

Imaging Spectrometer Science Measurements for Terrestrial Ecology: AVIRIS and the Next Generation AVIRIS Characteristics and Development Status

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OVERVIEW



- Objective
- Spectroscopy or multi-spectral
- Signals
- Imaging Spectroscopy
- Example of Imaging Spectroscopy based Science
- AVIRIS classic measurement characteristics
- Next Generation AVIRIS science measurement characteristics
- Next Generation Design and Status
- Summary



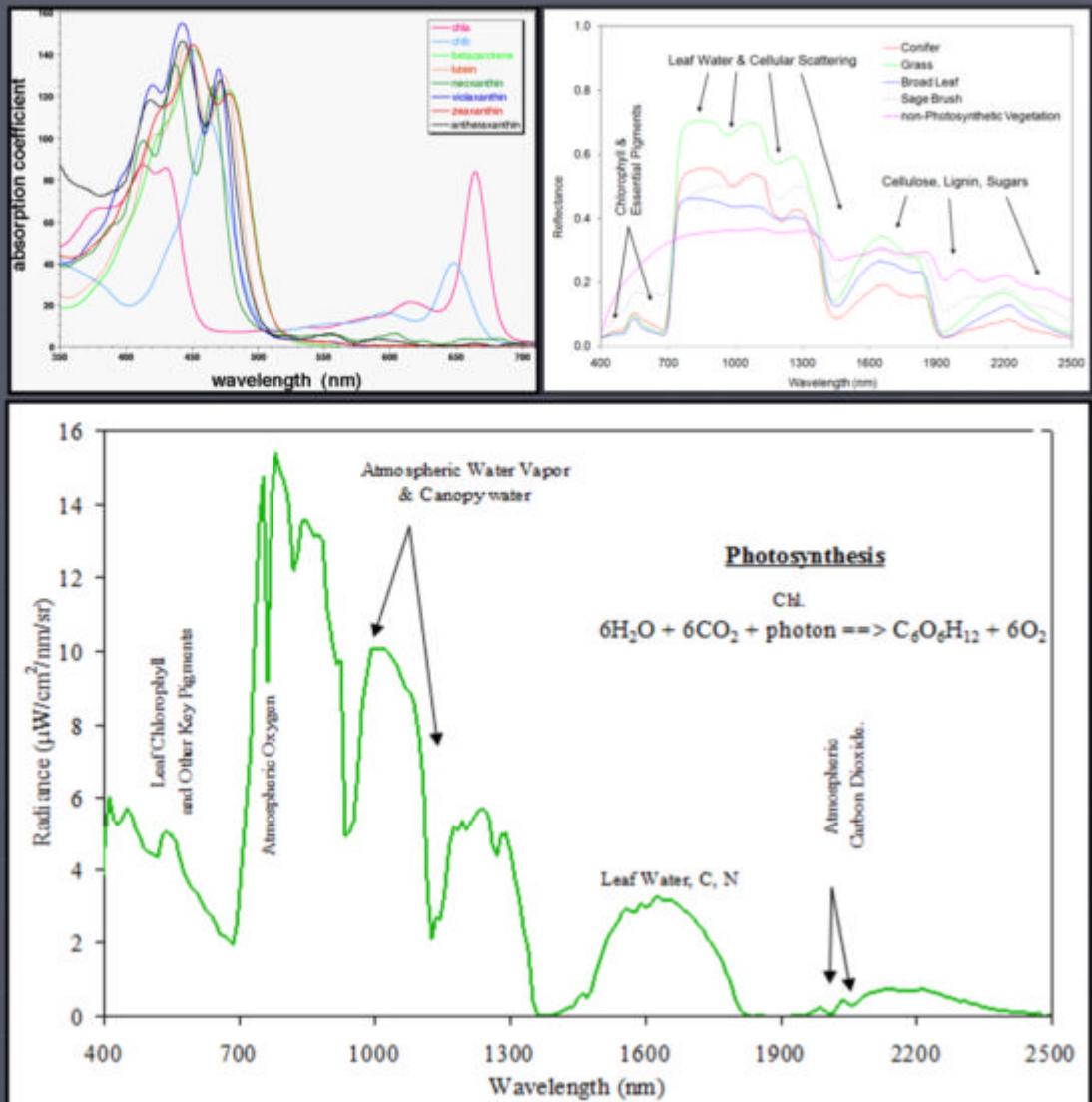
OBJECTIVE



- Answer next generation science questions with calibrated high uniformity and high signal-to-noise ratio imaging spectroscopy measurements

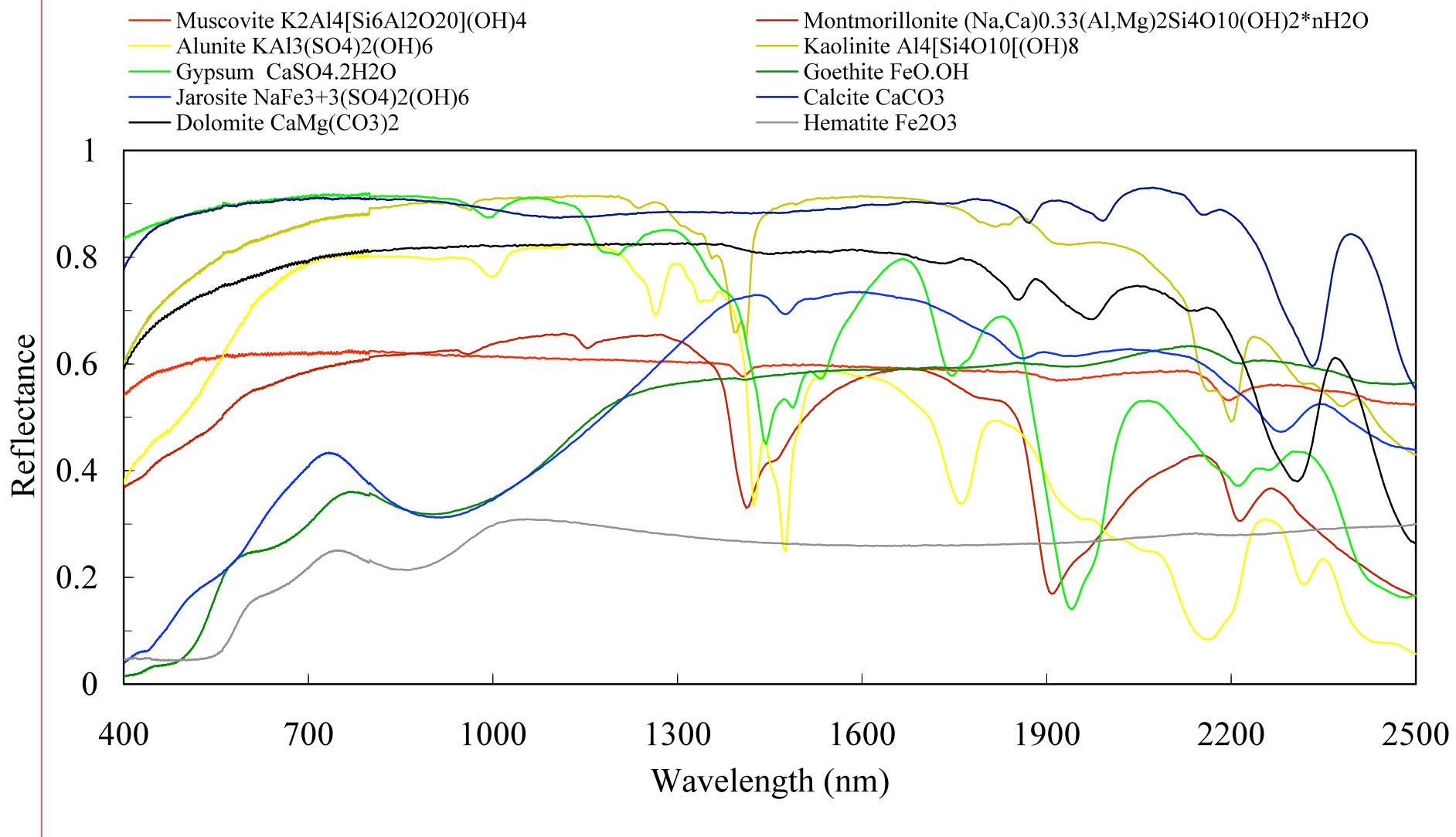


SIGNALS VEGETATION





SIGNALS MINERALS

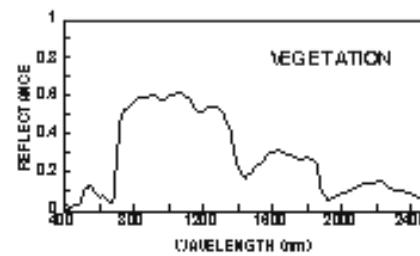
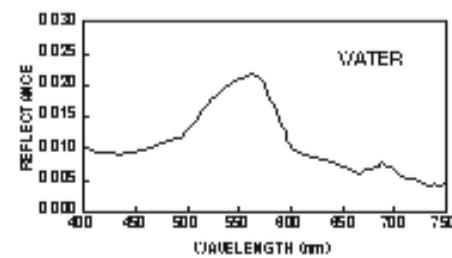
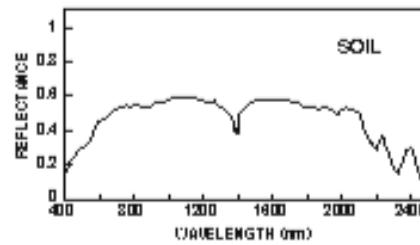
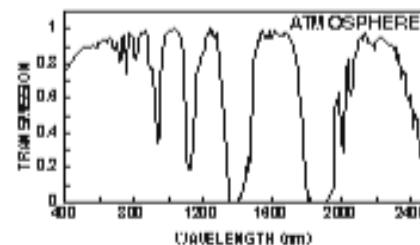
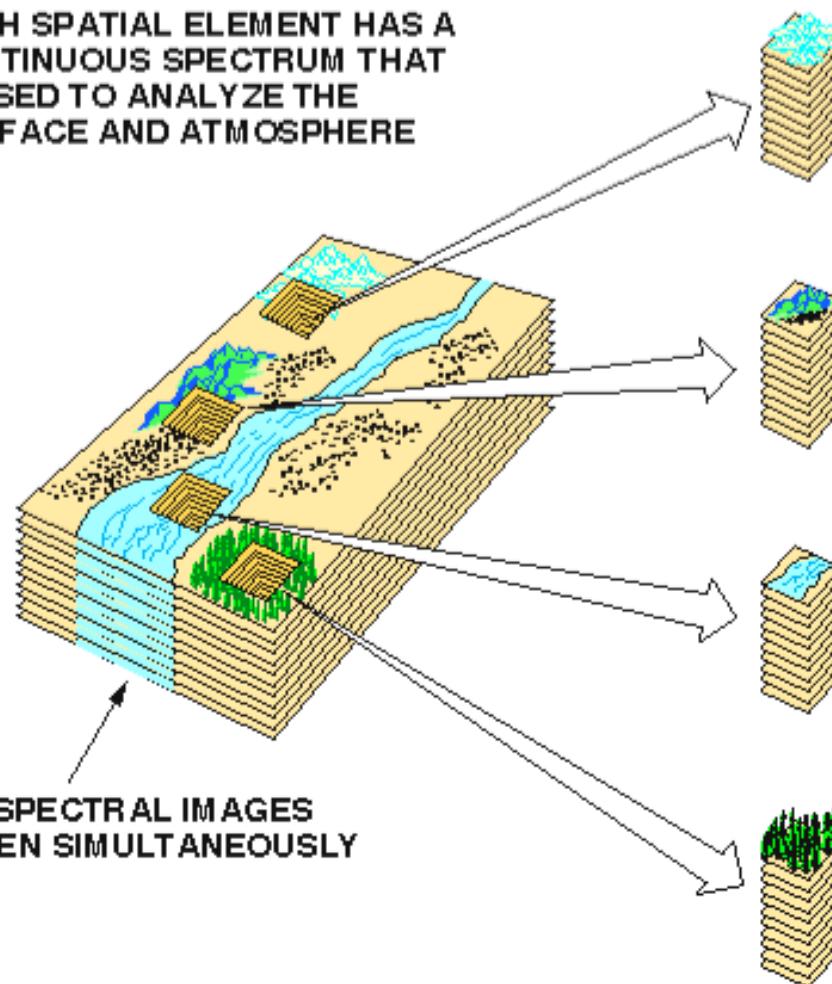




Imaging Spectroscopy Concept

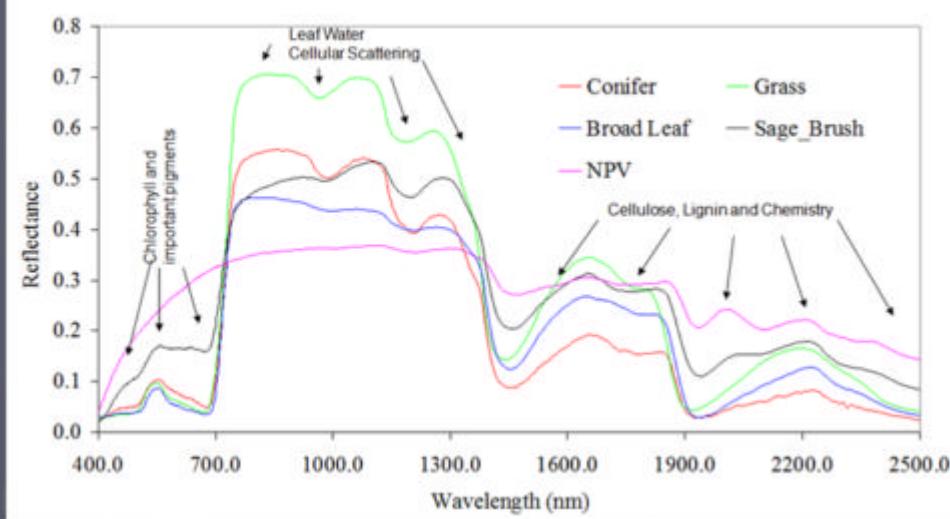


EACH SPATIAL ELEMENT HAS A
CONTINUOUS SPECTRUM THAT
IS USED TO ANALYZE THE
SURFACE AND ATMOSPHERE



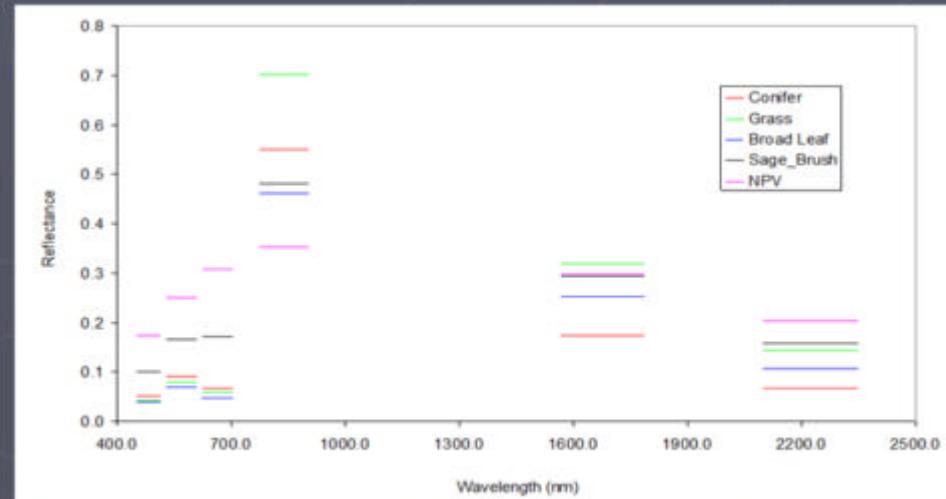


SPECTROSCOPY OR MULTISPECTRAL



Spectroscopy is required

Multi-Spectral is Insufficient





JPL EXPERIENCE



- AIS Retired in AVIRIS lab
- AVIRIS Now flying in NASA ER-2
- NIMS, VIMS VIMS orbiting Saturn
- Hydice Retired
- Hyperion Orbiting Earth
- [CRISM] Orbiting Mars
- MaRS Flying for DOD customer on various aircraft
- MMM (M3) Moon flight mission complete
- ARTEMIS Orbiting the Earth
- PBTB and ISTB Ongoing laboratory test bed activities

[] = minimal JPL involvement, but important lessons learned.

RED = sensor no longer operating.



AVIRIS "CLASSIC" SCIENCE MEASUREMENTS



Spectral

Range	380 to 2500 nm
Sampling	10 nm
Response	10 nm

Radiometric

Range	0 to Max Lambertian
Sampling	14 bit
Calibration	+/- 1nm
Signal-to-Noise ratio	>1000 @ 600 nm >400 @ 2200 nm

Spatial

Field-of-View	34 degrees
Instantaneous FOV	1 milliradian
Spatial swath	2.5 to 11 km @ alt (4 to 20 km)
Spatial resolution	4 to 20 m @ alt (4 to 20 km)

Uniformity

Spectral-Cross-Track	>95% Uniform
Spectral-IFOV	>95% Uniform

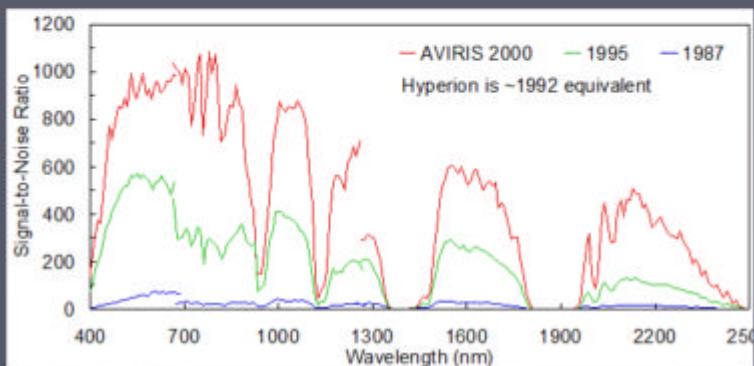


AVIRIS-NG KEY SCIENCE REQUIREMENTS

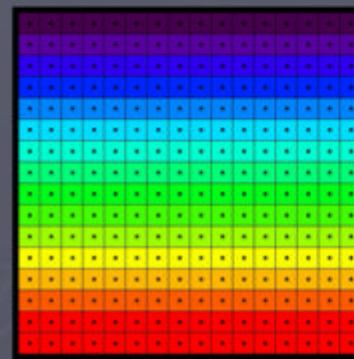


What is needed in the instrument/measurement:

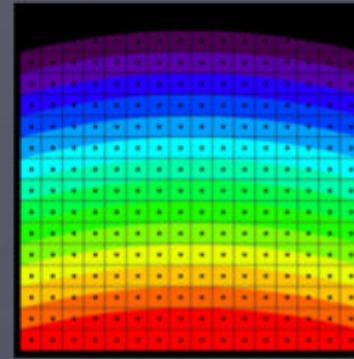
- *High Signal-to-noise ratio required for molecular spectroscopy*



- *Uniformity is required for spectroscopy in the image domain*

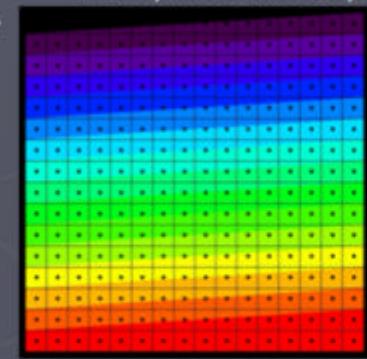
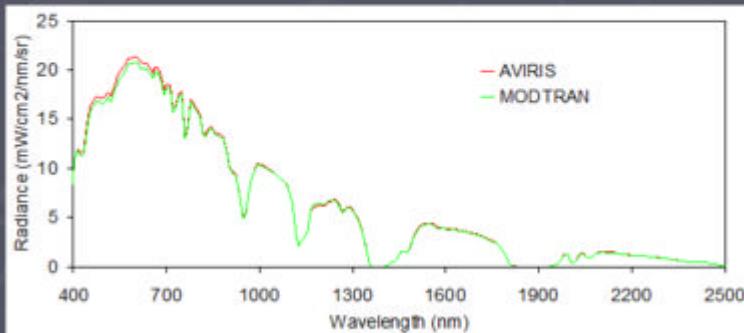


✓ Required Uniformity

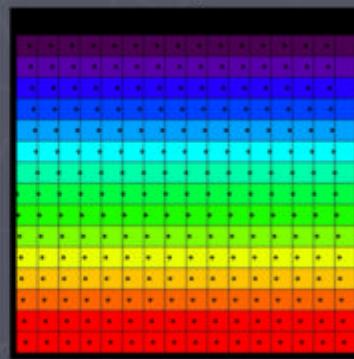


✗ Failure by "frown"

- *Excellent calibration for quantitative results (spectral, radiometric, spatial)*



✗ Failure by twist



✗ Failure by Spectral-IFOV-shift



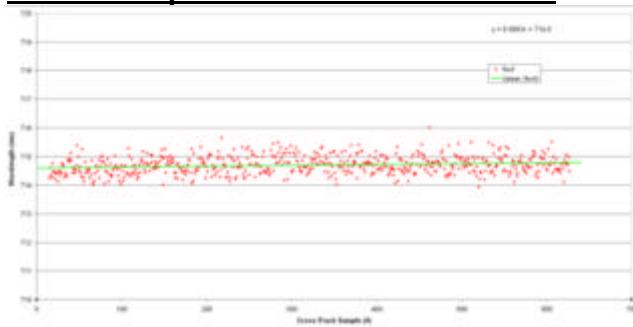
AVIRIS-NG KEY SCIENCE MEASUREMENTS



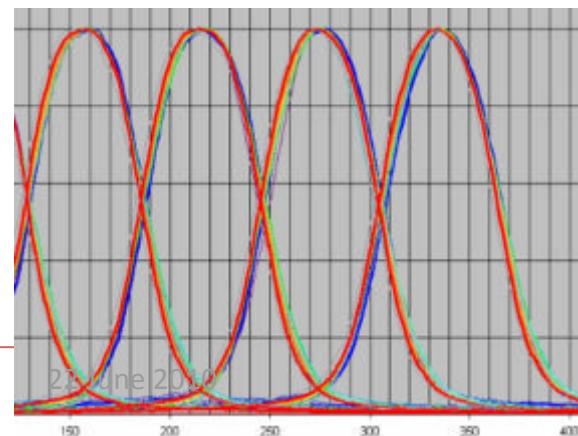
Depiction

- Grids are the detectors
- Spots are the IFOV centers
- Colors are the wavelengths

>95% spectral cross-track



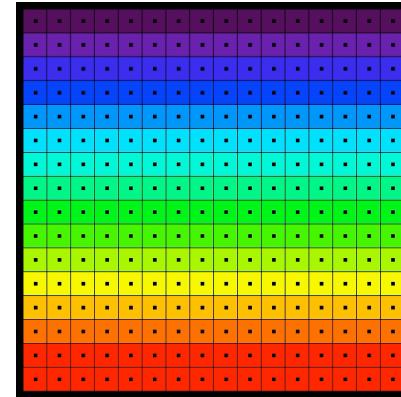
>95% spectral IFOV



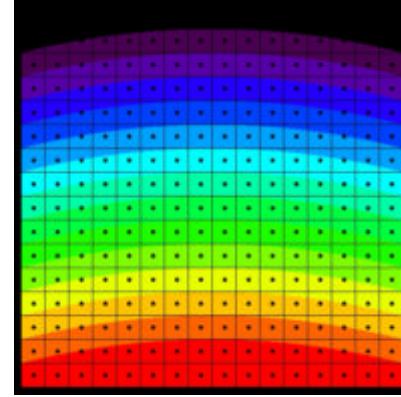
M3

Cross-Track

Wavelength (nm)

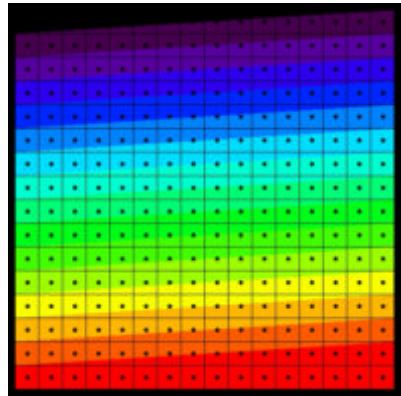


✓ Required Uniformity

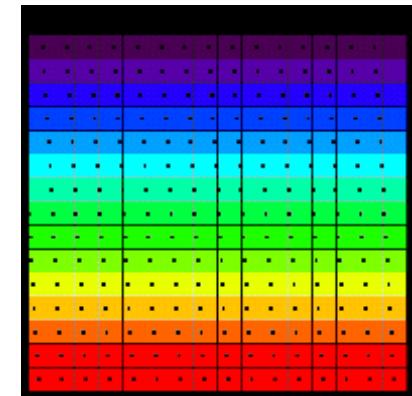


✗ Failure by "frown"

M3



✗ Failure by twist



✗ Failure by Spectral-IFOV-shift



RESEARCH AND APPLICATIONS



- Atmosphere: water vapor, clouds properties, aerosols, absorbing gases ...
- Ecology: chlorophyll, leaf water, lignin, cellulose, pigments, structure, nonphotosynthetic constituents ...
- Geology and soils: mineralogy, soil type ...
- Coastal and Inland waters: chlorophyll, plankton, dissolved organics, sediments, bottom composition, bathymetry ...
- Snow and Ice Hydrology: snow cover fraction, grainsize, impurities, melting ...
- Biomass Burning: subpixel temperatures and extent, smoke, combustion products ...
- Environmental hazards: contaminants directly and indirectly, geological substrate ...
- Calibration: aircraft and satellite sensors, sensor simulation, standard validation ...
- Modeling: radiative transfer model validation and constraint ...
- Commercial: mineral exploration, agriculture and forest status ...
- Algorithms: autonomous atmospheric correction, advance spectra derivation ...
- Other: human infrastructure ...



AVIRIS “CLASSIC” 100S OF FLIGHTS OVER 20 YEARS



22 June 2010

NASA Earth Science Technology Forum

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22 June 2010

NASA Earth Science Technology Forum

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22 June 2010

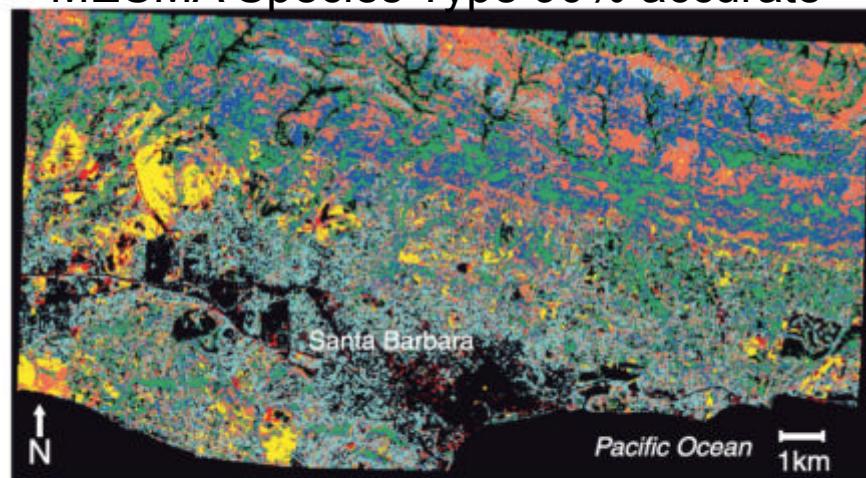
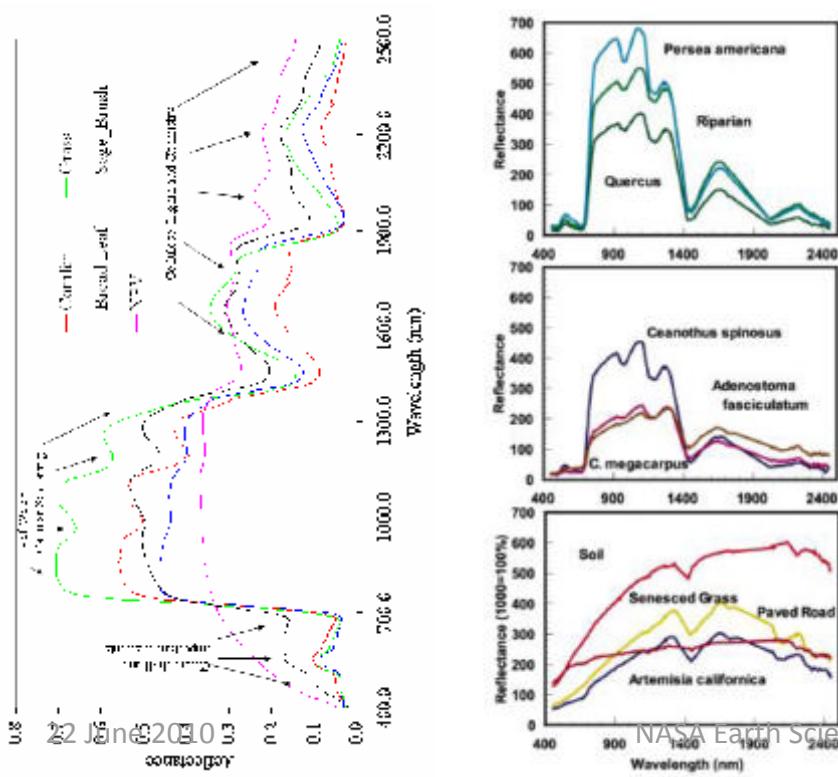
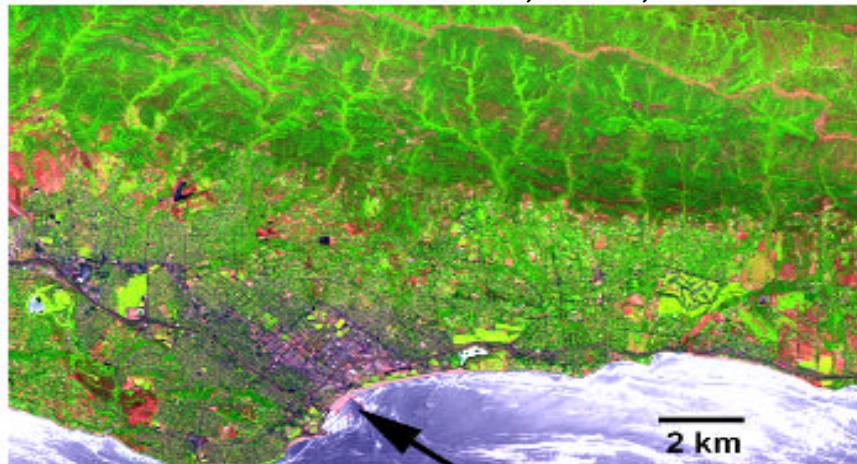
NASA Earth Science Technology Forum
a thoughtful, dedicated JPL team

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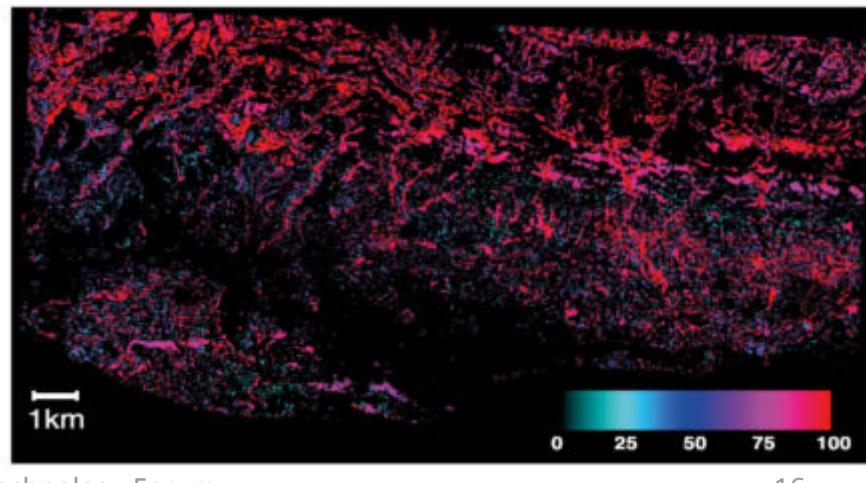
VEGETATION FUNCTIONAL TYPE ANALYSIS, SANTA BARBARA, CA

Dar Roberts, et al, UCSB

MESMA Species Type 90% accurate

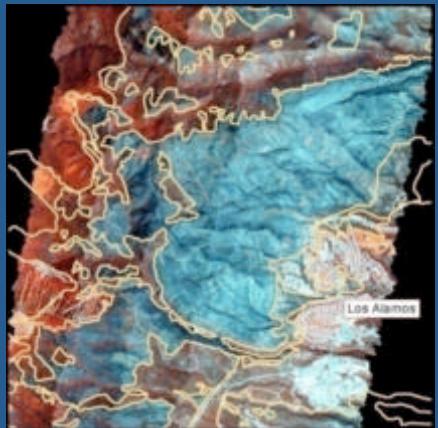


Species Fractional Cover



CERRO GRANDE FIRE SEVERITY, LOS ALAMOS, NM, RAY KOKALY

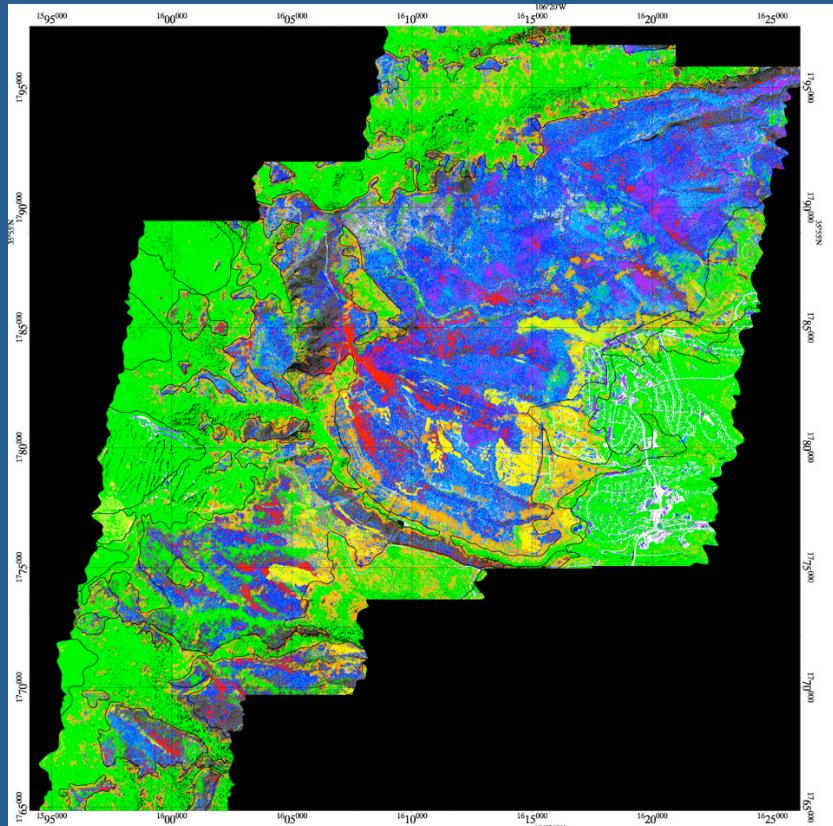
AVIRIS



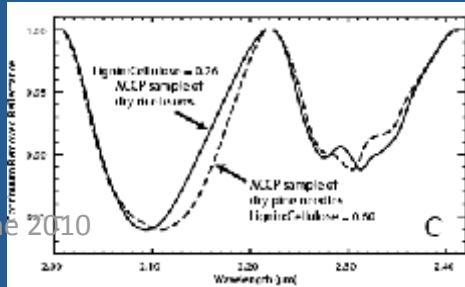
Photo



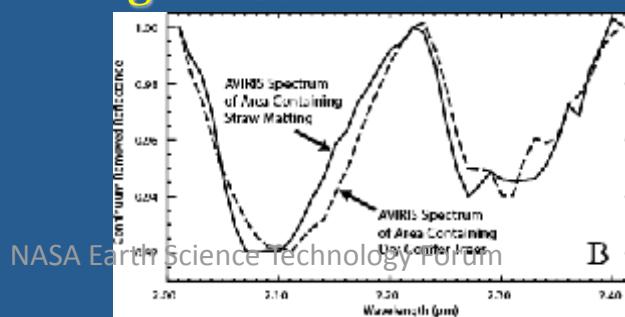
Spectral Fitting Map



Lignin-Cellulose Lab



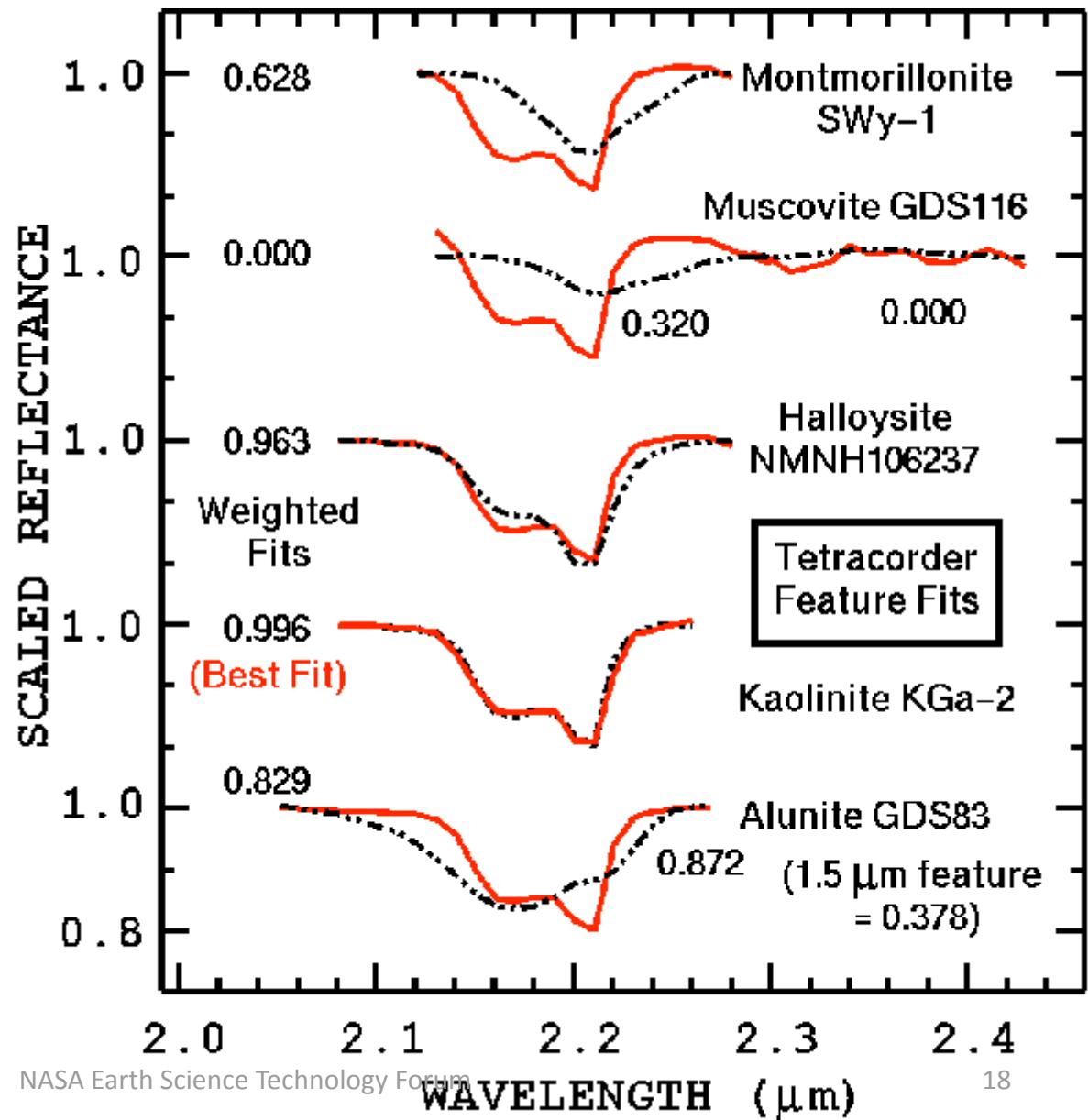
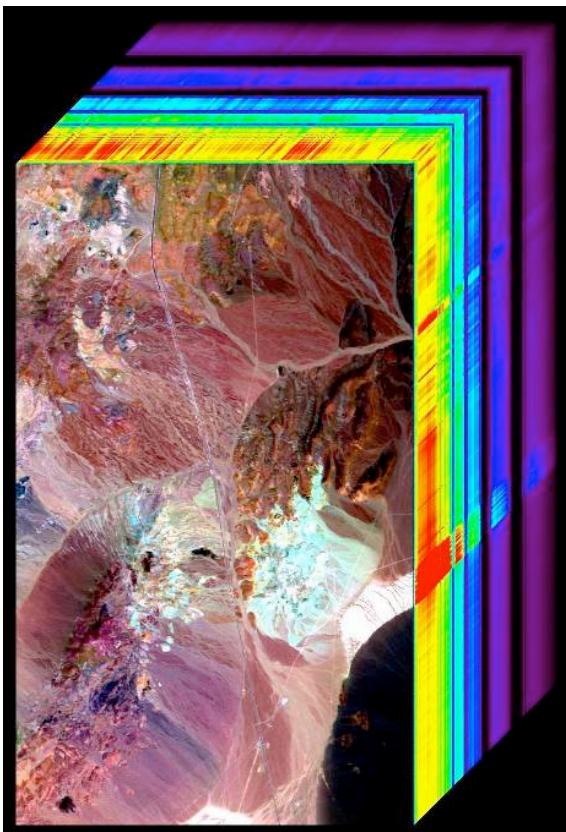
Lignin-Cellulose AVIRIS



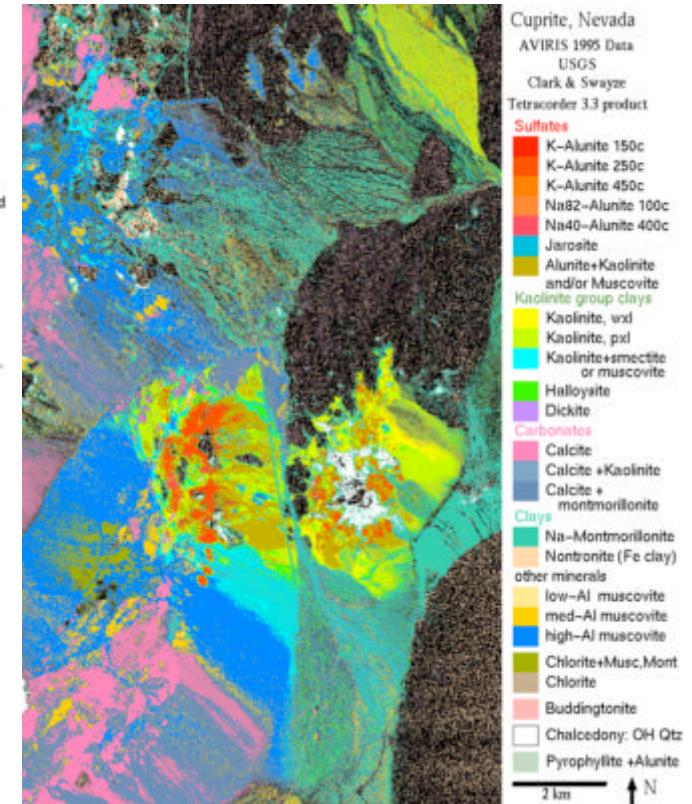
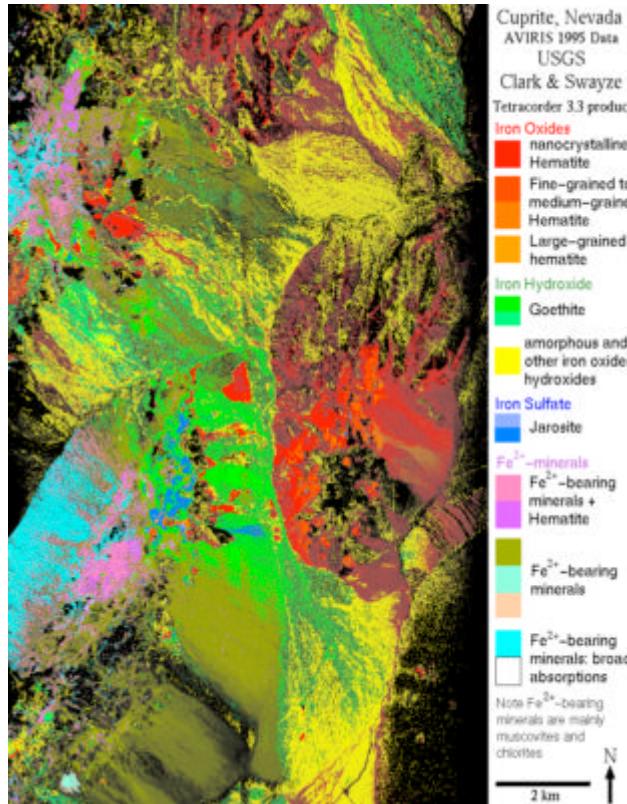
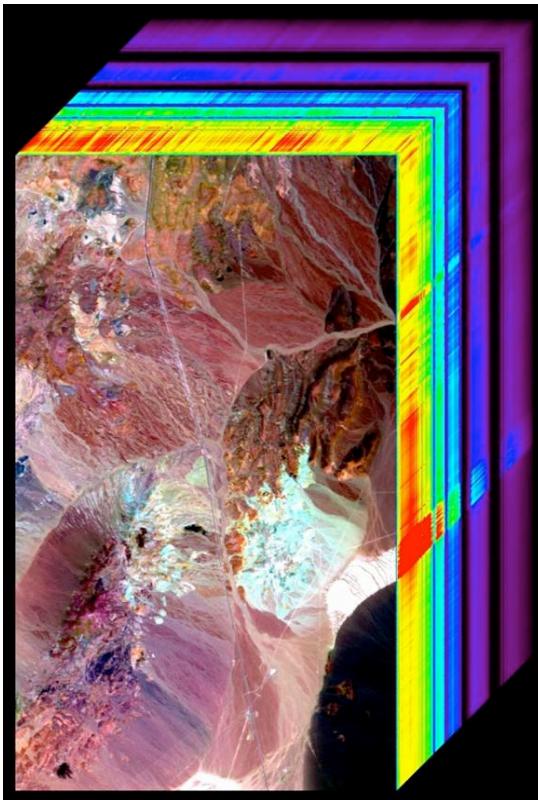
- Ash/Charcoal
- Mineral/Ash
- Mineral-1mm
- Mineral-2mm
- Dry Conifer
- Dry & Green Conifer
- Straw matting
- Straw matting & Green grass
- Green Vegetation

Remote Measurement via Spectral Fitting

Surface mineralogy
Cuprite, NV

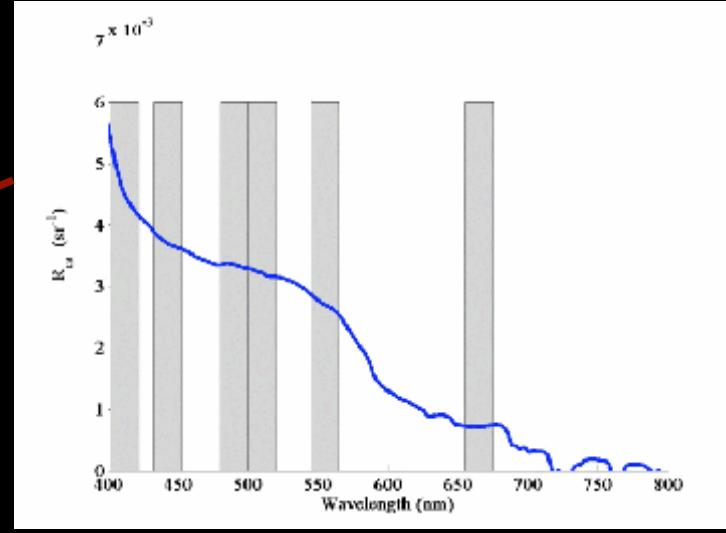
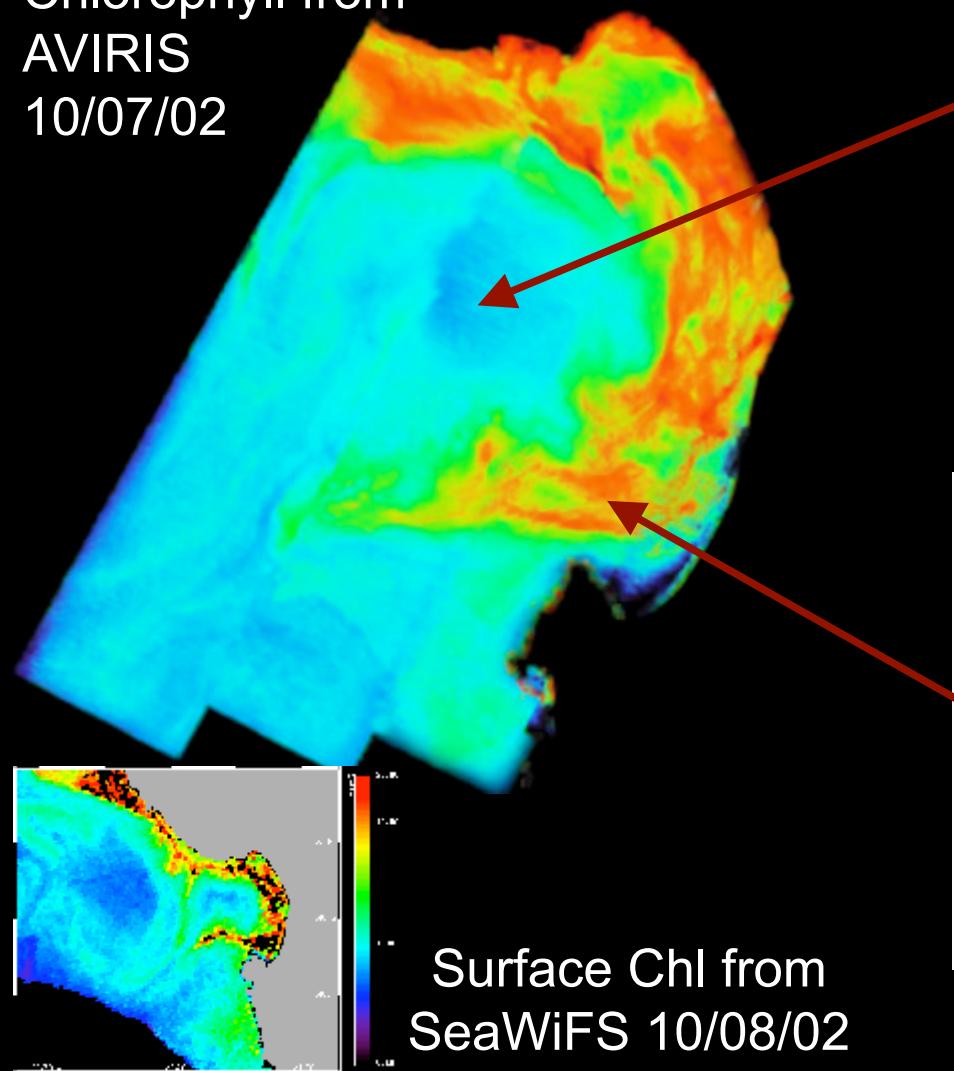


SURFACE COMPOSITIONAL DERIVED WITH IMAGING SPECTROMETER MEASUREMENTS (AVIRIS)

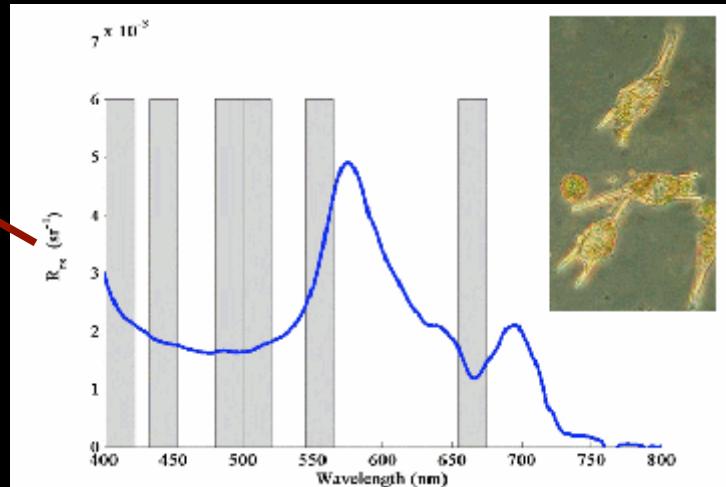


A red-tide bloom in Monterey Bay

Surface
Chlorophyll from
AVIRIS
10/07/02



Surface Chl from
SeaWiFS 10/08/02



SeaWiFS bands miss signal

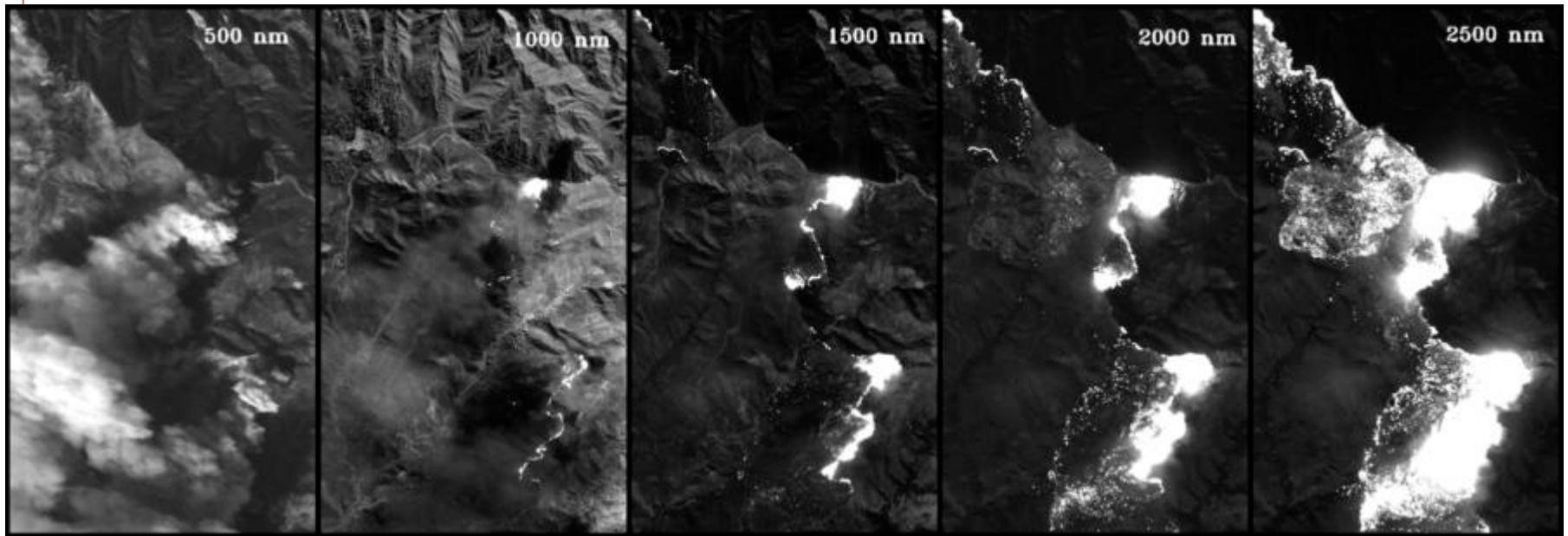
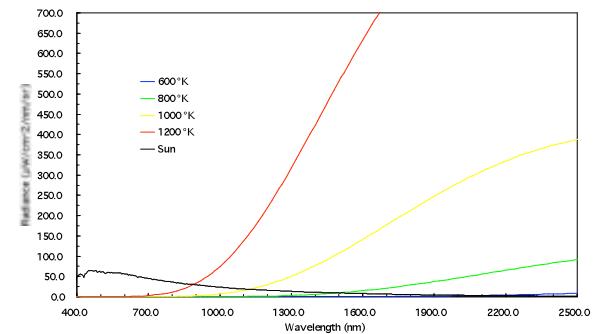
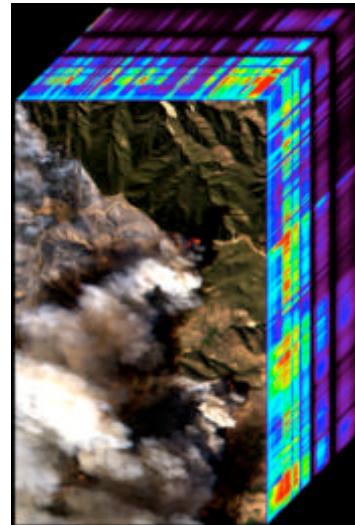
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SIMI VALLEY, CA WILD FIRE



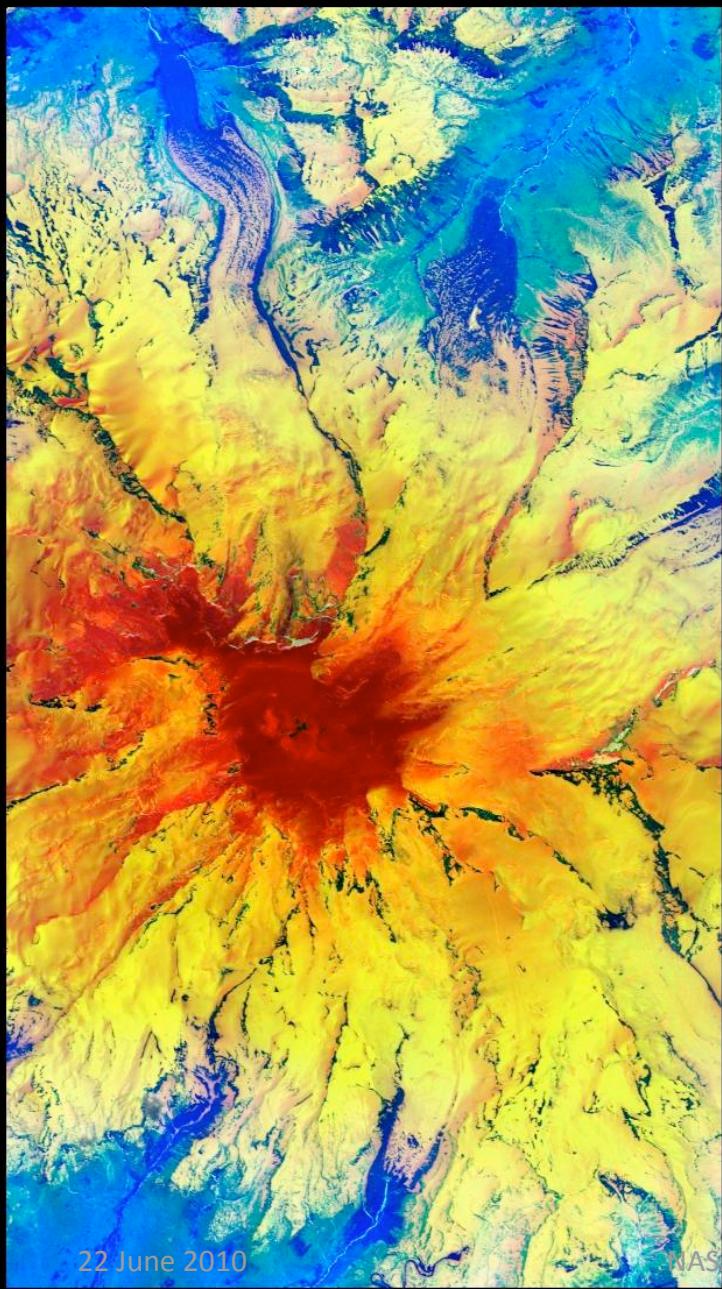
22 June 2010

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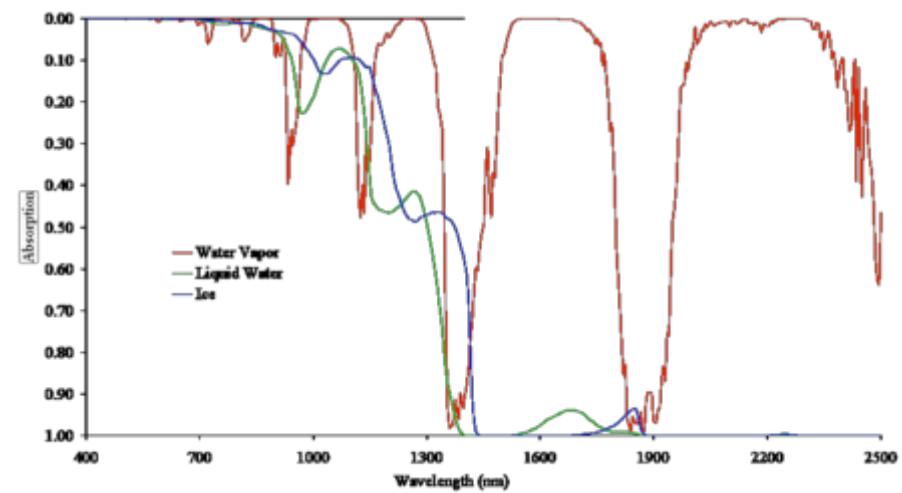
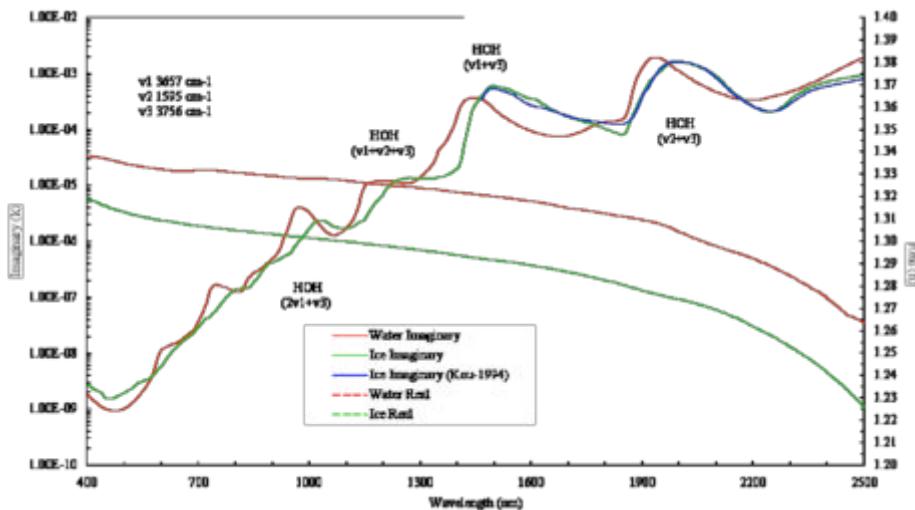
JPL

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AVIRIS: Mount Rainier, WA
Water Vapor, Liquid Water, Ice : BGR



Mount Rainier derived three phases of water (Vapor: blue, Liquid: green, Ice: red) in melting snow environment



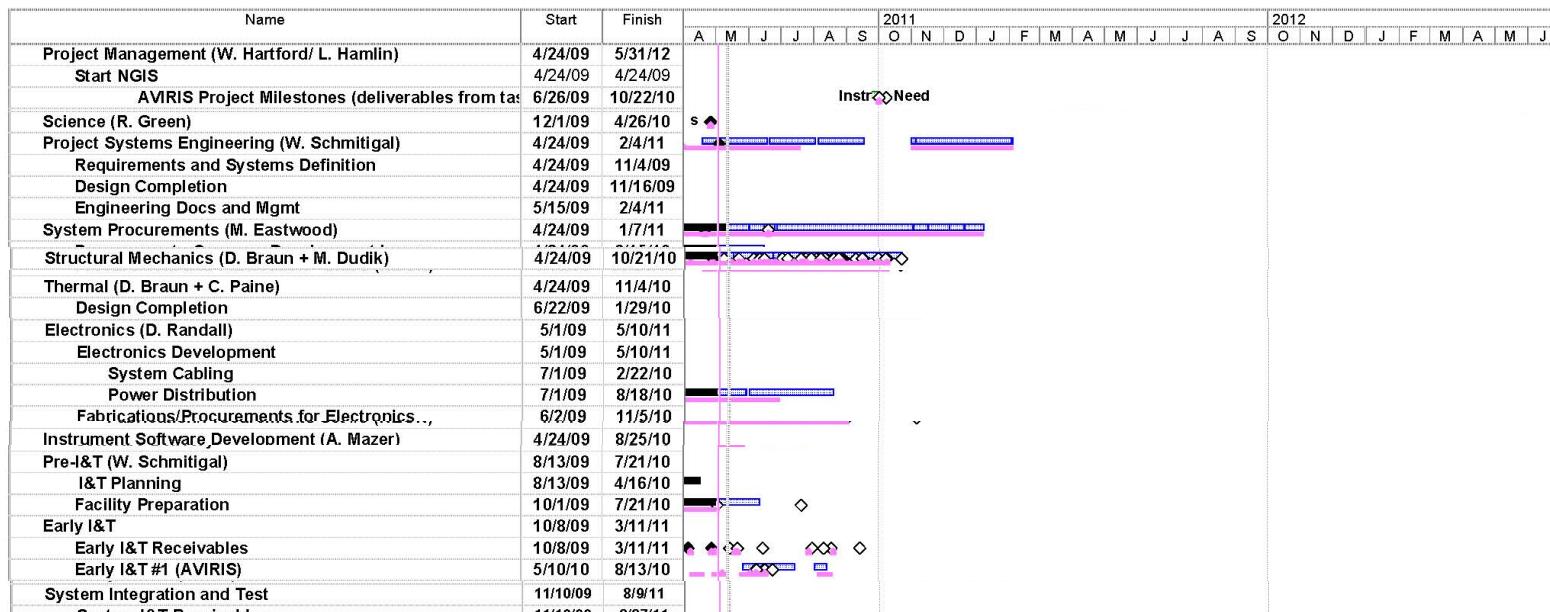


**NEXT GENERATION AIRBORNE VISIBLE / INFRARED
IMAGING SPECTROMETER (AVIRIS-NG)
AMERICAN RECOVERY AND REINVESTMENT ACT
(ARRA)**



Design and build follow-on to AVIRIS (higher sampling, high signal-to-noise)

- Task Budget: \$5M
 - Task Schedule: September 1, 2009 – October 1, 2010





AVIRIS NEXT GENERATION SCIENCE MEASUREMENTS



Spectral

Range	380 to 2500 nm
Sampling	10 nm 5 nm
Response	10 nm 5-7 nm

Radiometric

Range	0 to Max Lambertian
Sampling	14 bit
Calibration	+/- 1 nm +/- 0.1 nm
Signal-to-Noise ratio	>1000 @ 600 nm >2000 >400 @ 2200 nm >1000

Spatial

Field-of-View	34 degrees
Instantaneous FOV	1 milliradian
Spatial swath	2.5 to 11 km @ alt (4 to 20 km)
Spatial resolution	4 to 20 m @ alt (4 to 20 km)

Uniformity

Spectral-Cross-Track	>95% Uniform
Spectral-IFOV	>95% Uniform

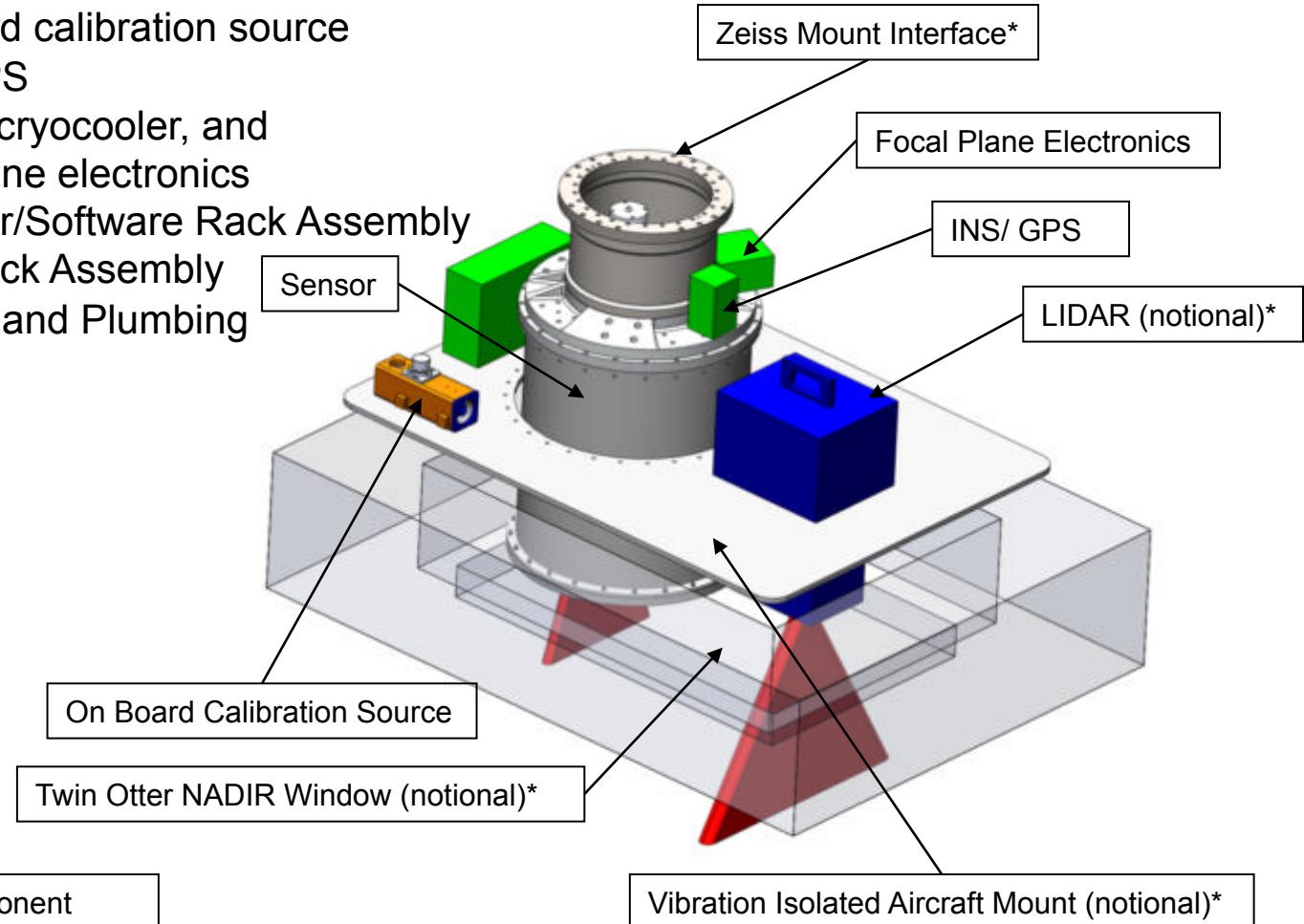


AVIRIS-ng Instrument Configuration



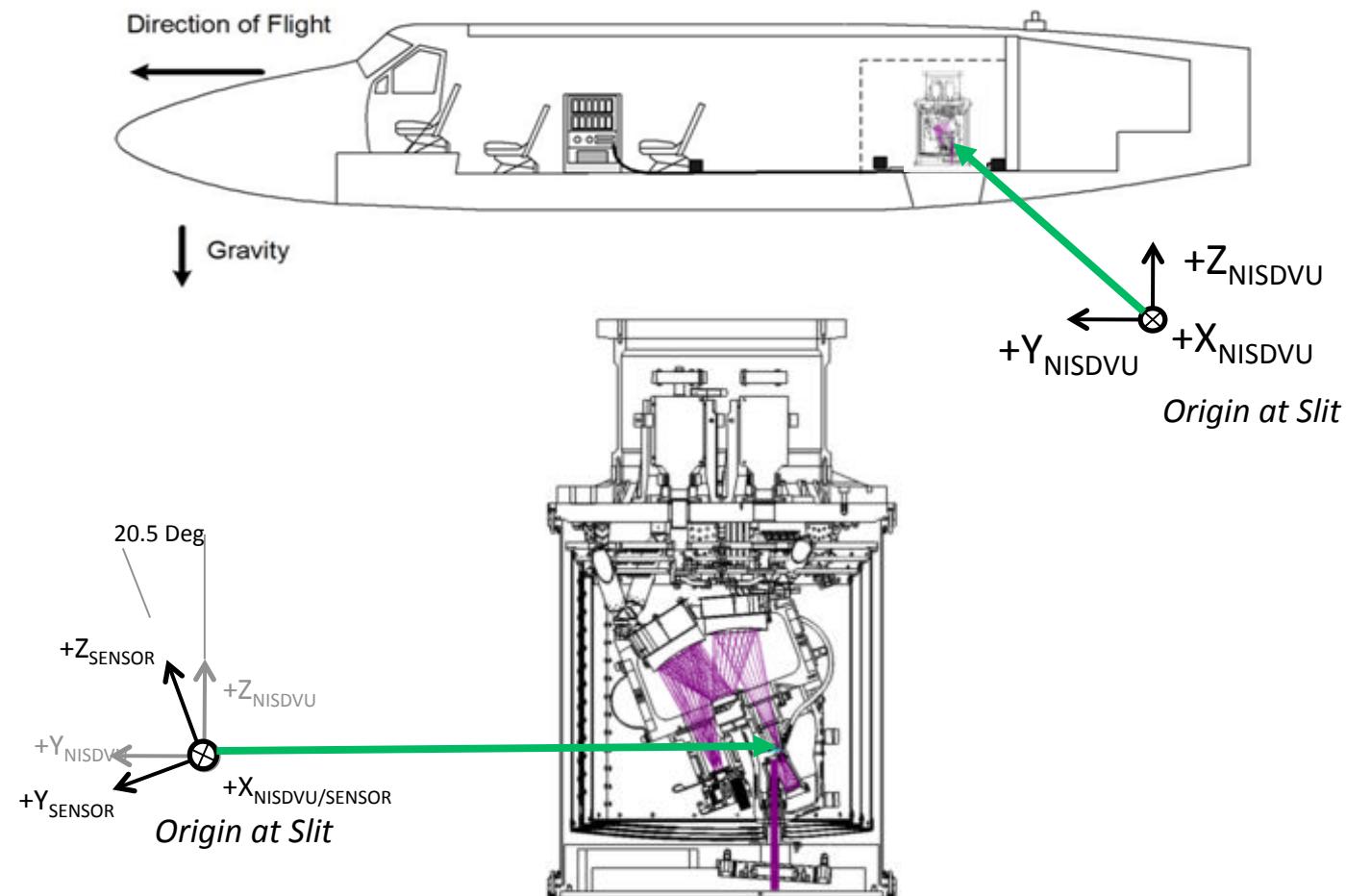
AVIRIS-ng consists of:

- Sensor
- On board calibration source
- INS/ GPS
- Heater, cryocooler, and focal plane electronics
- Operator/Software Rack Assembly
- ECS Rack Assembly
- Cabling and Plumbing





AVIRIS-NG COORDINATES

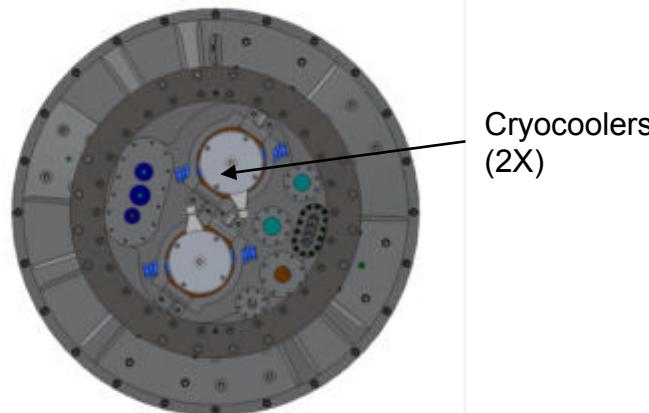




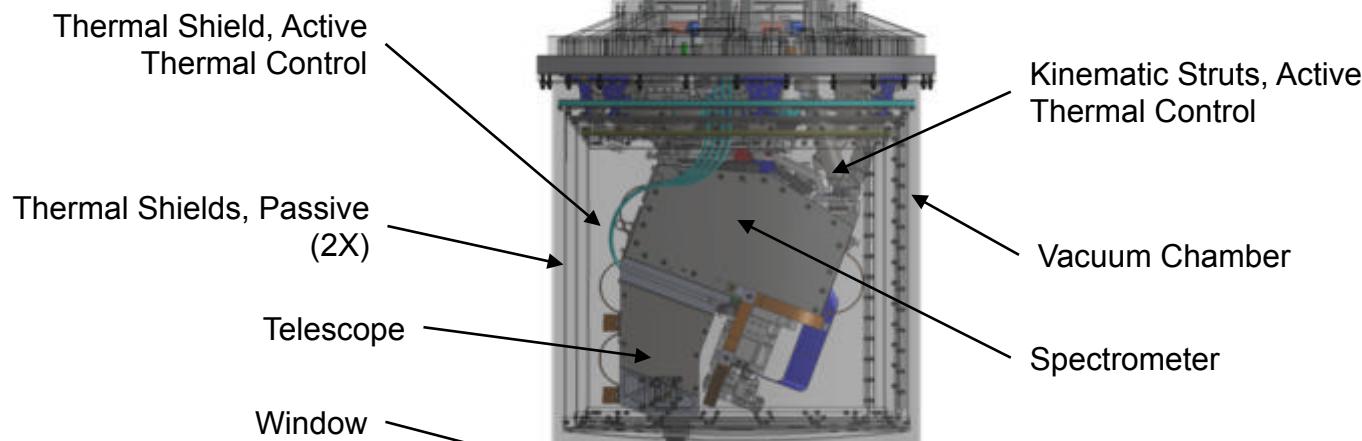
Sensor Configuration



Operator Rack
Power Switches
Blank Panel
KVM Drawer
Blank Panel
Aux Camera Control
LIDAR Operator Computer
Power Switches
NGIS Data Computer w/CL-SAS DVR
Sensor Data RAID Array



Environmental Control Rack	
LS 332 TC	LS 332 TC
Lakeshore 340 Thermal Controller	
cRIO & TCP/IP Hub	
Cryocooler Pwr & Cntrl	
ODBC Pwr, Cntrl, & Thermal Cntrl	
FPIE Power Supplies	
Ion Pump Controller	Aux Therm Ctrl
Power Dist / Conditioning / Switching	
Thermorack Chiller	

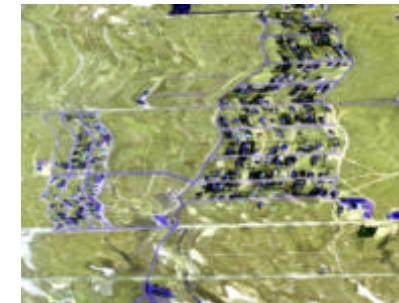




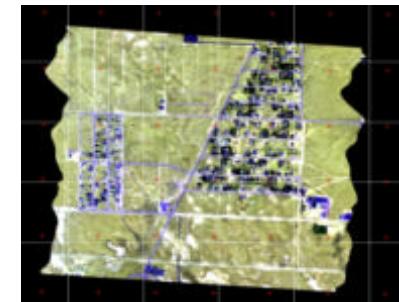
System Operation



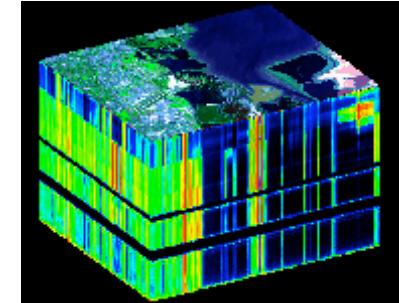
- Pushbroom imager collects image “lines”
- Image lines are collected into data “cubes”. Typical cubes are taken over 20 minute runs.
- Image “swath” is the width of the data cube. Swath dimensions are a function of altitude.
- Cube data are geo-rectified using GPS/INS data to correct for platform motion
- Cubes are radiometrically calibrated against a stable on-board source before and/or after each run
- The NISDVU is designed to be easily integrated with LIDAR



Raw Spatial Data



Geo-rectified Data



The NISDVU Data Cube  28

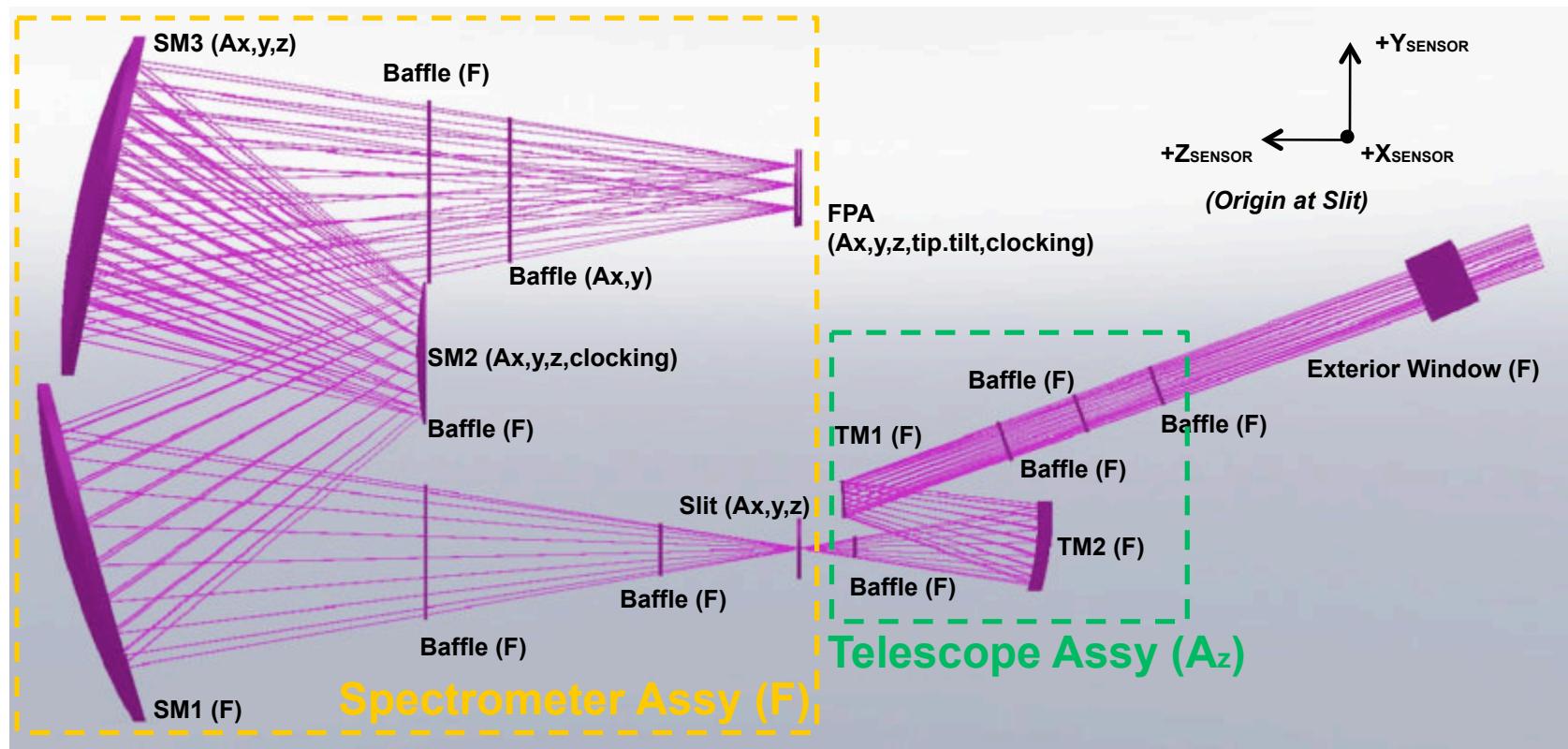


Instrument Optical Architecture



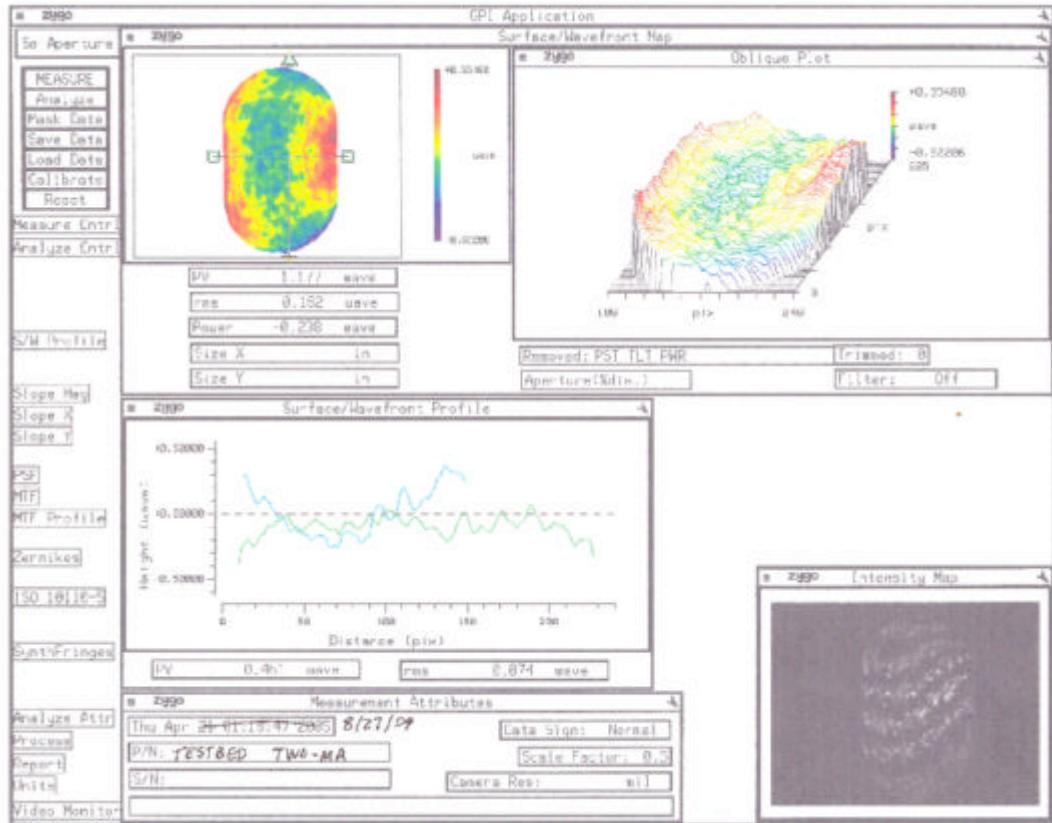
OFFNER SPECTROMETER, 2 MIRROR ANASTIGMAT HIGH-THROUGHPUT TELESCOPE

SIMPLE SYSTEM: 4 MIRRORS, 1 GRATING, 1 DETECTOR





ASSEMBLED TESTBED TELESCOPE



Achieved 1.1 wave p-v wavefront error over entire aperture and field with minutes of assembly time.

Notice non-circular pupil shape.

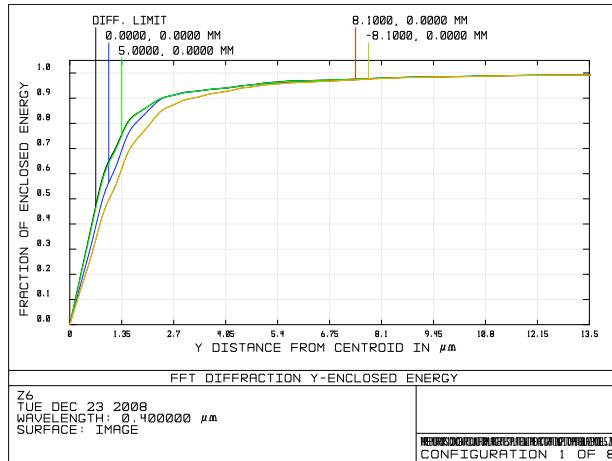
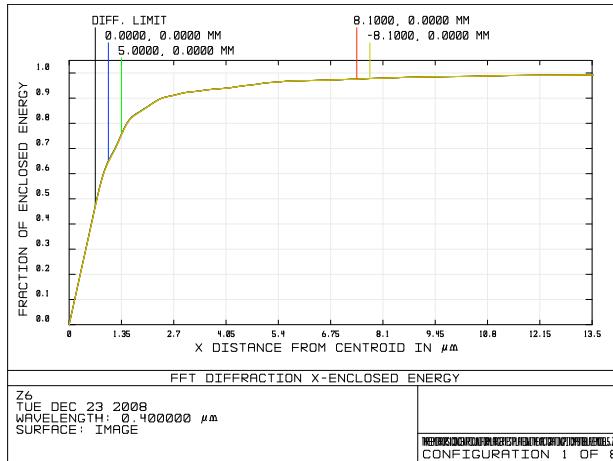
Post-polished aluminum mirrors produced by Axsys.

Previous Axsys mirrors fail mid-frequency specification by a factor of 2-3, impact acceptable (see risks and mitigations later).

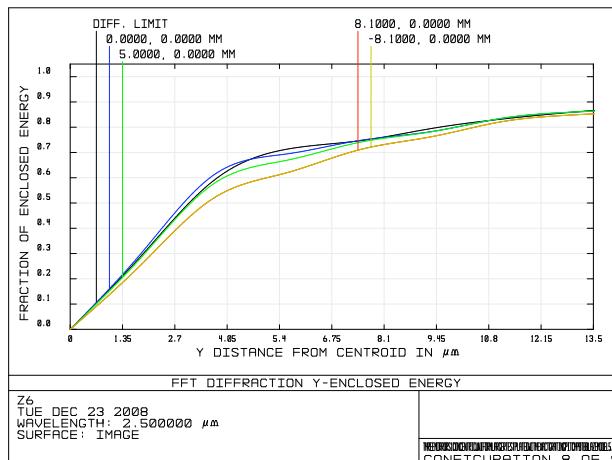
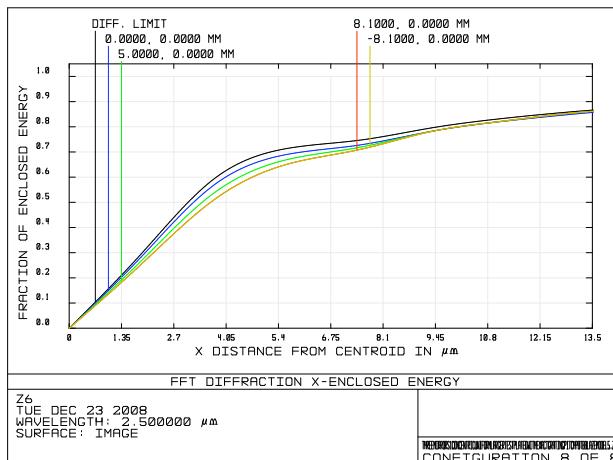


SPECTROMETER PERFORMANCE

(X- AND Y- ENCLOSED ENERGY IN PIXEL)
(ACCOUNTS FOR APODIZATION)



400 nm



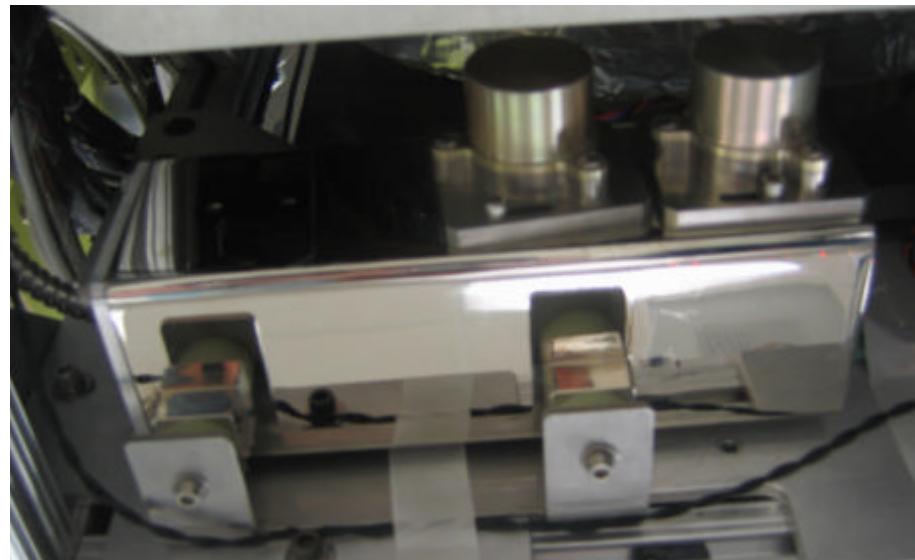
2500 nm



ON-BOARD CALIBRATOR



- Maintain absolute radiometric calibration within 95% across all spectral channels within the FOV
- As designed will also
 - Allow image specific flat fielding to control small radiometric variability and deviations
 - Allow trend monitoring to detect performance issues early
- To meet the requirements we will use a refinement of the on-board calibrator source currently flying and meeting these requirements in AVIRIS and MaRS.

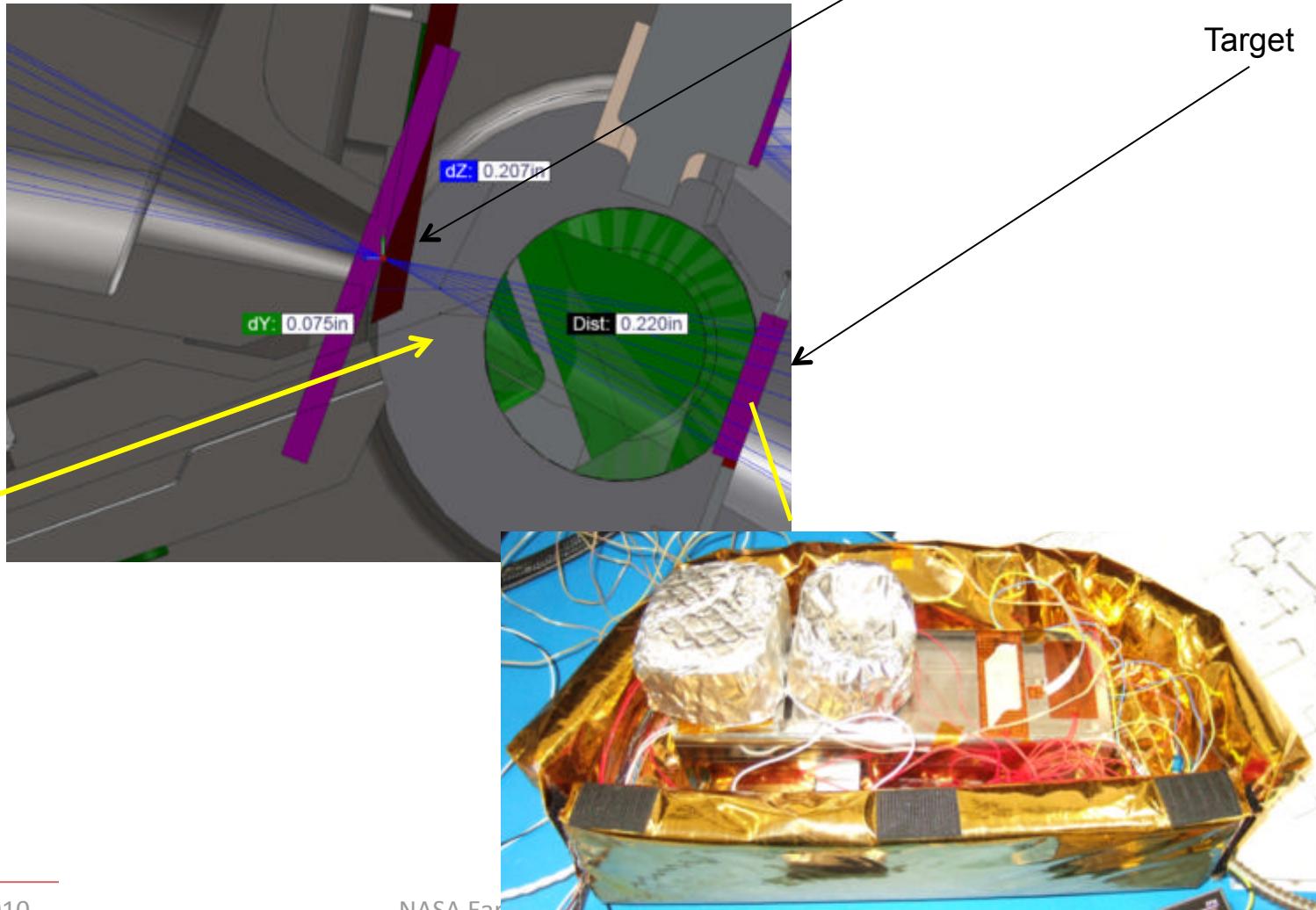




OBC TARGET

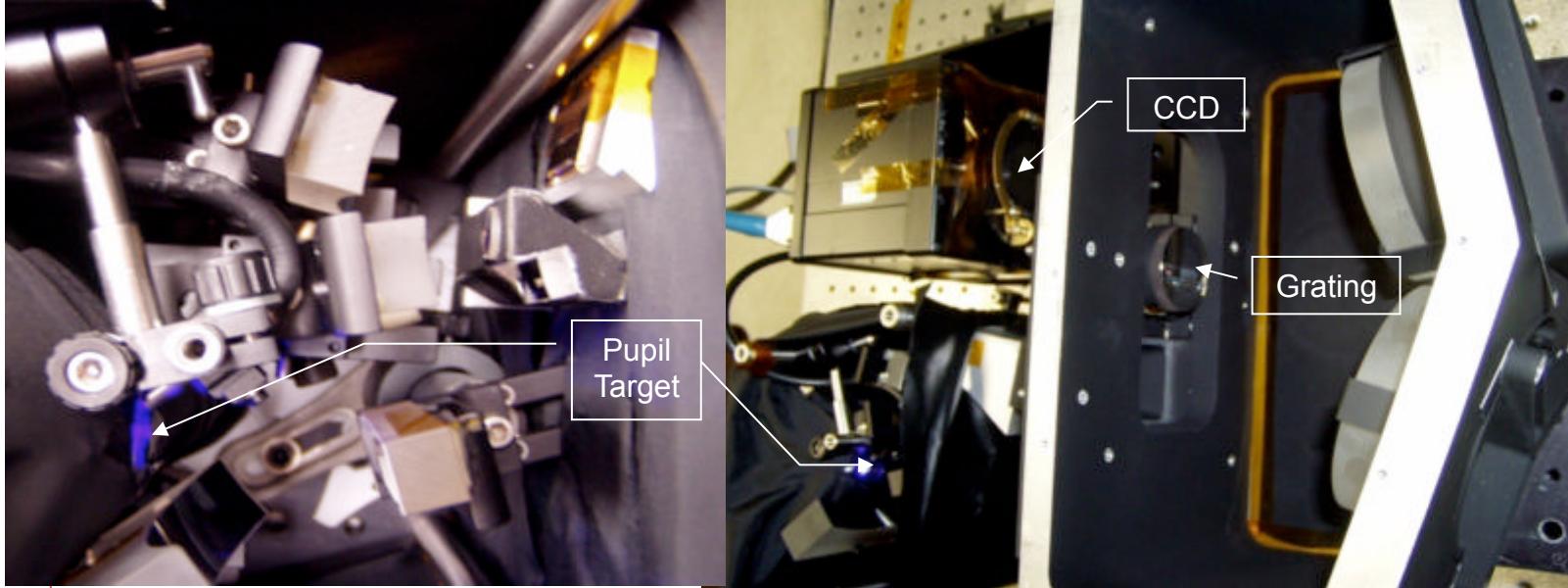


- OBC target in front of slit





TESTBED OBC TARGET TESTING



- Mature OBC approach with extensive heritage
- We have tested the use of the OBC target at the slit entrance
- OBC now in testing for installation into AVIRIS-ng



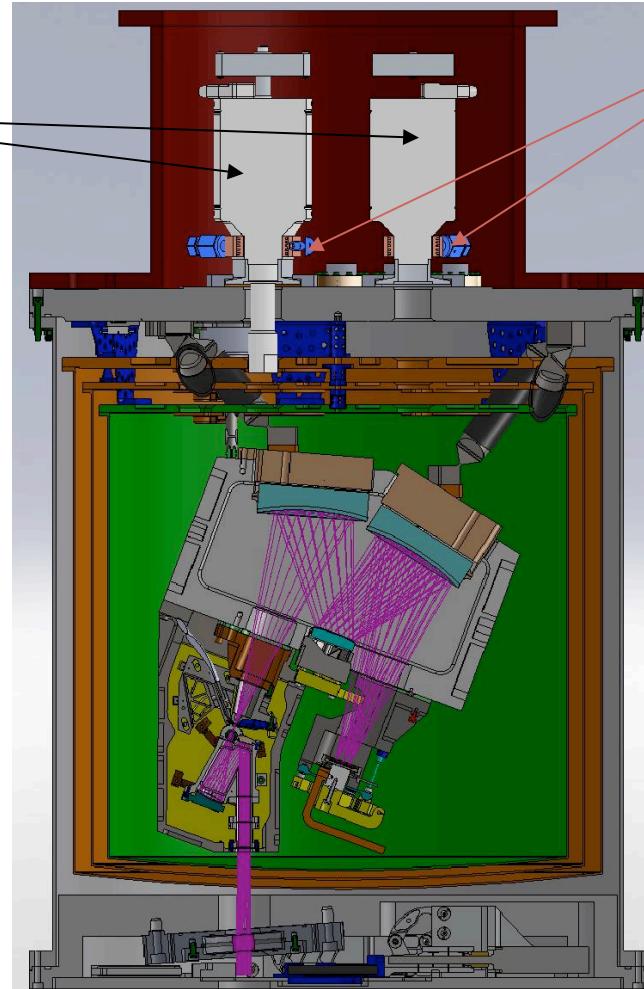
INSTRUMENT CRYOVACUUM ARCHITECTURE



cryocoolers, power supplies, and drive electronics



vacuum-ion pump and gauge
decision on use of the ion pump is pending system performance evaluation



NASA Earth Science Technology Forum

recirculating fluid chiller
removes heat from cryocoolers



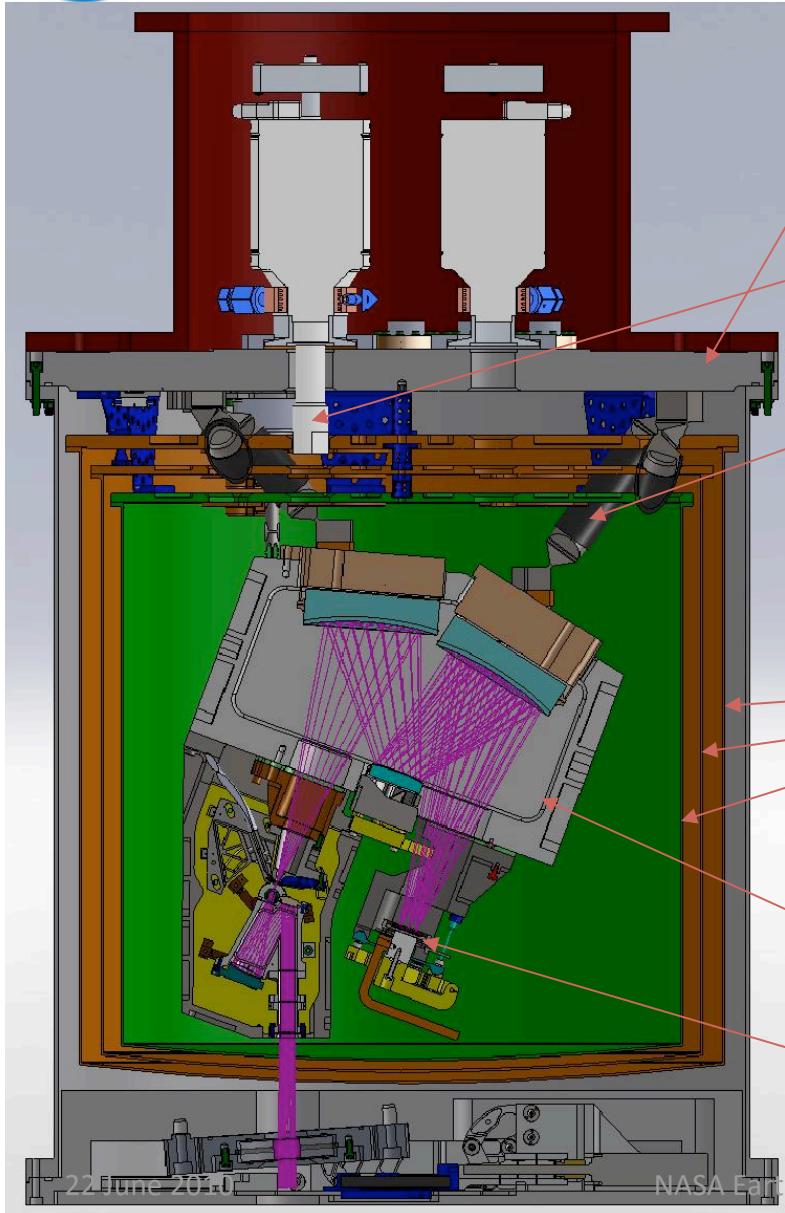
temperature monitoring and control
maintains FPA temperature stability and optical alignment (multiple locations on instrument)



cryo wiring harness is fabricated in-house; not shown



CRYOVACUUM SUBSYSTEM OPERATIONAL CONDITIONS



Vacuum enclosure carries window, cryocoolers, all electrical and vacuum feedthroughs, all mechanical loads

cryocooler cold ends ~110 K steady-state
Both coolers run for cooldown, then one switches off.
(Straps to inner shield and spectrometer are not shown)

Thermally-isolating kinematic struts support instrument from warm outer vacuum shell (318 K maximum environmental temperature)

Two “floating” and one actively-cooled radiation shields reduce heat input to spectrometer:
1st shield ~288 K, floating
2nd shield ~235 K, floating
actively-cooled shield ~125 K, controlled

Spectrometer and telescope temperature are actively controlled to <50 mK variation

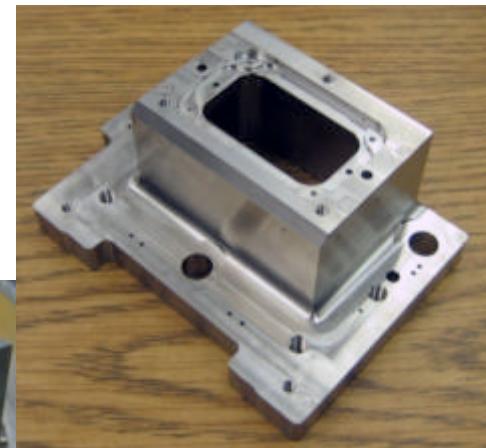
FPA is cooled to <140 K and actively controlled.
(Dedicated high-conductance strap to cold sink not shown.)



DESIGN /FABRICATION STATUS



- Design complete
- Mechanical and Cyro-vacuum parts now in fabrication
- 45% of parts now received, complete in August 2010

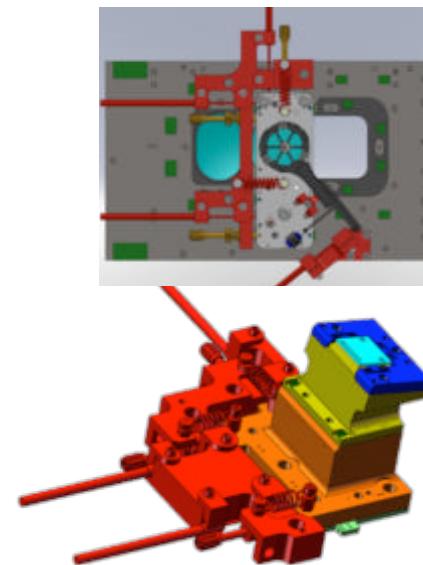




I&T STATUS



- Facility Ready
- Personnel trained
- GSE design complete – now in build
- Plans, procedures and storyboards in review
- “Pre” I&T now underway
- I&T begins July 2010



22 June 2010

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Next Generation Imaging Spectrometer
(NGIS)

System Integration and Test Plan

May 11, 2010

Rev. DRAFT

Prepared by

W. S. S. [redacted]

NGIS Instrument I&T Lead

Reviewed by

L. [redacted]
NGIS Instrument Deputy Manager

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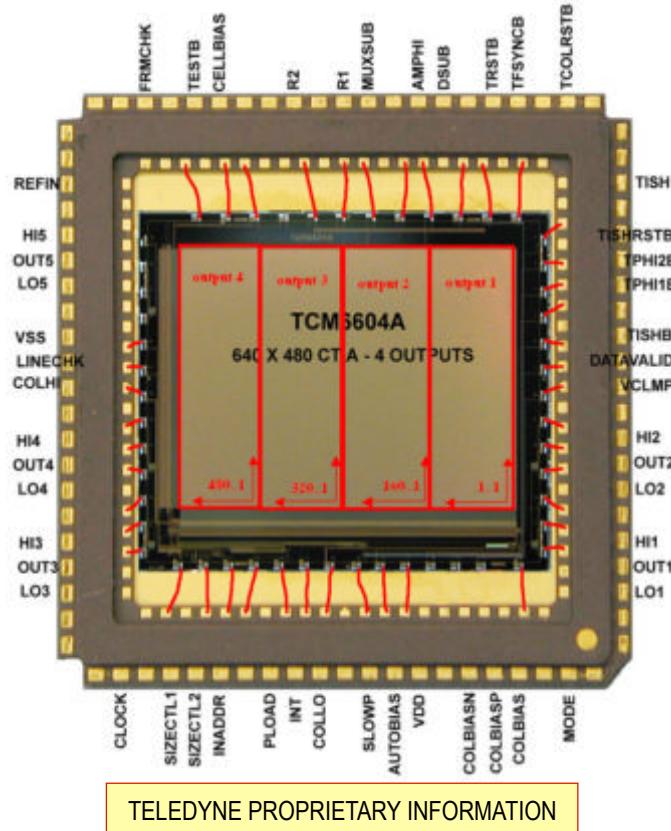
ELECTRONICS DESIGN- MARS HERITAGE



- Focal Plane Interface Electronics
 - only 1 analog board (only 1 focal plane assembly)
 - FPGA programming has been modified
 - 1 Focal Plane Assembly (no multiplexing)
 - uses MaRS 14-bit capability (MaRS only used 12 bits)
 - MaRS pedestal shift problem identified and corrected
 - GPS time tag incorporated (requirement)
 - 1st pixel clock extension incorporated (for warmer FPA operation)
 - line buffers enable non-interleaved data output
 - larger FPGA (pin-compatible and EEPROM compatible)
 - Camera Link Data Path (to next-generation frame grabber card)
 - OBC and control electronics
 - C-MIGITS III INS/GPS



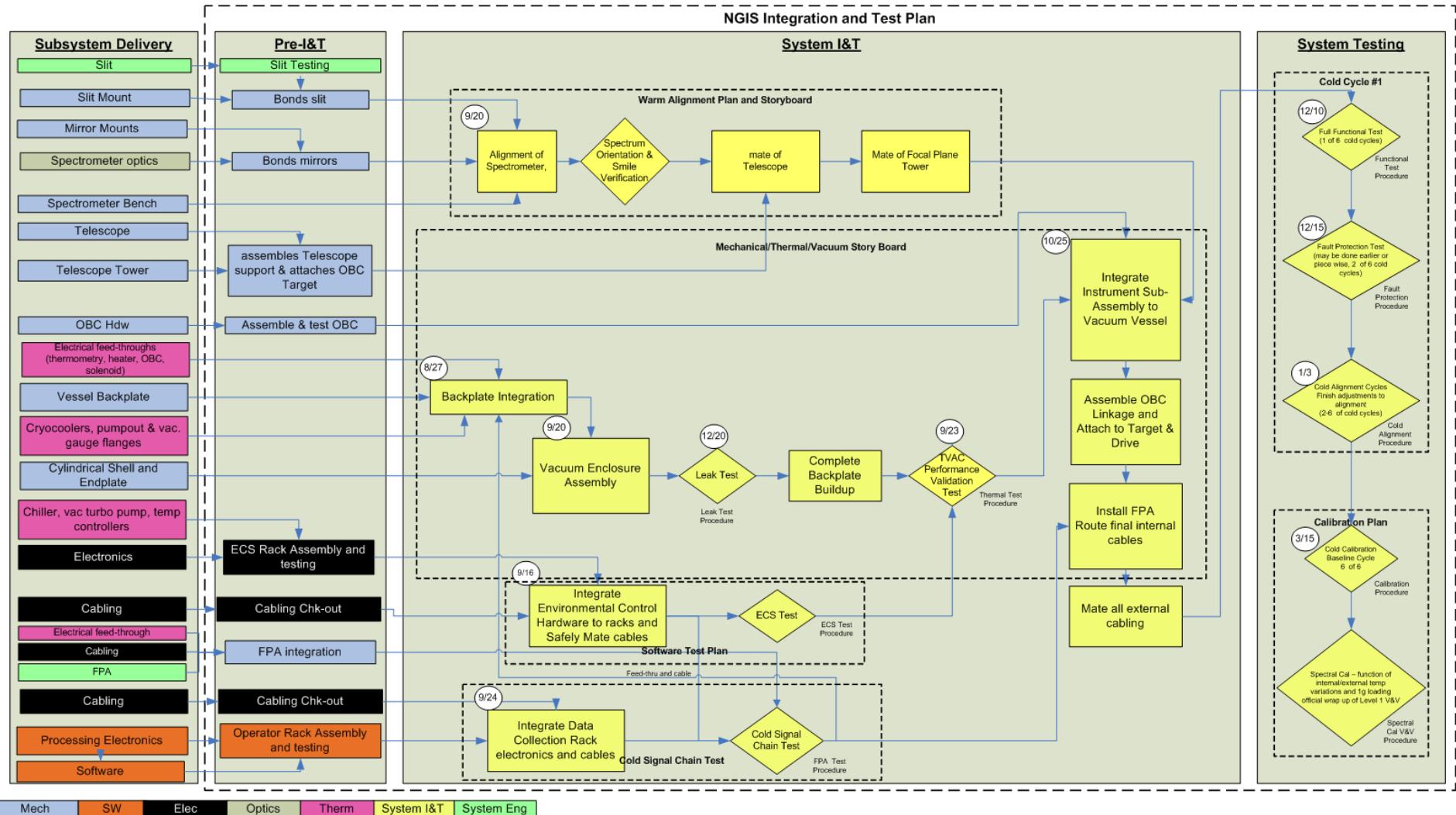
FOCAL PLANE ARRAY



- Well-known Teledyne TCM6604A
- Capacitive transimpedance amplifier
- Snapshot imaging
- 4 video outputs, 1 ref output (unused)
- 4 clocks, 7 bias voltages
- Power dissipation: ~ 60mW
- Read noise: 120 e- [rms]
- Amplifier glow: ~ 100 e-/sec
- Recent understanding of a phenomenon affecting blue-end QE stability is yielding processing steps for mitigation
- Devices are straightforward to fabricate using current vendor processes
- Custom JPL drive electronics exist and work extremely well with this FPA



I&T FLOW DIAGRAM



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MECHANICAL/THERMAL/OPTICAL INTEGRATION PROCEDURE



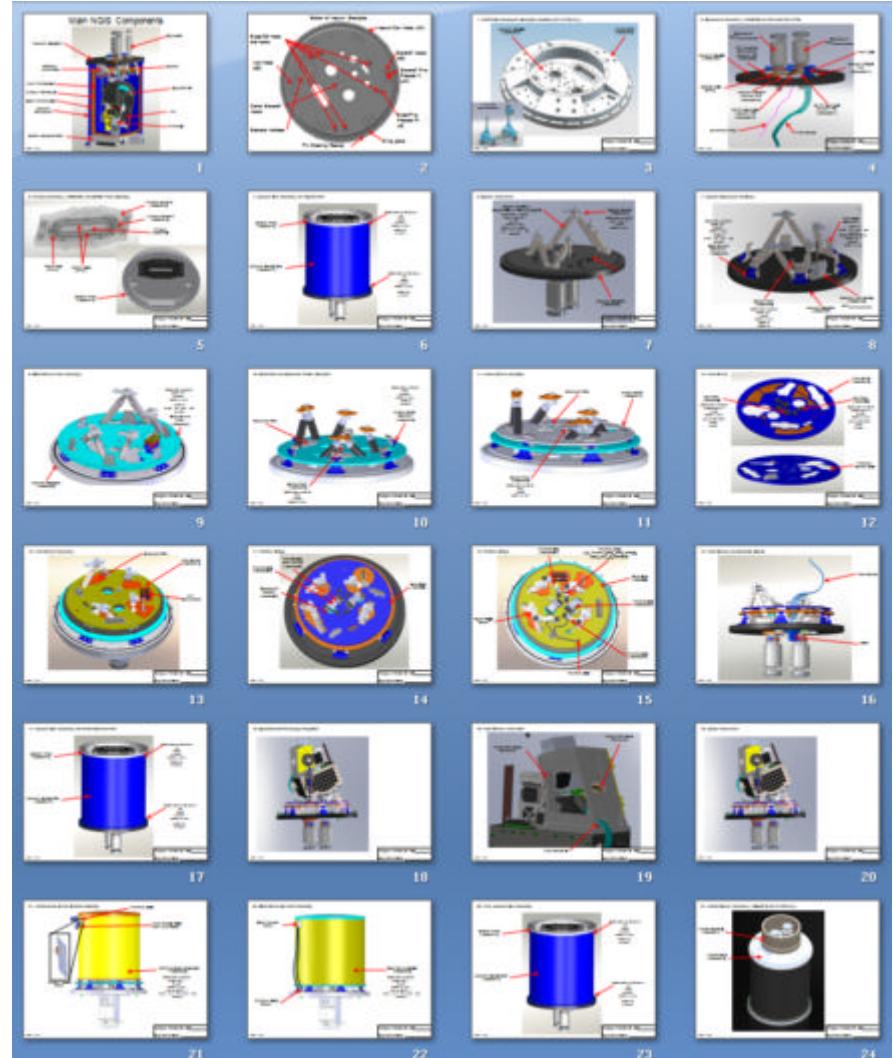
NGIS Integration Procedure

Alfonso Feria

Version 0.6
June 2nd, 2010

- Purpose
 - Story board for the integration process
- Description
 - Provide hardware information (adhesives, fasteners, parts lists, etc.) to aid during integration
 - Provide installation notes
 - Record bolt torque values
 - Document torque wrench information
 - Collect and document thoughts and notes before, during and after integration
- Status
 - Done

22 June 2010



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WARM ALIGNMENT STORYBOARD



JPL

neon
National Ecological Observatory Network, Inc.

NGIS: NEON Warm Alignment Storyboard

16 June 2010
Prepared by:
Holly A. Bender, Optics

16 June 2010 0004 controlled document



Description:

A step-by-step visual guide to the warm alignment process, including GSE needed and verification steps

Status: Done

Successful peer review 06/09/10

Note: Slides above represent only a portion of the full document



SUMMARY



- The science enabled by a high uniformity and high signal-to-noise ratio imaging spectroscopy is well established
 - AVIRIS referenced in > 600 refereed journal articles
- We understand the key measurement characteristics that are needed
- We have developed the right set of requirements flowing from the science
- Throughout the AVIRIS-ng effort, these science traced requirements will be tracked, balanced, and reported to assure the instruments are ready for the next generation science
- The design is complete and being manufactured
- AVIRIS-ng is being built and will be integrated this summer