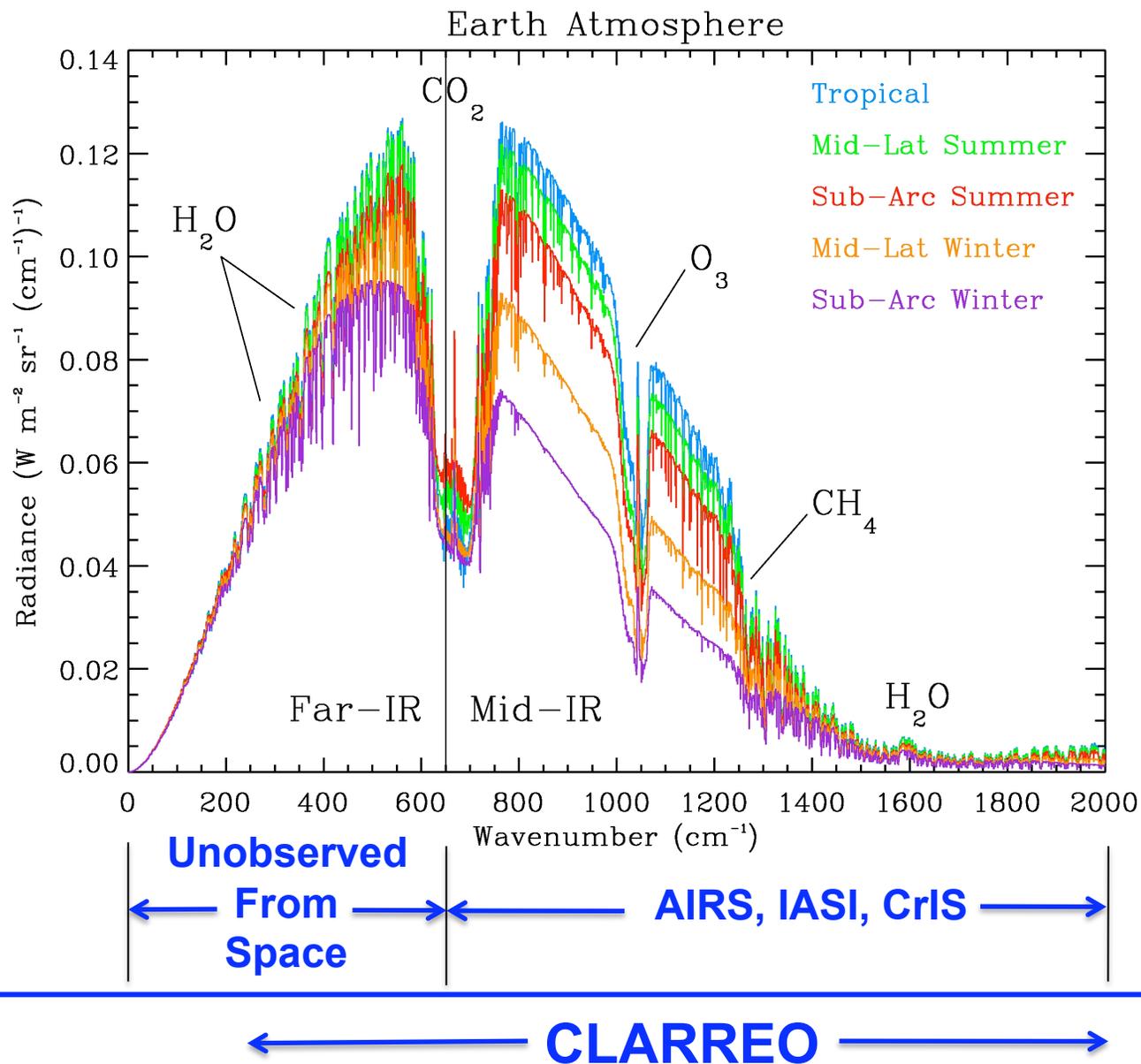
Langley Research Center

# INFLAME Flight Results

## Summary

- **IR FTS (LW) Successfully demonstrated during flight**
  - **First direct measurement of net fluxes and cooling rates within the atmosphere**
  - **Analysis is ongoing**
- **UV-NIR FTS (SW): No useful flight spectra obtained.**
  - **Commercial controller failed due to excessive drift with temperature before takeoff.**

# Earth's Outgoing Longwave Radiation



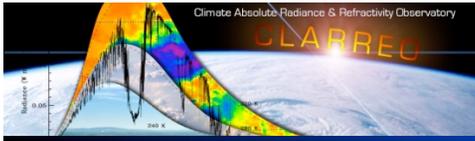
---

# CORSAIR – Langley's Current IIP

---

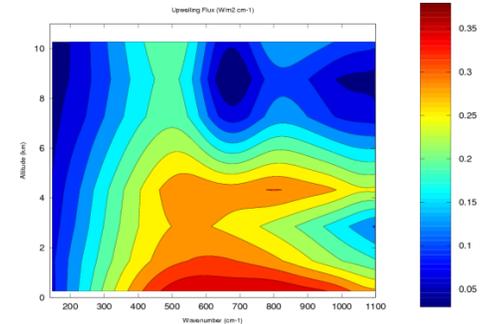
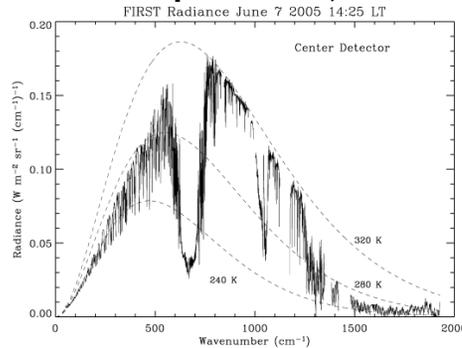
- **FIRST, INFLAME demonstrate numerous technologies for FTS**
- **Some CLARREO – specific technologies remain**
- **CORSAIR IIP Project addresses:**
  - **Broad bandpass beamsplitters: ITT**
  - **Warm high sensitivity detectors: RVS**
  - **Blackbody radiance sources: SDL**
- **CORSAIR presently in year 2 of 3**





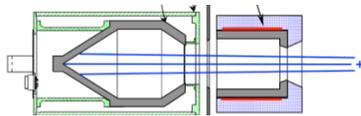
# Technology Development Heritage

- LaRC has been developing technology for CLARREO for nearly one decade
  - FIRST and INFLAME FTS instruments developed and flown
  - CORSAIR and FIREBIB projects in progress
    - Detectors, beamsplitters, blackbodies in development

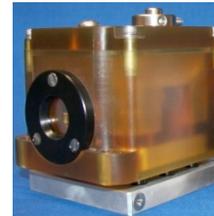


- UW/Harvard also developing CLARREO Technology
  - Verification system elements in development

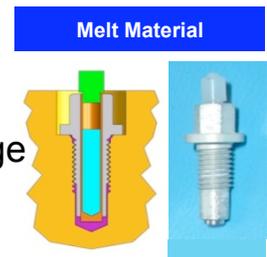
Blackbodies  
Emissivity Monitoring



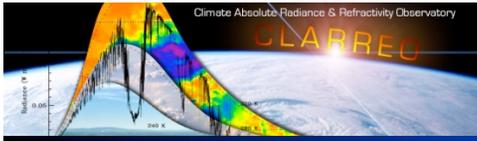
QCL



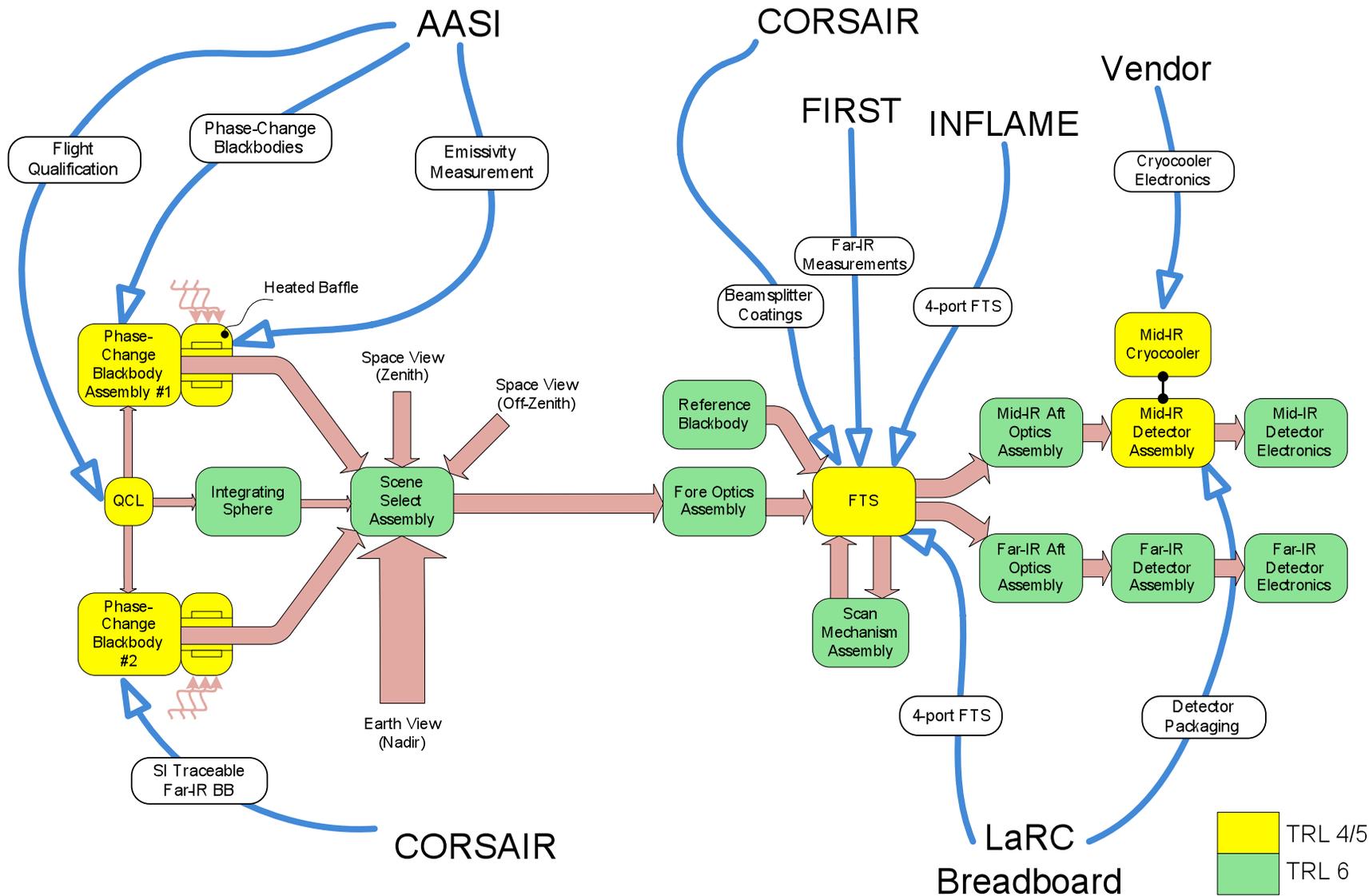
Phase Change  
Cells



**Key CLARREO technologies in development to TRL 6**



# Technology Development Plan Leverages Hardware Matured Through Breadboard and IIP Programs



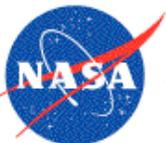
IIPs and breadboard provide early risk reduction

---

# Summary – Incubator to Orbit

---

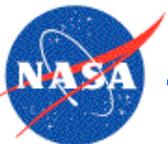
- **FIRST instrument successfully developed and demonstrated and transitioned to operational scientific instrument in FORGE campaign**
- 
- **INFLAME instrument successfully developed and demonstrated – awaiting opportunity for science campaign**
- **CORSAIR Project developing and demonstrating component technologies specific to CLARREO Mission**
- **Additional CLARREO-specific technologies being developed by U. Wisconsin and Harvard**
- **These ESTO-IIP efforts are paving the way to CLARREO via science and instrument technology**



---

---

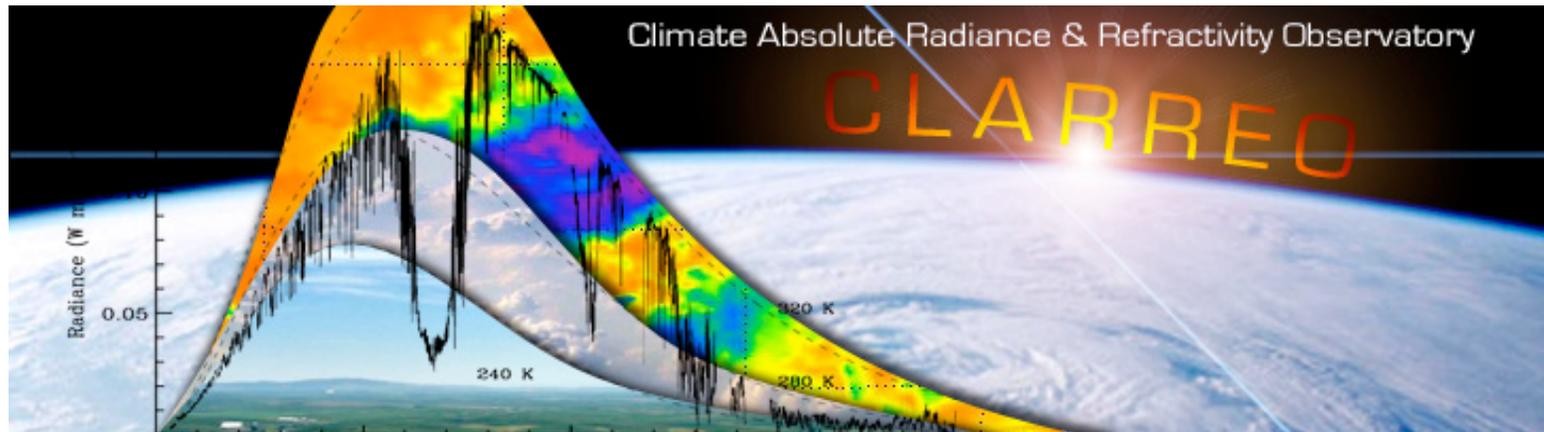
# Backup Slides



---

# Onto the Future – The CLARREO Mission

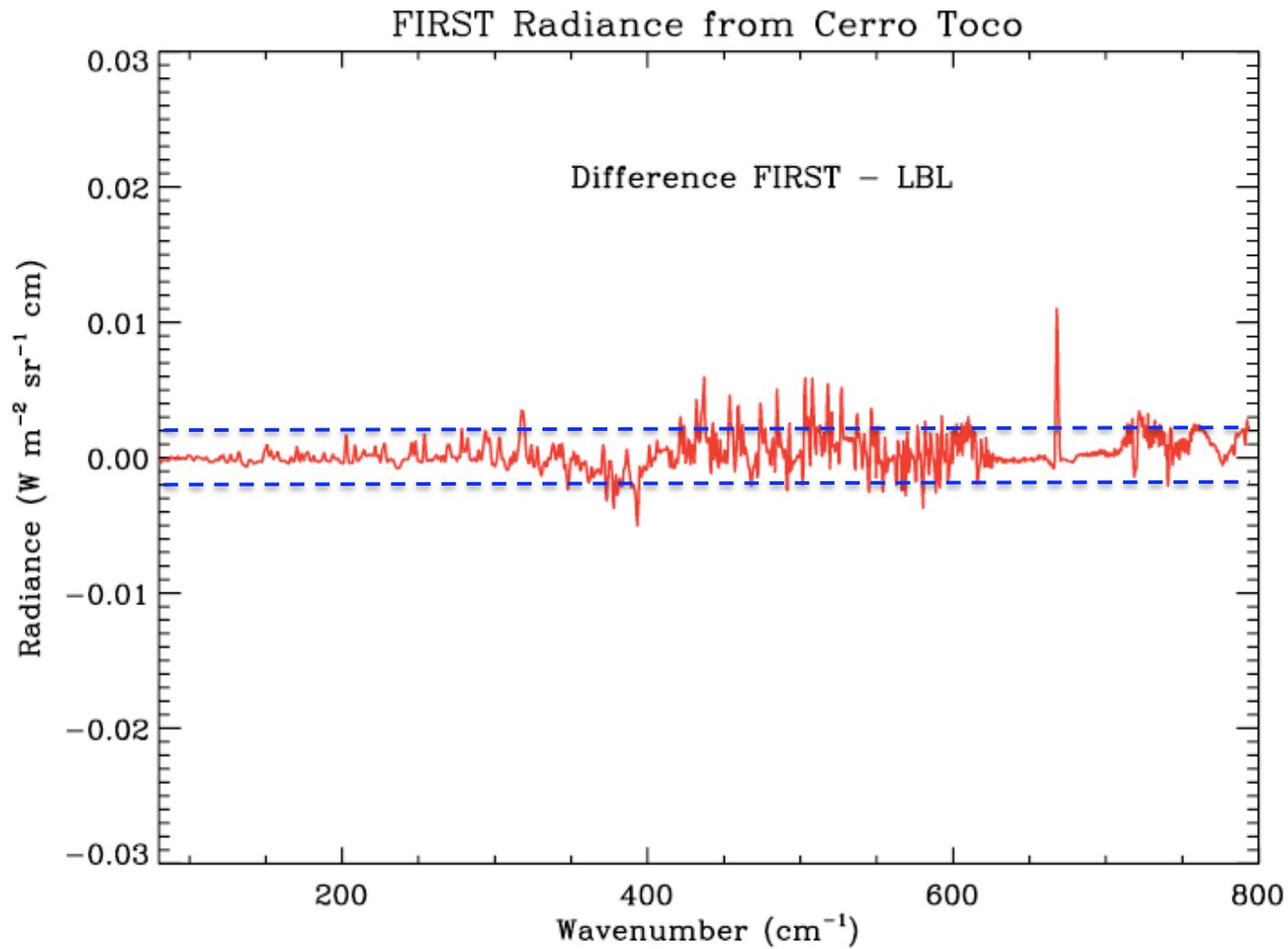
---



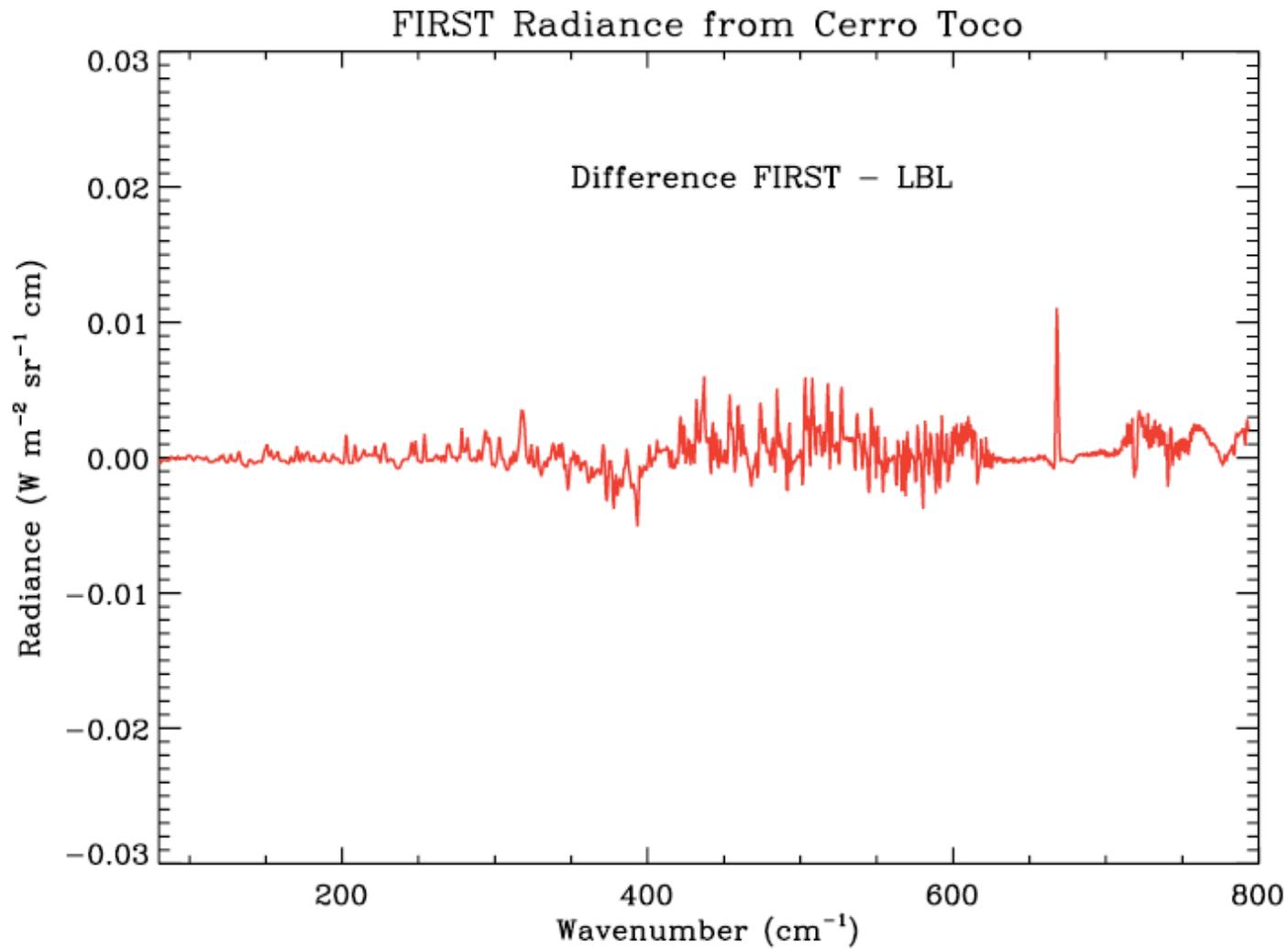
- **FIRST, INFLAME demonstrate FTS instrument technology**
- **CORSAIR is developing CLARREO-specific component technology**
- **Advancing calibration and climate system measurement knowledge**
- **Enabling development of CLARREO IR FTS instrument**



# Radiance Difference – 09/05/2009



# Radiance Difference – 09/05/2009





# Instrument Overview

- We use two instruments to cover the required spectral range:
  - LW instrument covers thermal IR, 100  $\mu\text{m}$  to 3  $\mu\text{m}$ .
  - SW instrument covers solar, 3  $\mu\text{m}$  to 0.3  $\mu\text{m}$ .
- Main differences between LW and SW are the calibration sources, optical coatings, and detectors.

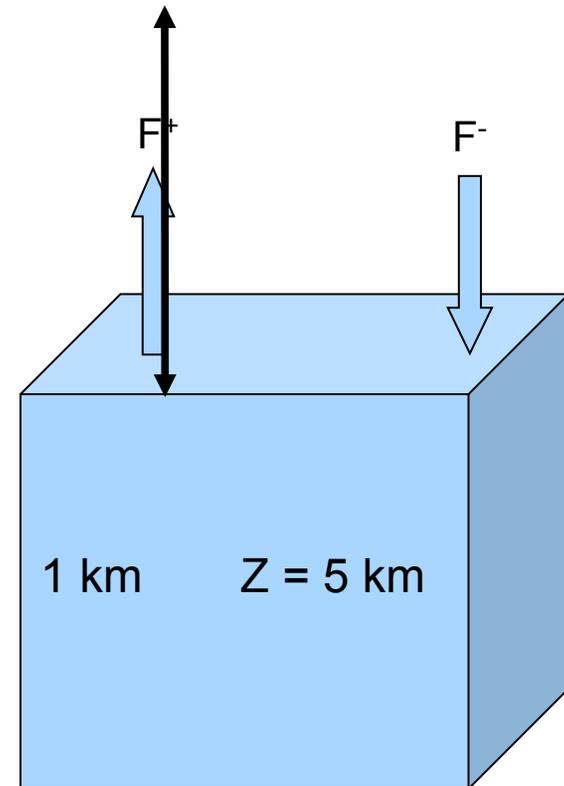


# Measurement Overview

- **Consider a unit cube in the terrestrial atmosphere:**
  - Spectral flux is the energy per unit frequency interval flowing through one face ( $F^+$  or  $F^-$ );
  - Net Flux is the difference in energy flowing through one face in opposite directions:

$$F_z = F^+ - F^-$$

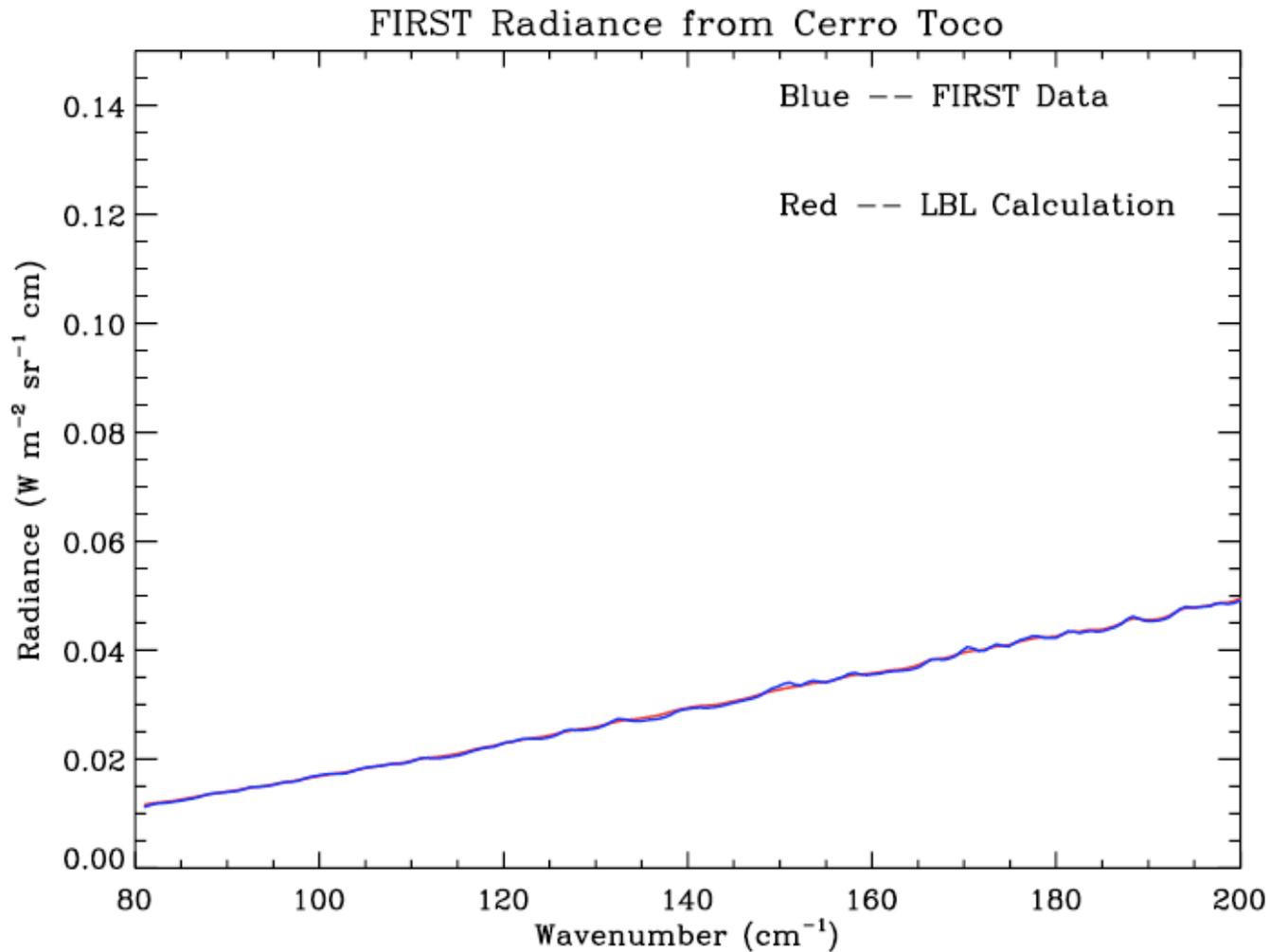
- Flux divergence is given by  $dF_z/dz$ ;
  - Assume  $dF_x/dx$  and  $dF_y/dy$  are small.
- **Radiative heating rate is estimated from the measured flux divergence:**



---

# September 5 2009 – PWV = 0.75 mm

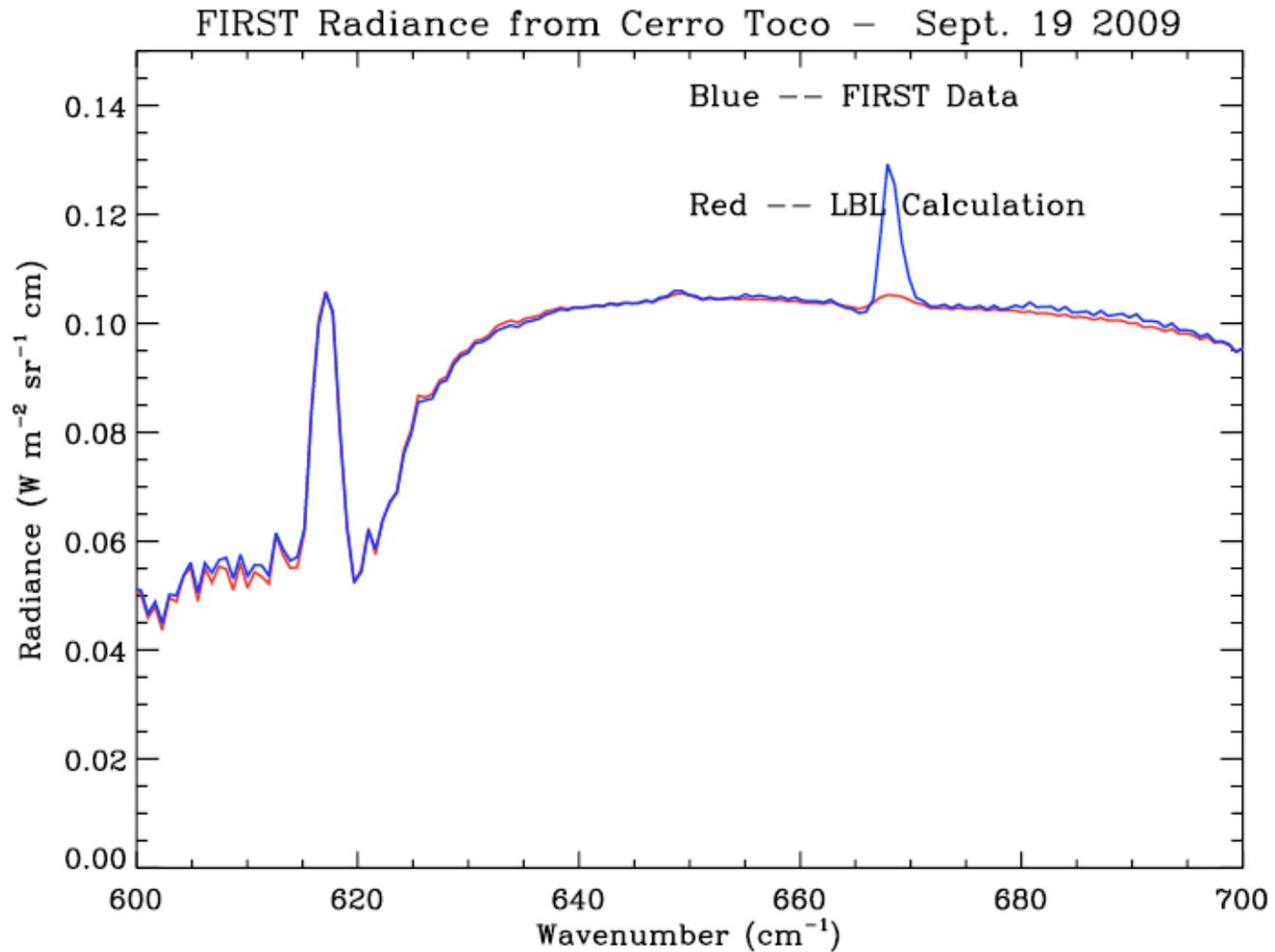
---



---

# September 19 2009 – PWV = 0.4 mm

---





# INFLAME Objective

- **Measure the spectral dependence of radiative heating rates in the troposphere.**
  - **Flight demonstration goal is to measure net flux with sufficient stability to estimate radiative heating rates from the net flux divergence in 1 km layers of the troposphere with an accuracy of 10%**
- **Milestones:**
  - **Proposed: 11/2004**
  - **Funded: 11/2005**
  - **Calibration: 7/2009**
  - **Test flight: 1/5/2010**



# Measurement Challenge

- What if we use uplooking and downlooking instruments to measure fluxes as functions of altitude, subtract to get net flux, and take the derivative to get flux divergence?
  - Need to measure small changes in the difference of large numbers:

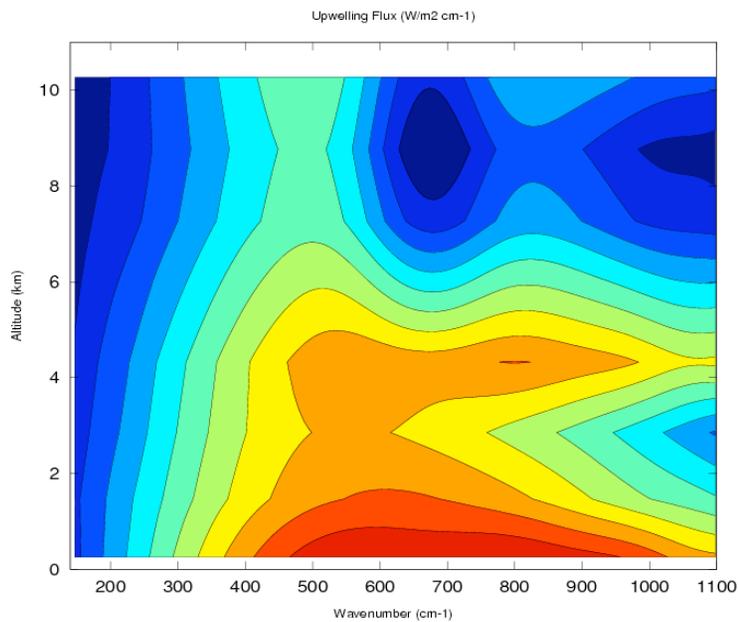
	F+, W/m <sup>2</sup>	F-, W/m <sup>2</sup>	Fz, W/m <sup>2</sup>	dFz/dz, W/m <sup>2</sup> km
Thermal IR	344	263	81	16.1
Solar	80	768	-688	-13.6

- Small systematic errors in measured F+ and F- can easily be as large as  $dF_{\text{net}}(z)/dz$ .

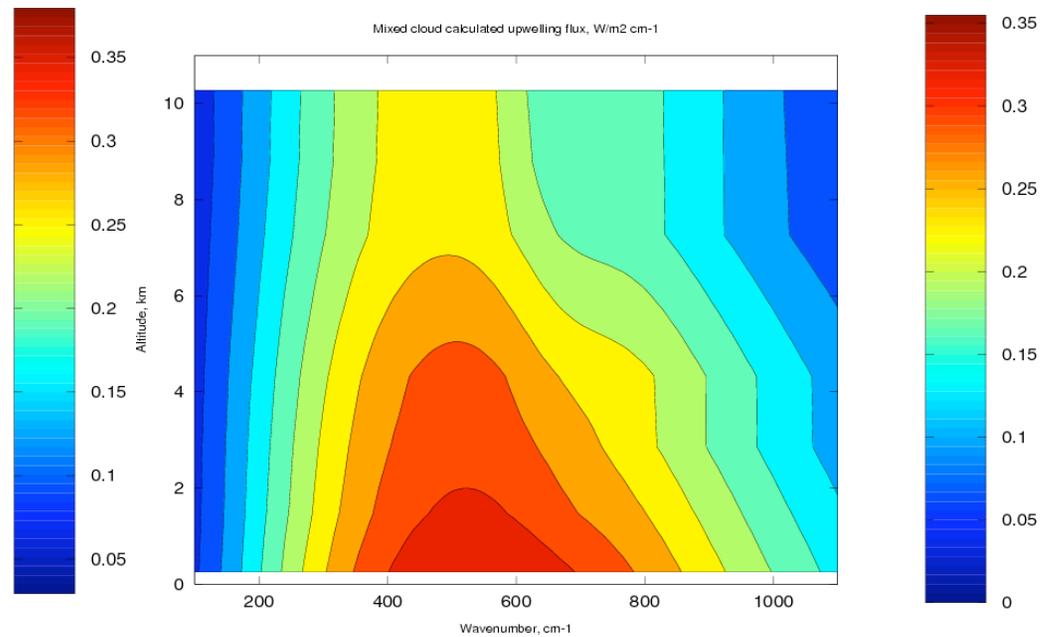


Langley Research Center

# INFLAME LW Upwelling Flux



Measured Upwelling IR Flux



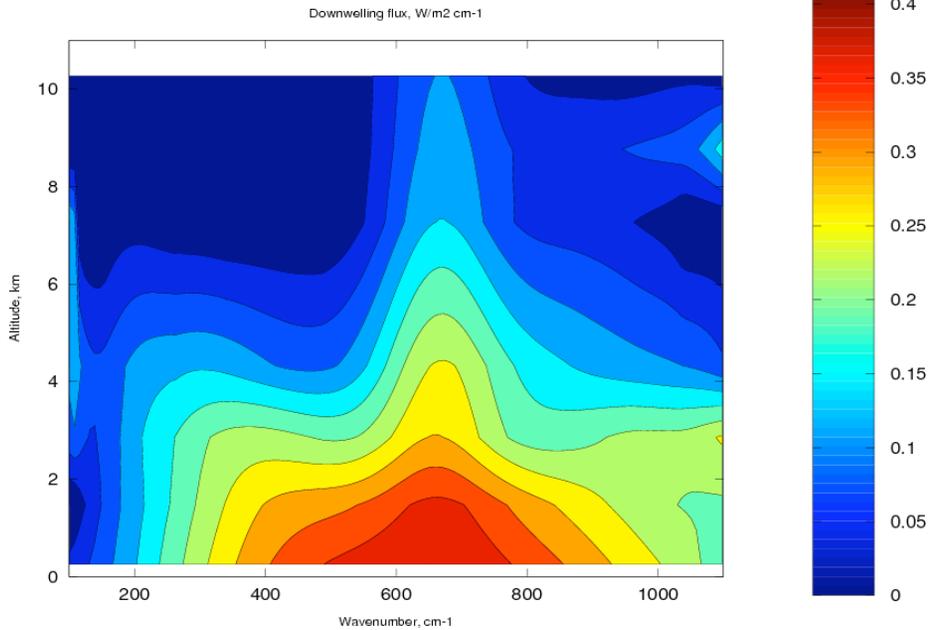
Calculated Upwelling IR Flux

1/5/2010

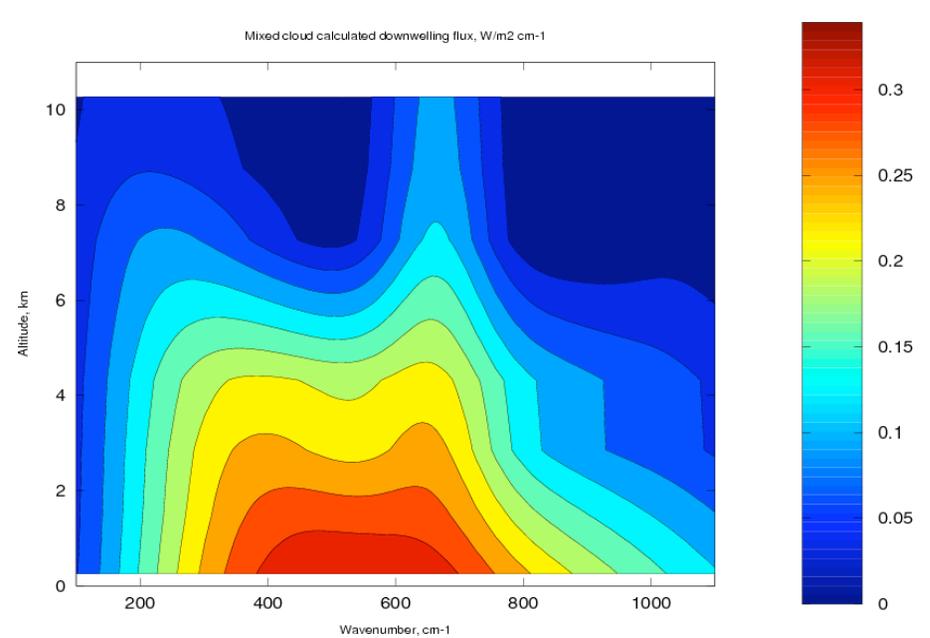


Langley Research Center

# INFLAME LW Downwelling Flux



**Measured LW Downwelling Flux**



**Calculated LW Downwelling Flux**

1/5/2010