



Laser Frequency Stabilization for GRACE-II

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Acknowledgements

- Co-Is:

- W. Klipstein, N. Yu (JPL), M. Stephens, J. Leitch (BATC)

- Collaborators:

- Science advisors: M. Watkins (JPL), R. Nerem, P. Bender (CU)

- Laser advisors: K. Danzmann, G. Heinzel (U. Hannover)

- Cavity/enclosure development

- J. Decino, R. Pierce, C. Pace, M. Davis (BATC)

- Modulators/electronics

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- Sponsor

- NASA Earth Science Technology Office / Instrument Incubator Program

GRACE Mission

Science Goals

High resolution, mean & time variable gravity field mapping for Earth System Science applications.

Mission Systems

Instruments

- KBR (JPL/SSL)
- ACC (ONERA)
- SCA (DTU)
- GPS (JPL)

Satellite (JPL/DSS)

Launcher (DLR/Eurockot)

Operations (DLR/GSOC)

Science (CSR/JPL/GFZ)

Orbit

Launch: March 2002

Altitude: 485 km

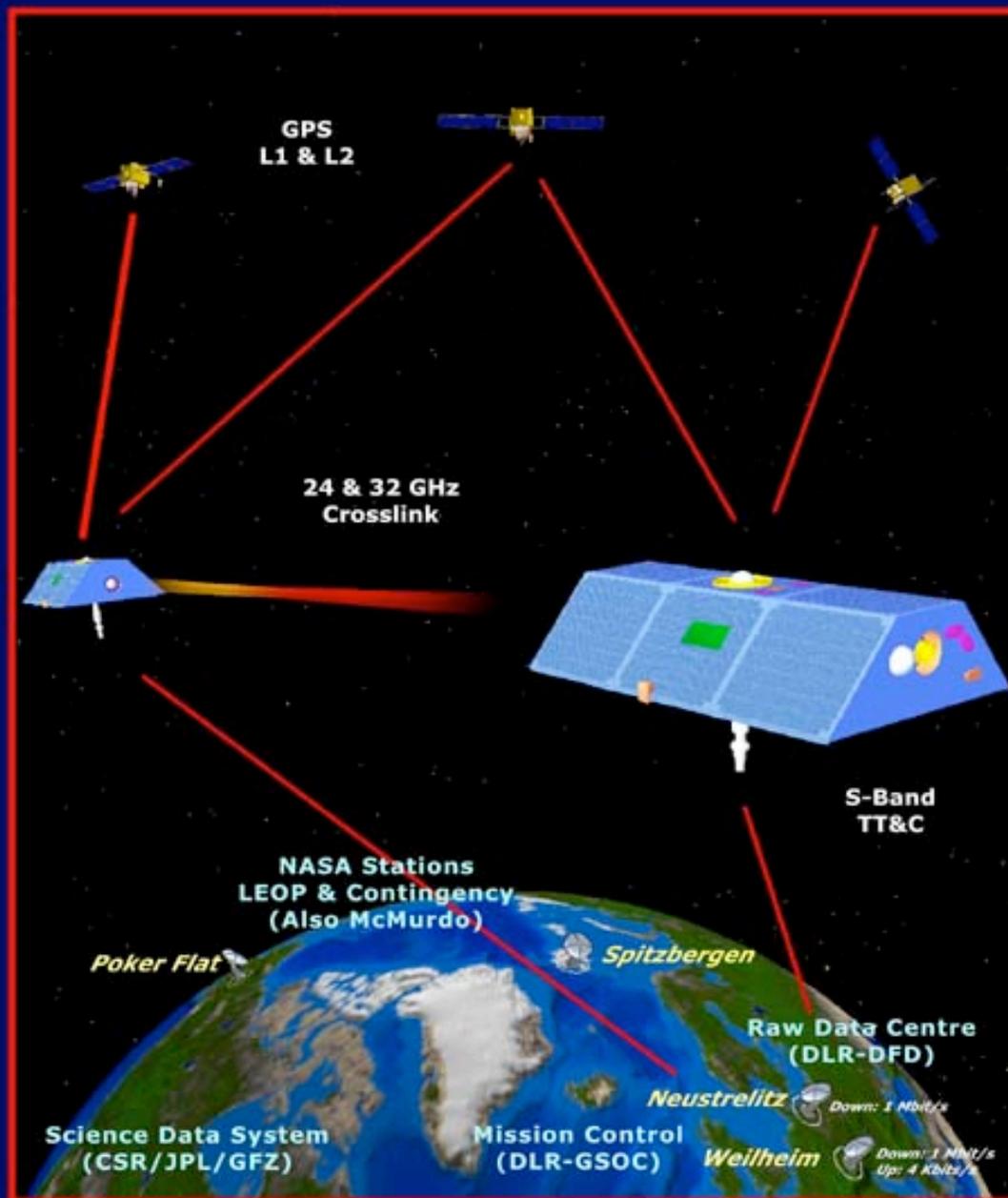
Inclination : 89 deg

Eccentricity: ~ 0.001

Lifetime: 5 years

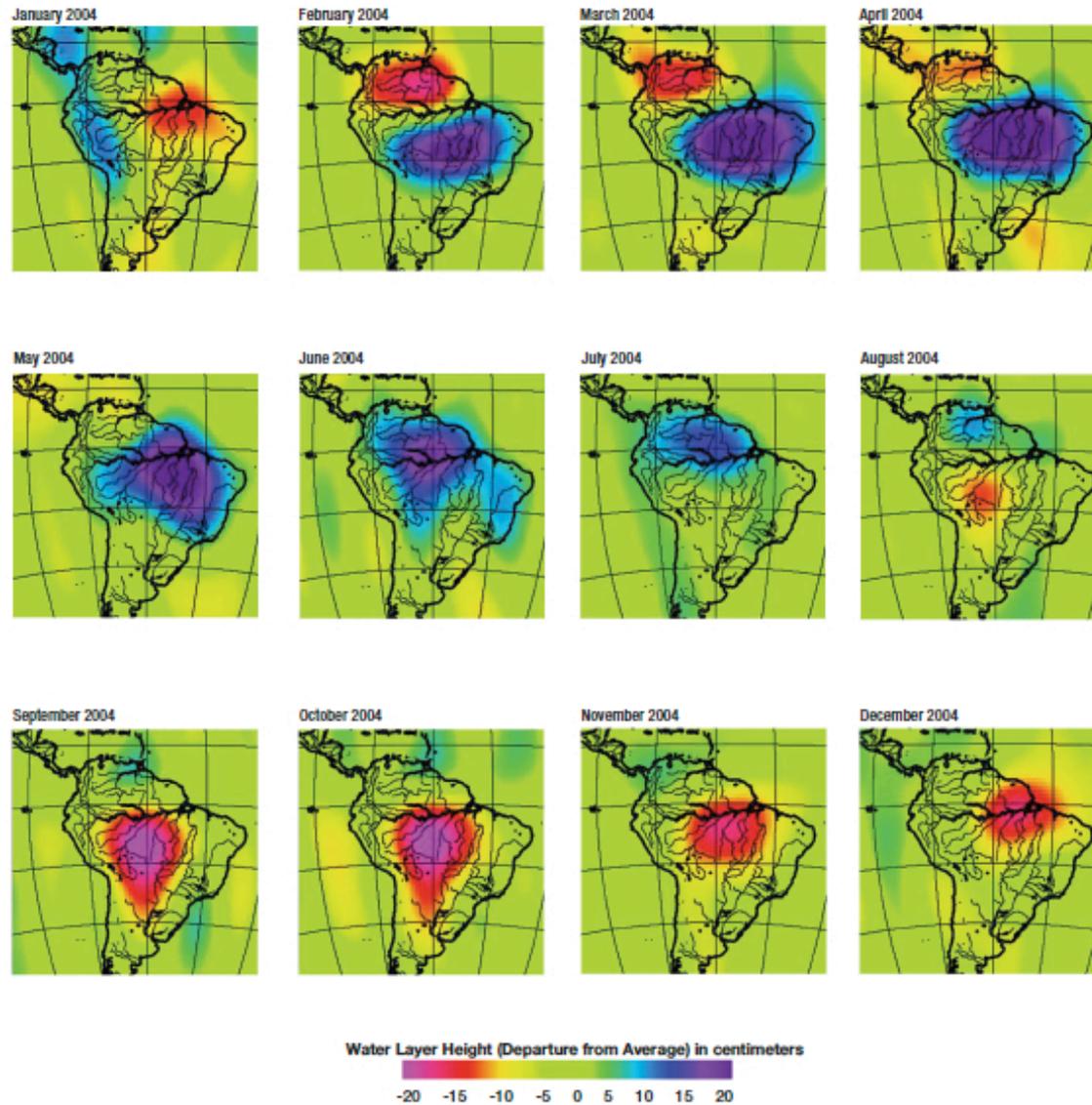
Non-Repeat Ground Track

Earth Pointed, 3-Axis Stable





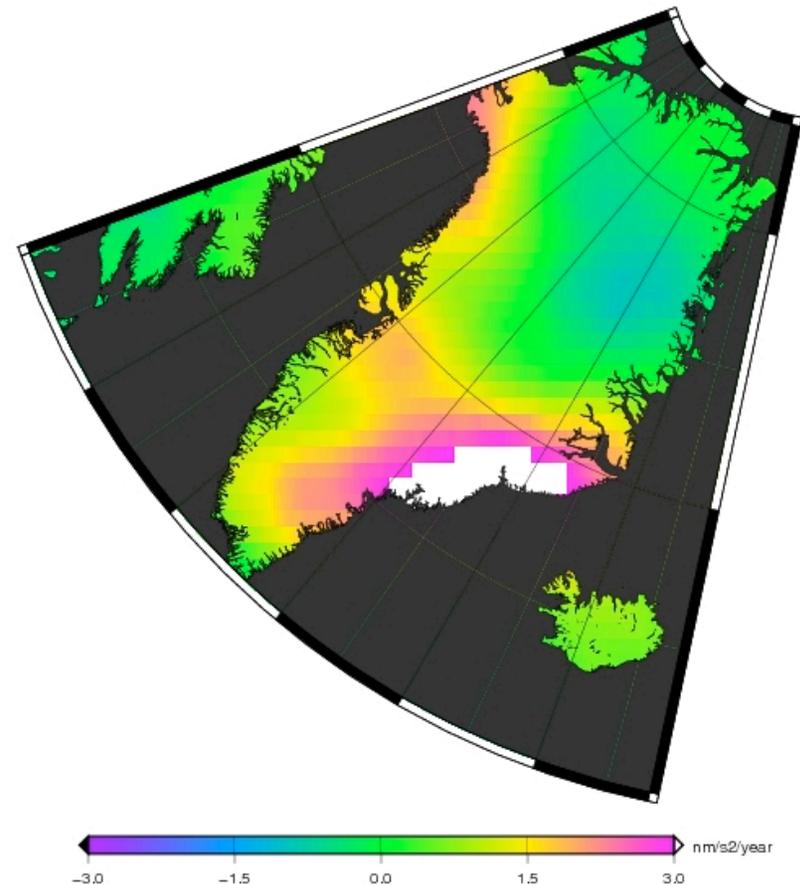
Seasonal Mass Change in Amazon





Ice Mass Loss

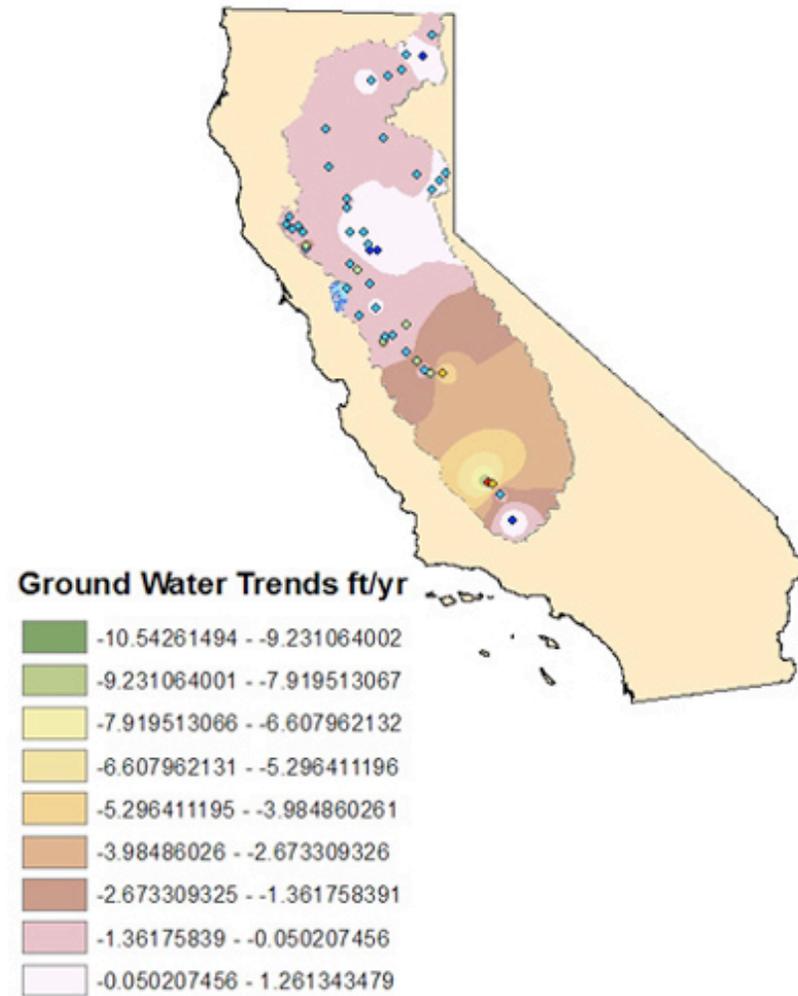
- GRACE can see changes in mass over large regions
- Effect of most significant features (e.g. glaciers) is spread due to limited spatial resolution
- Improved measurement system will allow better identification of features where mass is changing





Subsurface Aquifer Changes

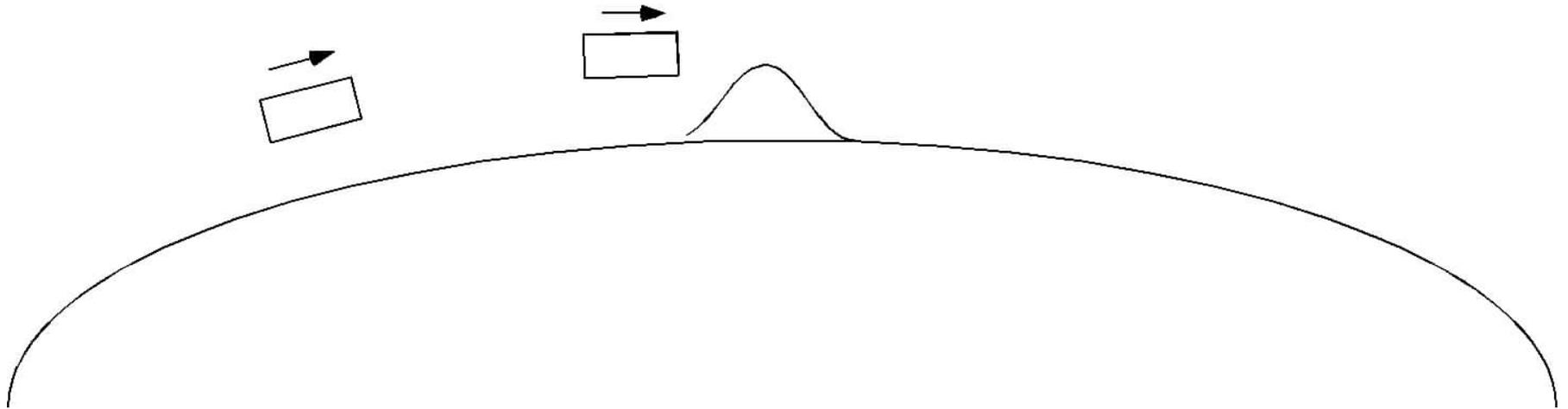
- Groundwater storage changes in the Sacramento-San Joaquin River Basins from GRACE and supplementary data,





GRACE Measurement Concept

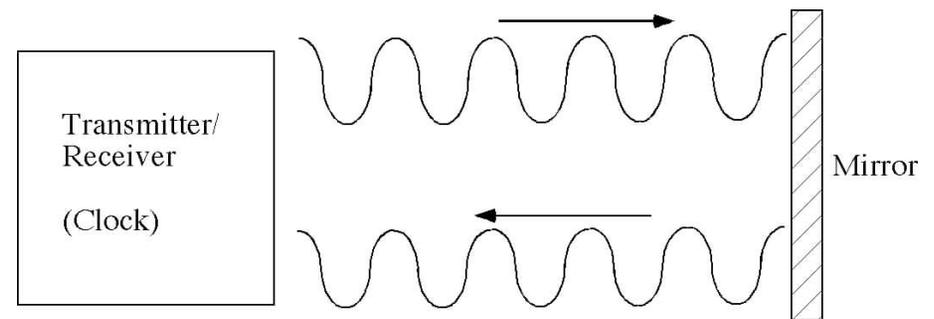
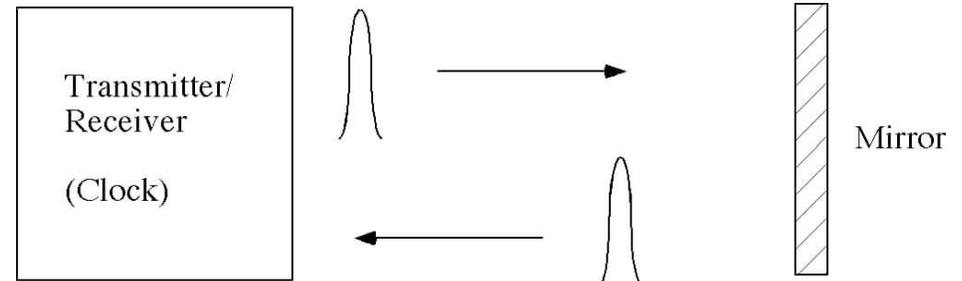
- Earth gravity features affect lead/trailing spacecraft at different times
 - Lead spacecraft encounters feature first
 - e.g. lead spacecraft speeds up towards mountain
 - Range to trailing spacecraft increases
 - Any unknown non-gravity forces acting on spacecraft also affect range
 - Calibrated out using accelerometer or drag-free sensor system





Ranging Measurement Methods

- Range is determined by round-trip light time
 - Pulsed light is used in SLR/LLR where photon rates are low
 - Coherent signals allow use of phase delay for higher accuracy
 - Range ambiguity by integer number of wavelengths can be resolved with modulation if required





Dual-one-way versus Transponding

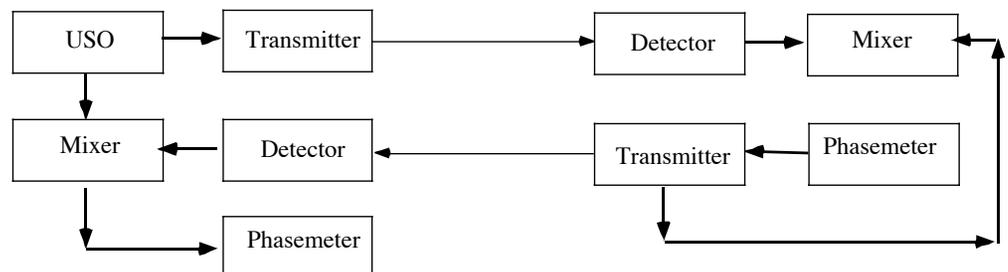
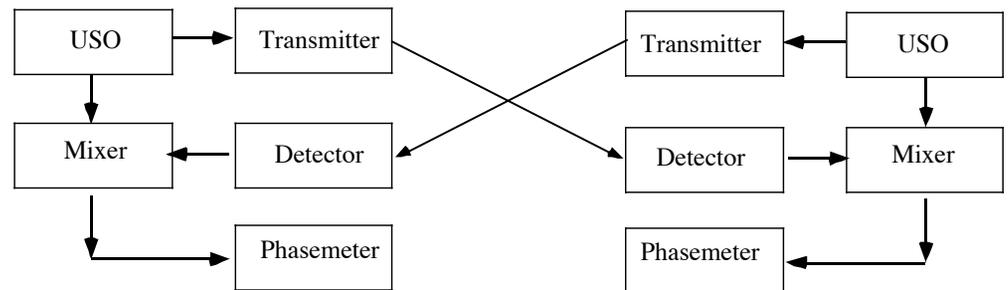
- GRACE uses independent transmission/detection at each spacecraft

- Combination of data on ground determines range

- Laser ranging will lock laser to frequency reference on one spacecraft and lock laser to received laser on seconds s/c

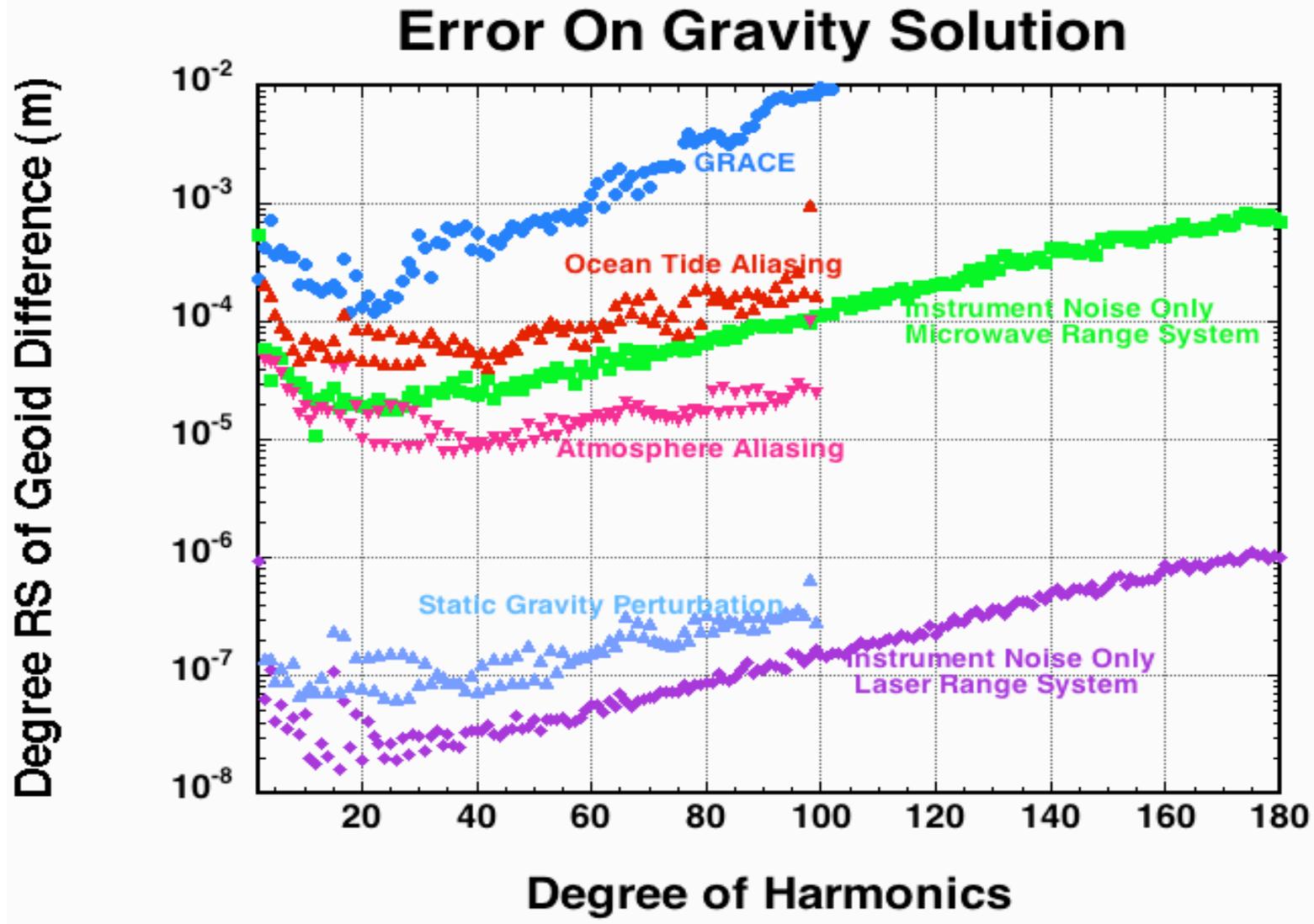
- Otherwise laser frequencies would be too far apart

- Fast phase meter needed for locking function





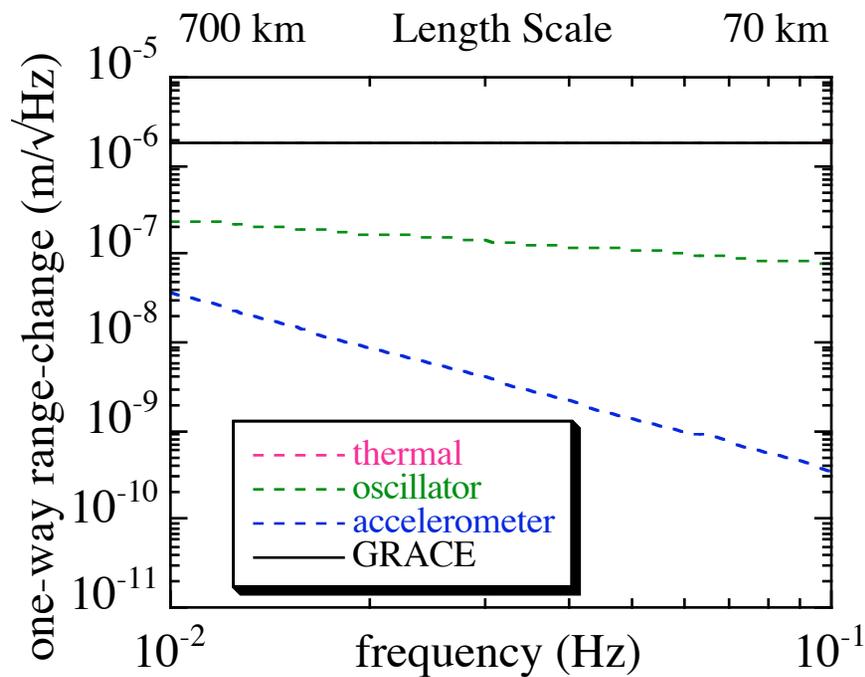
GRACE-II Simulation



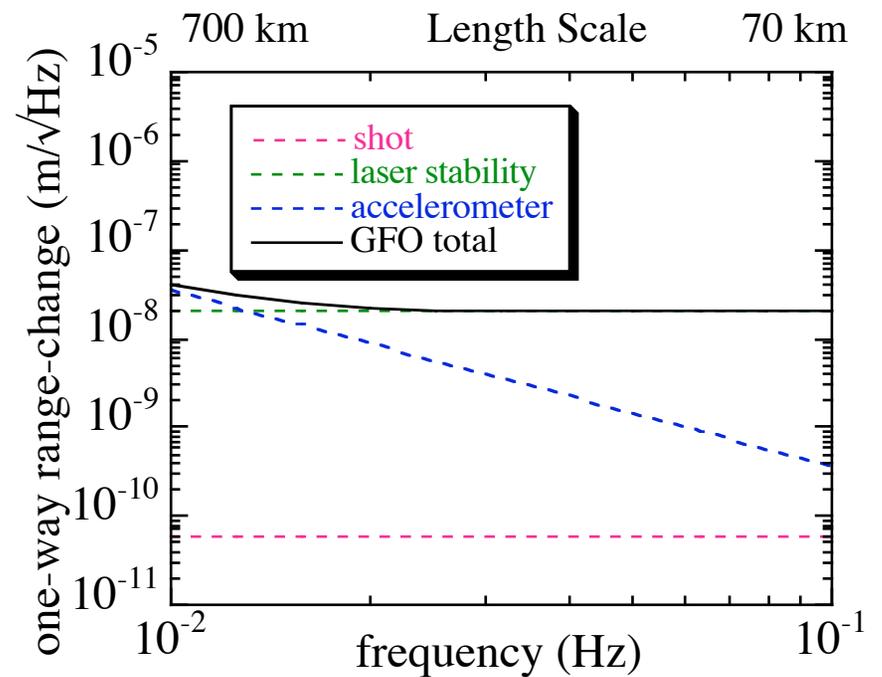


Microwave and Laser Ranging Sensitivity

Microwave Ranging System Noise



Laser Ranging System Noise





Optical Cavity Frequency Reference

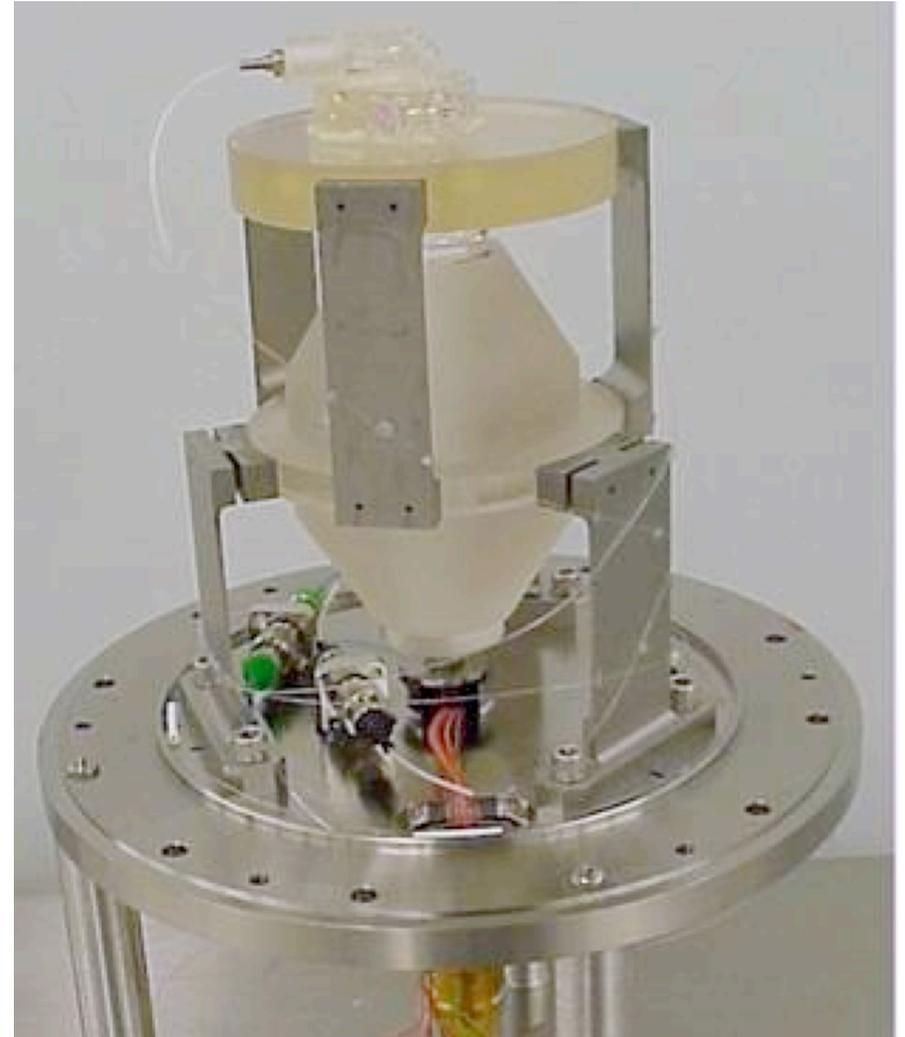
- Stabilize laser frequency by locking wavelength to thermally stable optical cavity
 - ULE glass has ultra-low thermal-expansion coefficient
 - Isolate from external temperature fluctuations





Cavity Mount Design

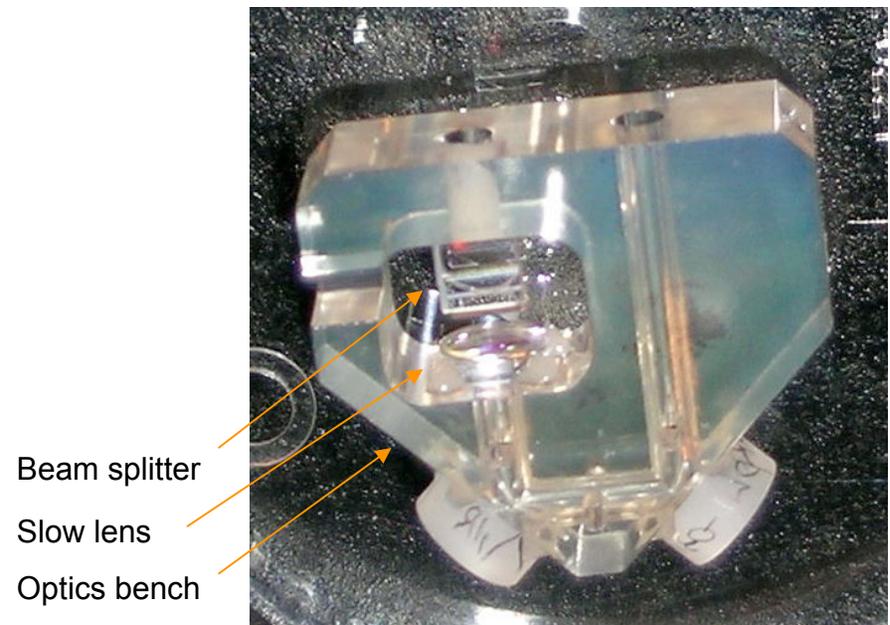
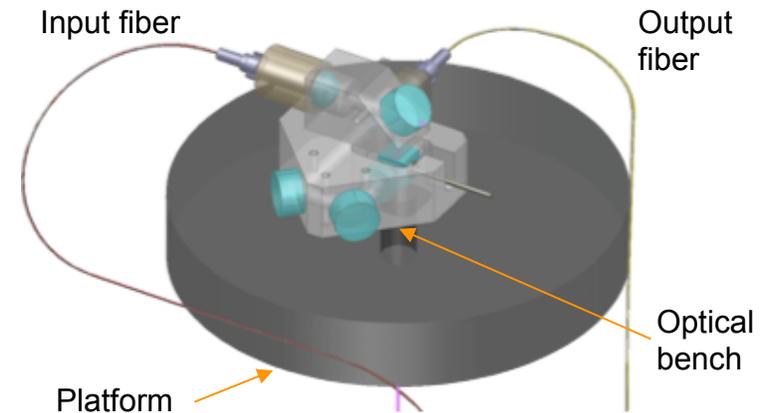
- Cavity is mounted using flexures bonded to cavity central ring
 - Flexure material (Ti) and stiffness chosen to provide support and maintain alignment for launch and minimize thermo-elastic effects on cavity
- Optics for injecting laser light into cavity and output interference via fiber also mounted to cavity using flexures





Optical Fiber Injection

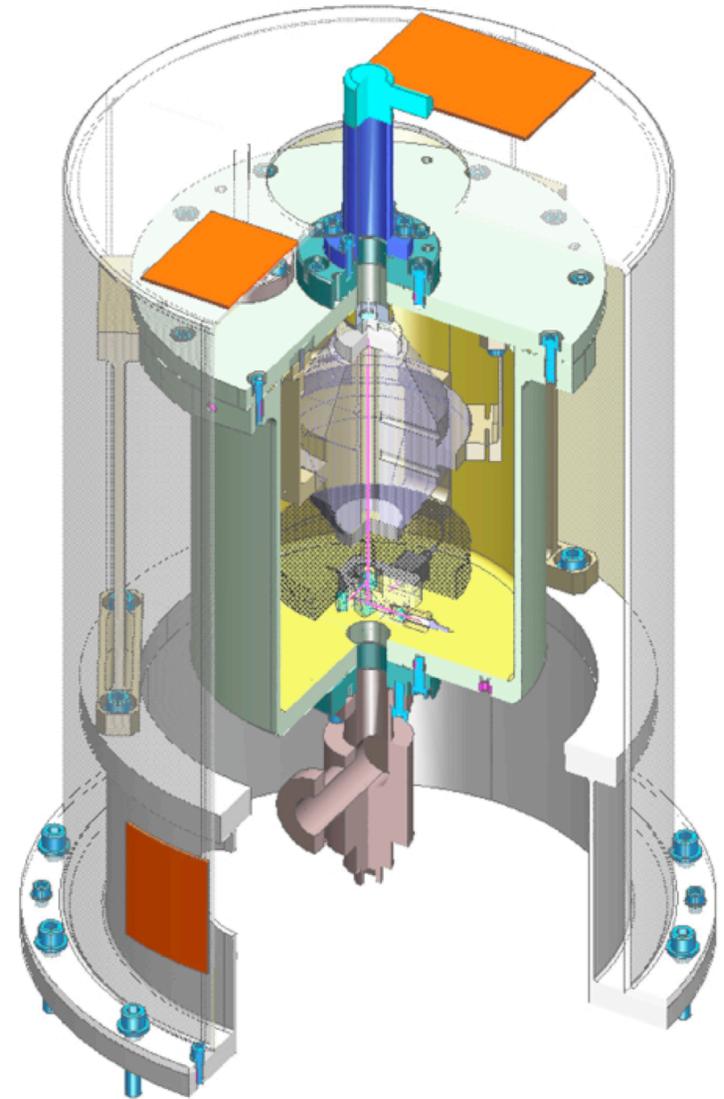
- Laser light injected into, and output signal sent, via fiber
 - Optics for light injection form small, folded telescope
- Optics, optical bench, and platform are made of low thermal expansion (zerodur) glass and mounted near cavity using flexures





Thermal Isolation Enclosure

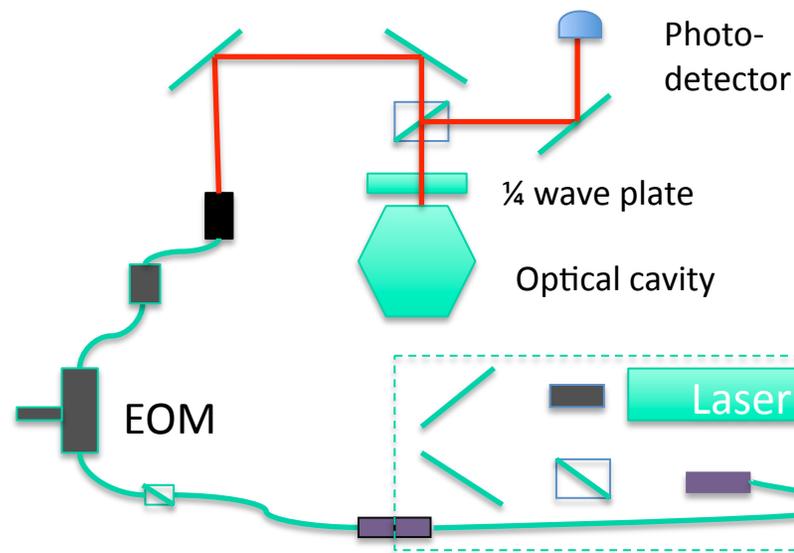
- Two concentric shells surround the cavity and optics to isolate them from external thermal
- Size and mass of laser frequency stabilization subsystem has been iterated with GRACE spacecraft team to ensure compatibility
- Inner shell forms a vacuum enclosure to avoid gas pressure fluctuations and contamination
 - Vacuum gauge will allow monitoring
 - Vacuum valve included on breadboard for development, will be replaced with pinch-off for prototype





Laser Locking to Optical Cavity

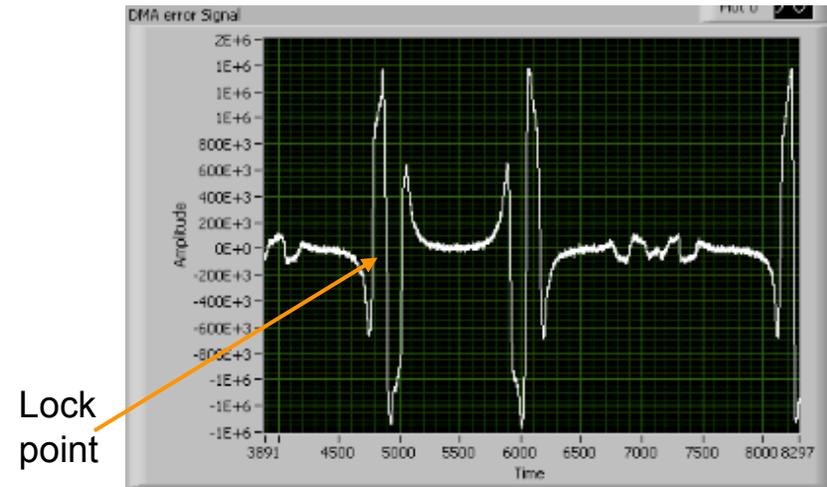
- Laser locked to cavity using Pound-Drever-Hall technique
 - Light resonance occurs when cavity length is integer number of wavelengths
 - Laser frequency stability is directly related to stability of length of cavity
 - Cavity is made of material with low thermal expansion coefficient
 - Light exiting cavity compared with light entering cavity
 - Requires electro-optical modulator to add phase-modulation to laser beam
 - Difference in modulation signal gives laser correction to match cavity



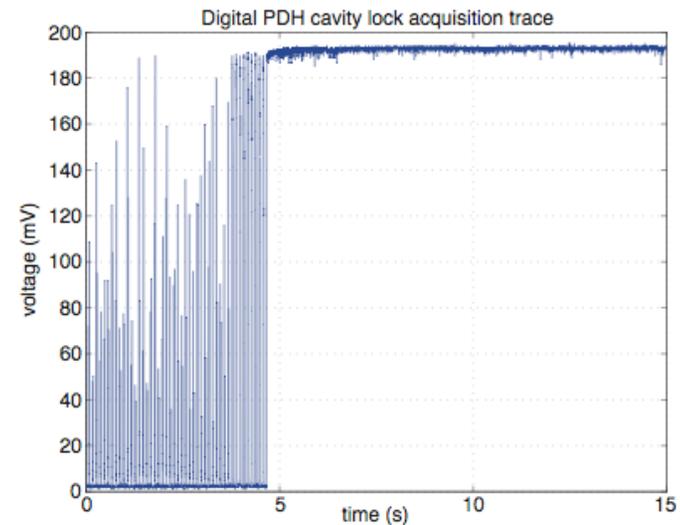


Cavity Response Function

- Laser light injected into cavity exits and interferes with input light
 - Pound-Drever-Hall locking
- As laser frequency varies at input to cavity, interference with exiting light produces sharply varying response
- When error signal from locking point fed back to laser, laser frequency is stabilized



Interference response as laser frequency changed

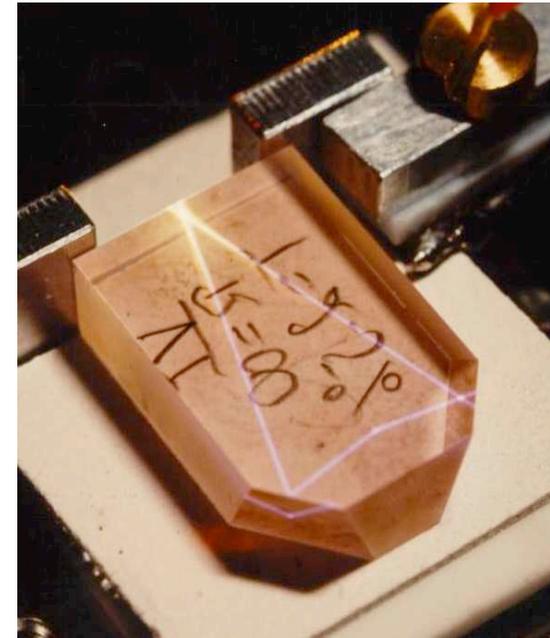


Laser frequency swept then locked



Tunable Laser

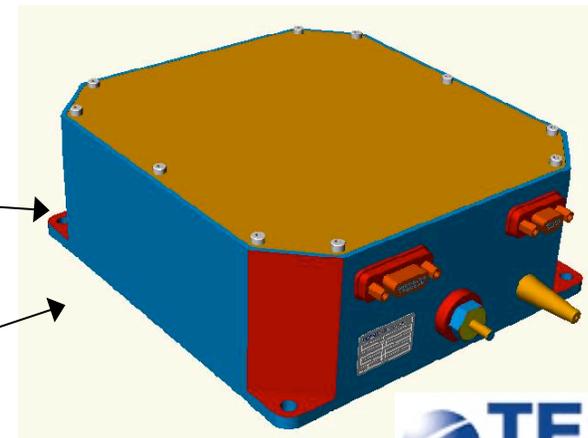
- Non-planar ring oscillator (NPRO) Nd:YAG laser provides tunability for locking to cavity
 - Laser wavelength adjusted by changing dimensions of YAG crystal using PZT glued to crystal and thermal adjustment
- Space-qualified NPRO laser available from Tesat Spacecom



NPRO laser head



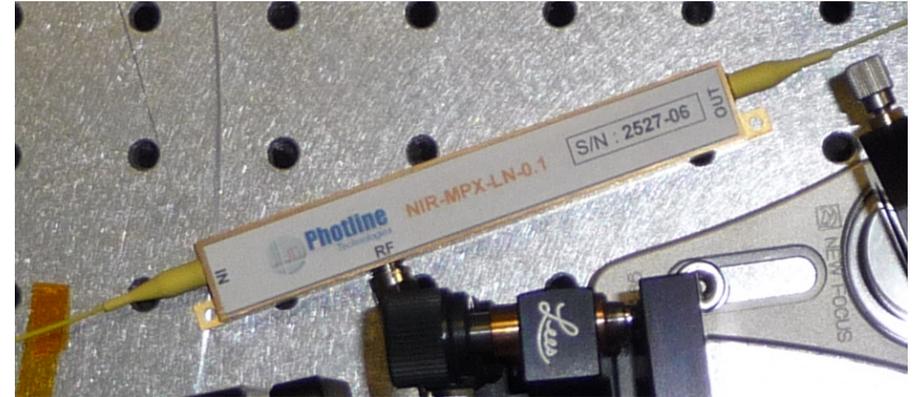
Laser pump diode assembly





Electro-Optical Modulator

- Several companies make EOM suitable for use in space
- Selected unit from Photline has inexpensive laboratory model traceable to flight model
 - Tested to show meets locking requirements

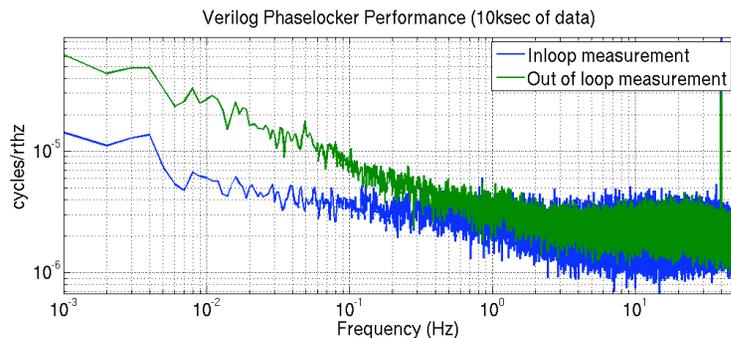
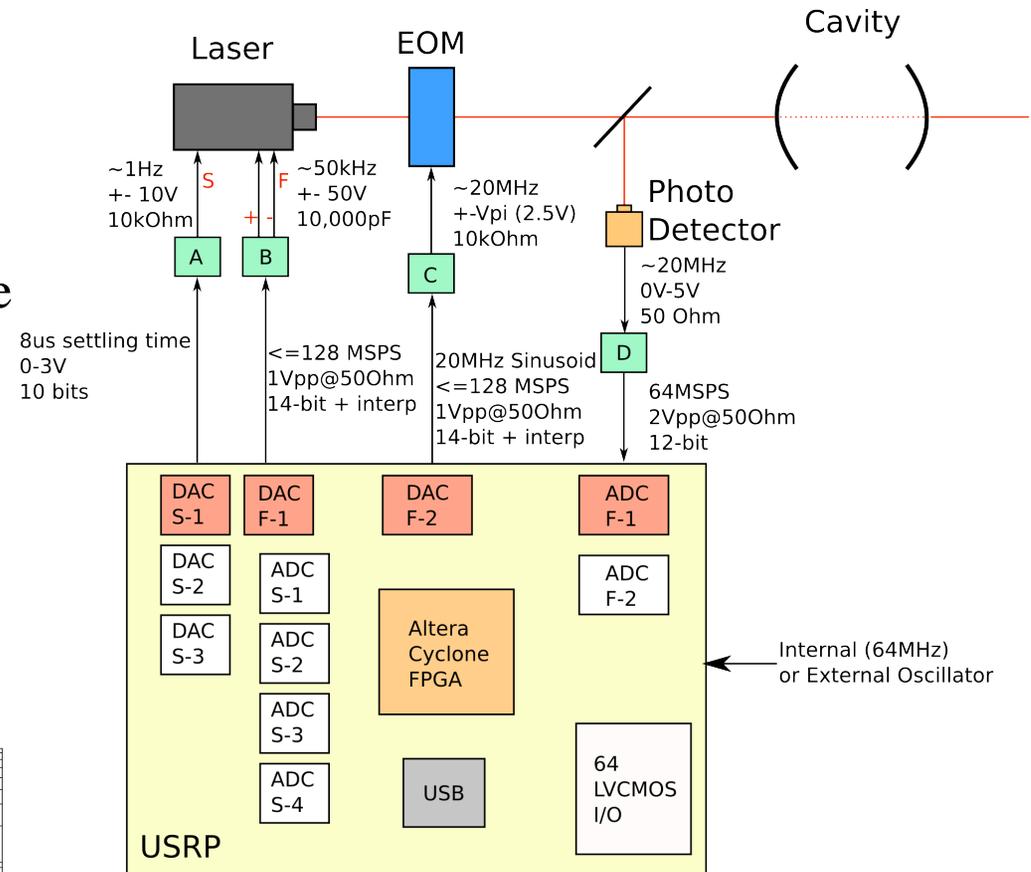


Selected electro-optical modulator



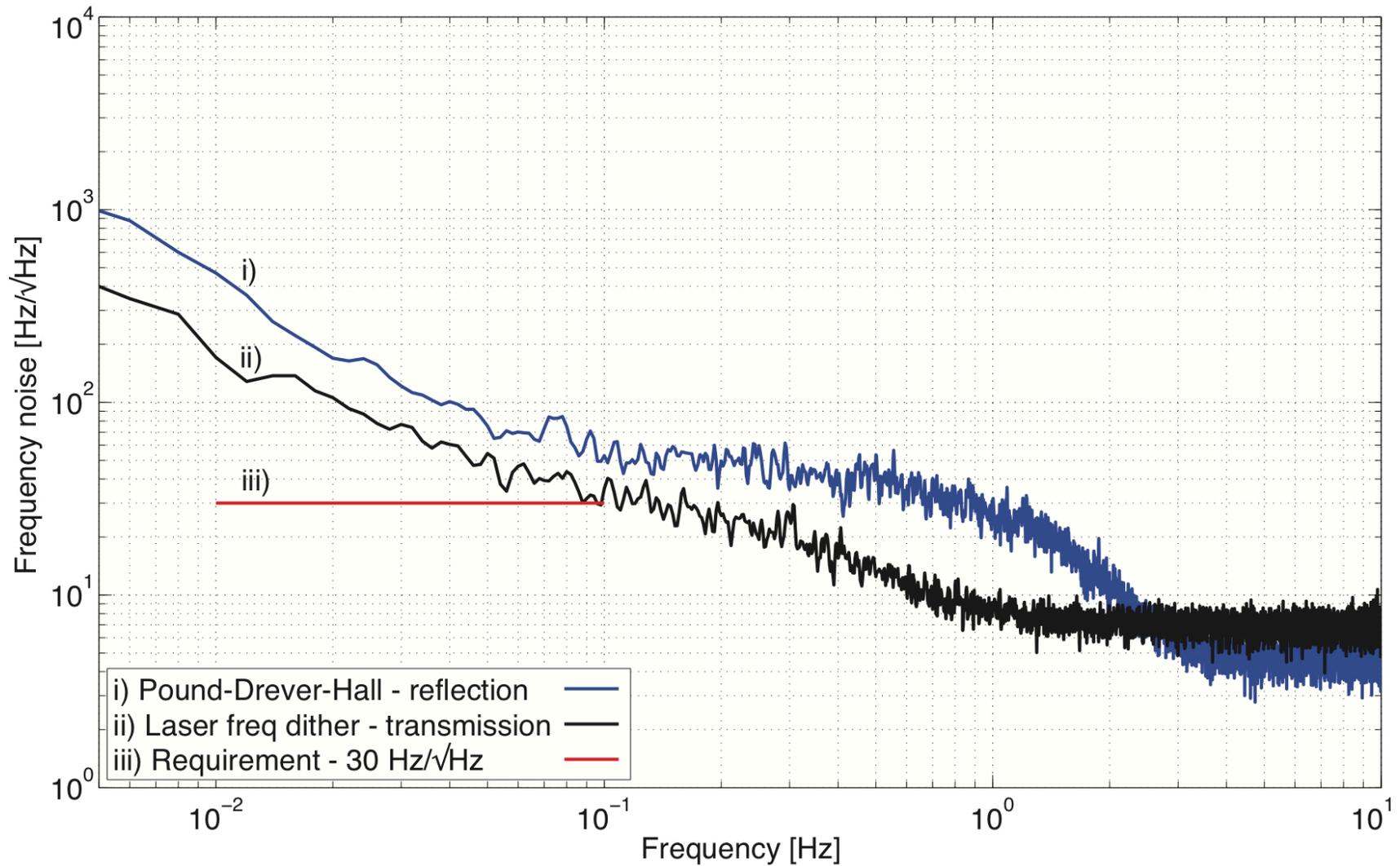
Breadboard Locking Electronics

- Commercial FPGA evaluation board used for development and testing of laser locking algorithms
 - ADC for sampling interference signal from photodiode
 - DACs used to adjust laser crystal frequency via PZT and temperature





Subsystem Performance Results



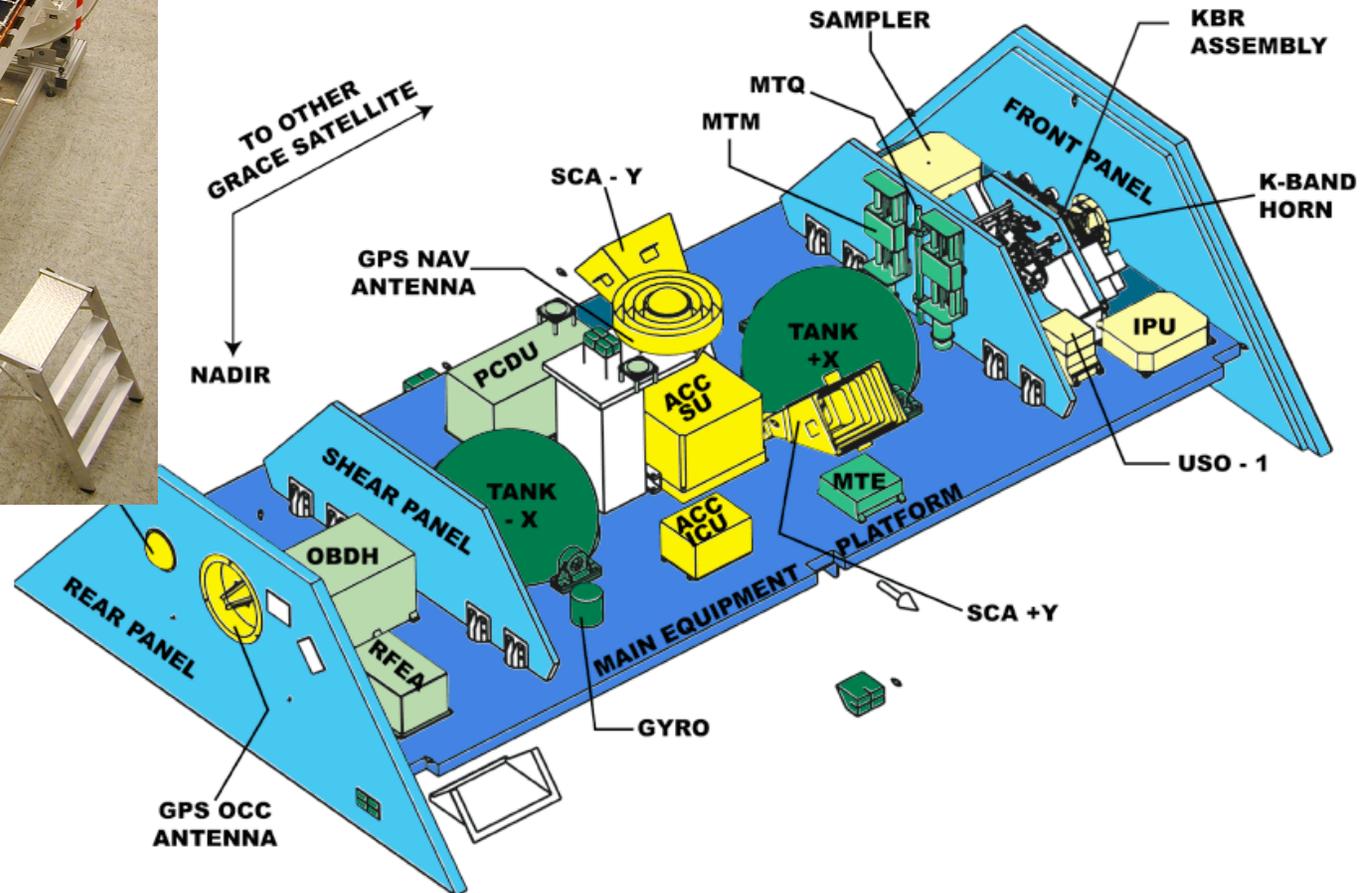
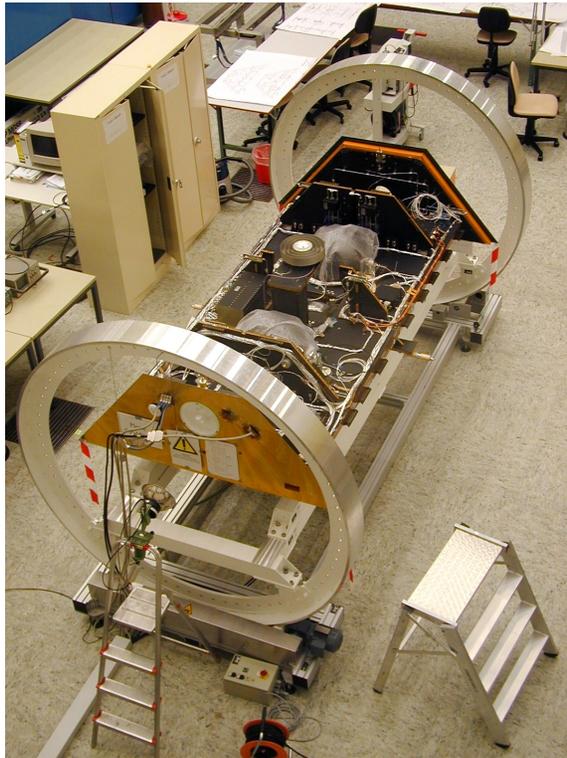


Environmental Test Results

- Breadboard cavity subjected to vibration and thermal tests per the GRACE mission requirements
- Vibration tests were successful
- Thermal tests showed a slight misalignment occurred at highest temperatures (+60 C, +20 C above maximum flight temp.)
 - Traced to adhesive glass transition at 50C
 - Prior uses of same adhesive had not been as sensitive to alignment
 - We are investigating options for prototype unit

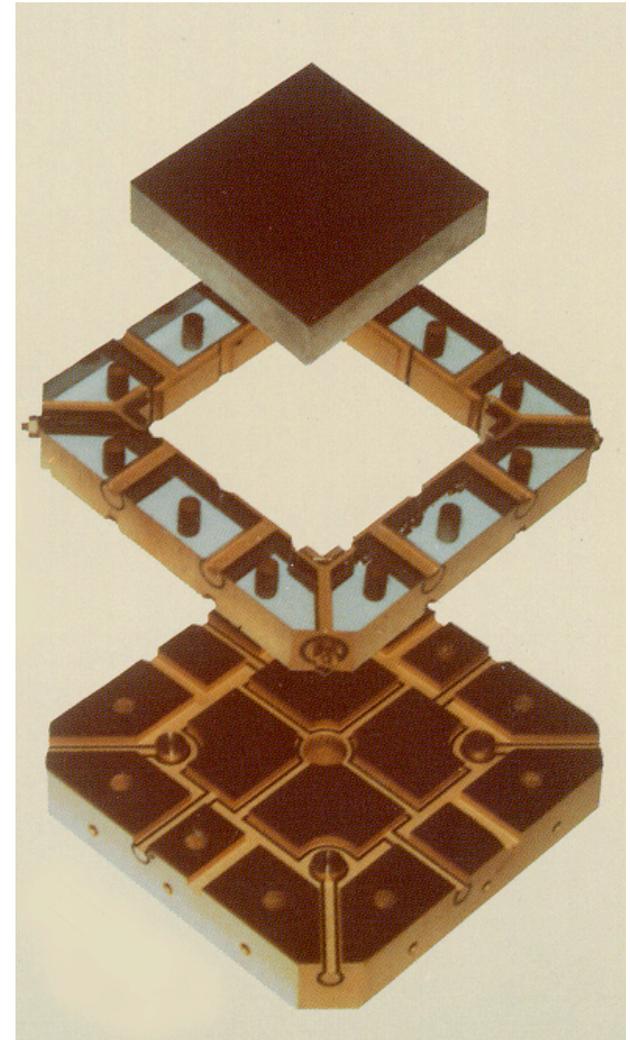
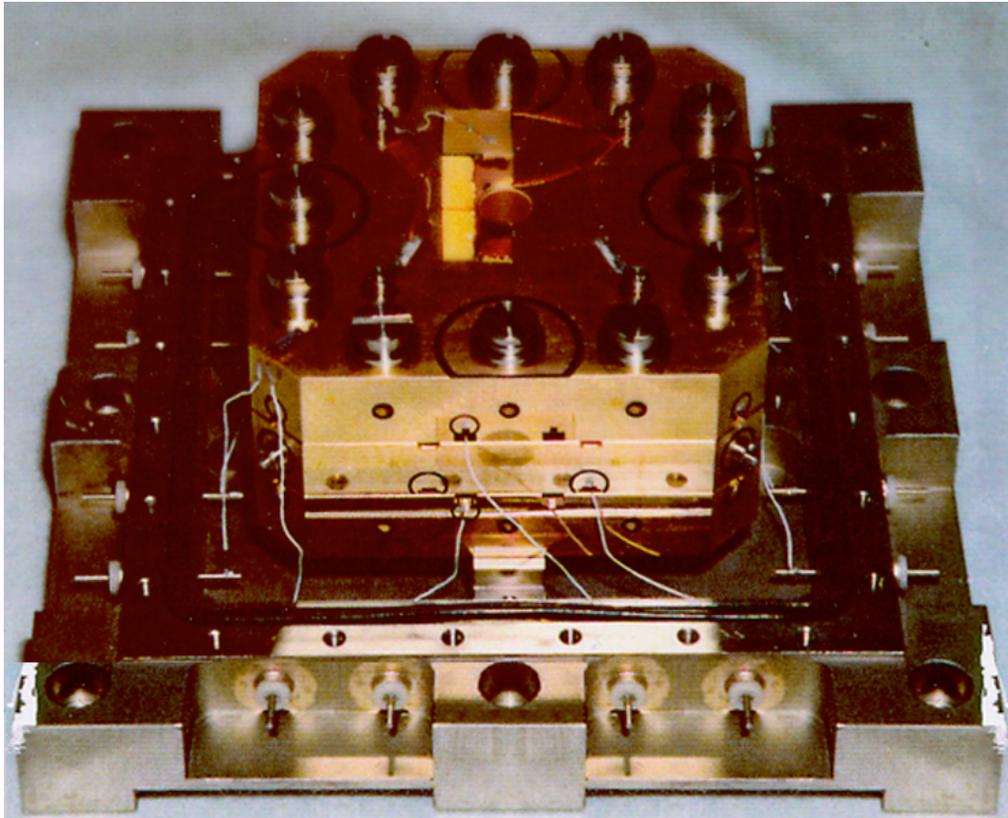


GRACE Spacecraft





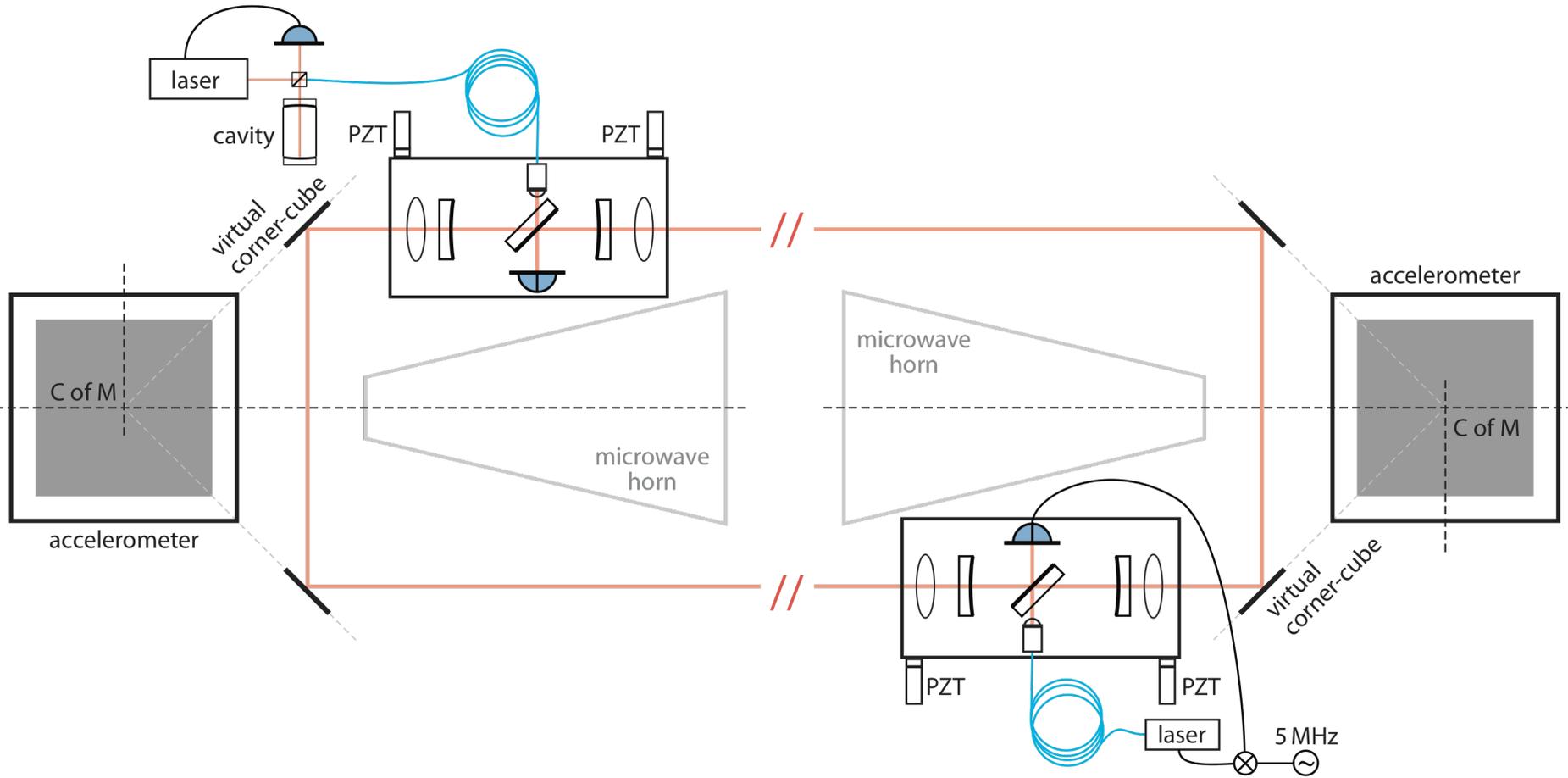
SuperSTAR Accelerometer



ONERA

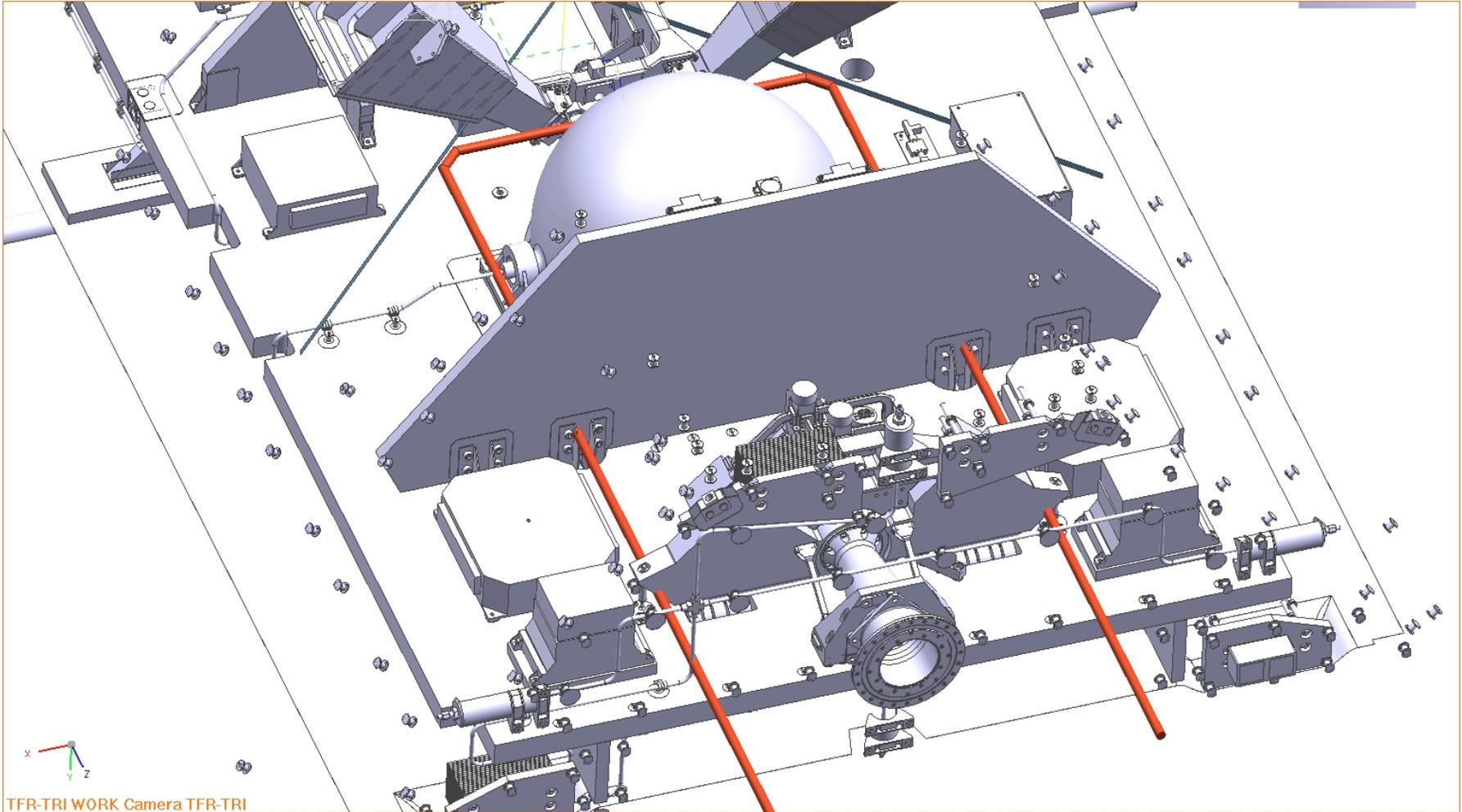


Racetrack Configuration



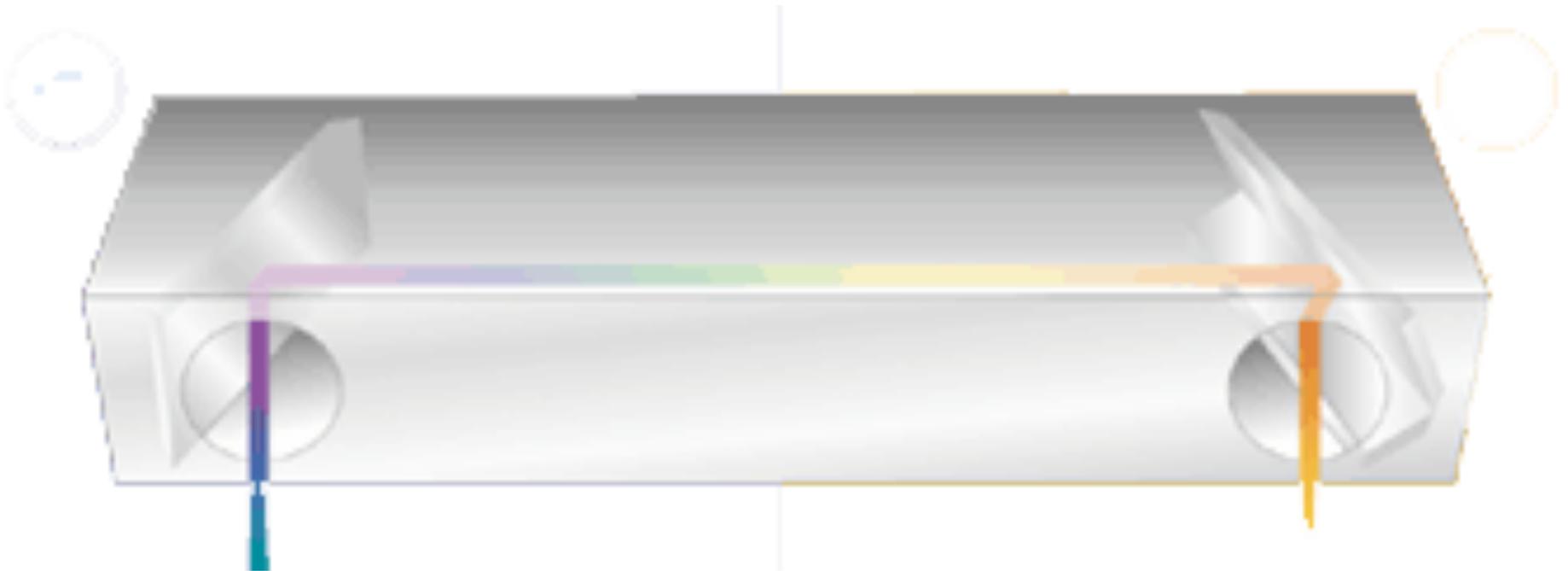


Laser Optical Path on GRACE-FO





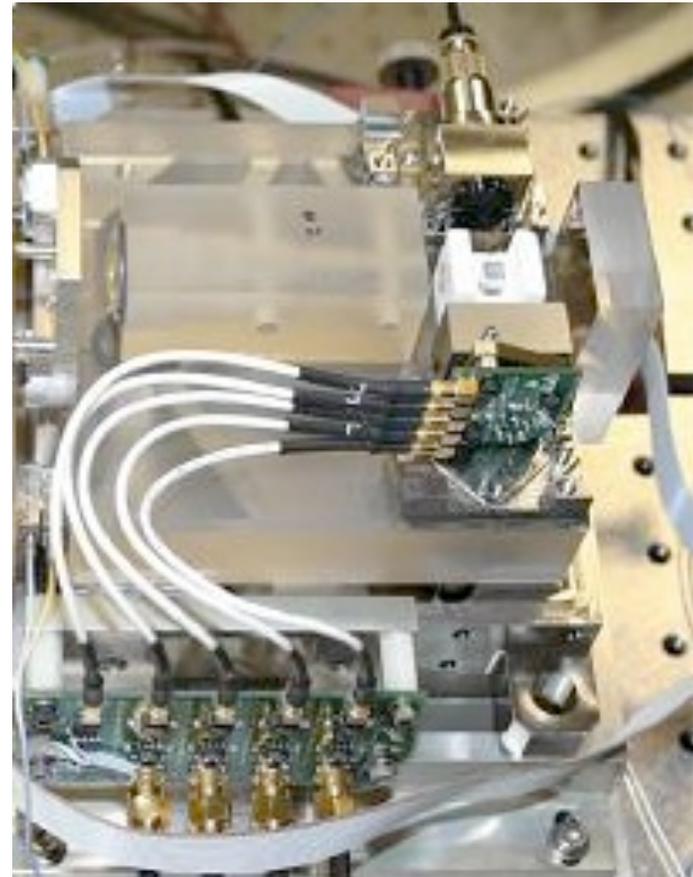
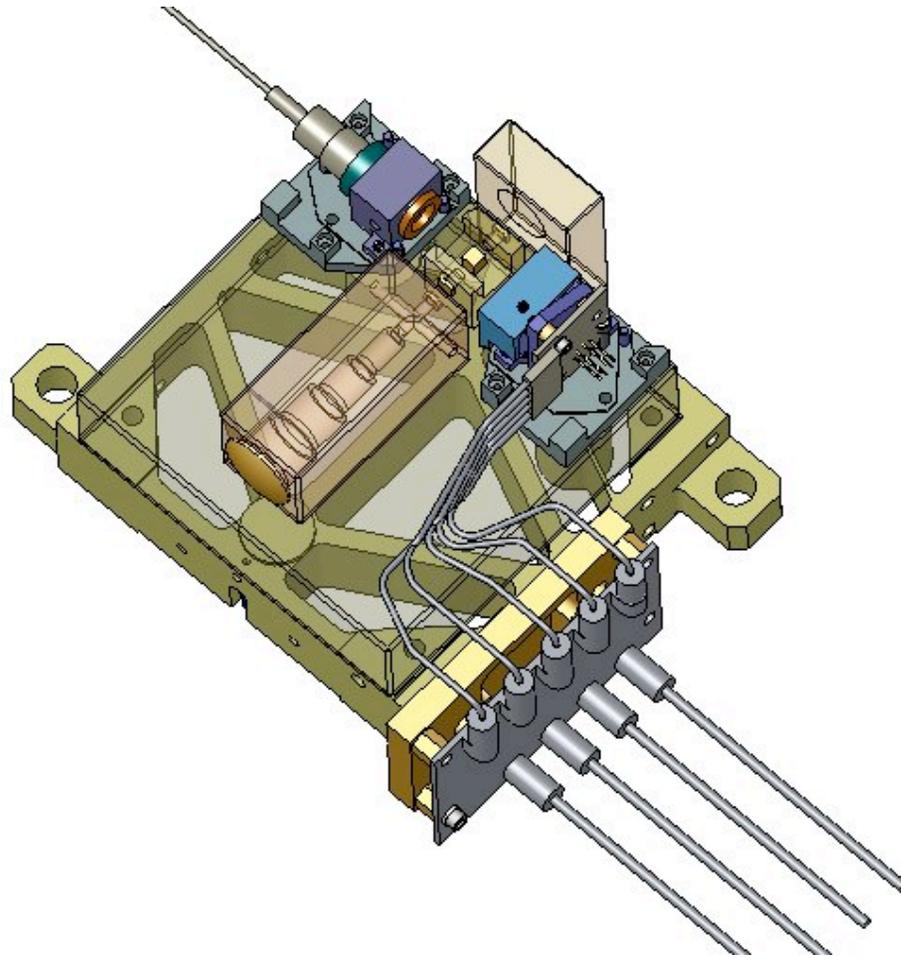
Lateral Transfer Hollow Retroreflector™



PLX



IRT Optical Bench

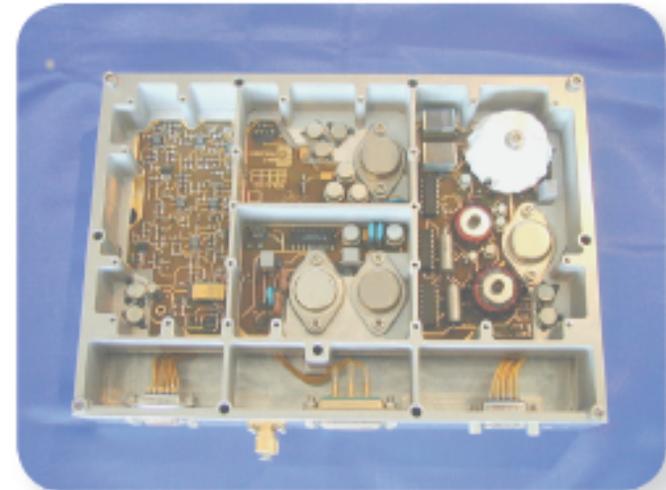




Candidate PZT Actuators



View of the APA μ XS ($m = 0.15\text{gr}$).



FB-LA75-space card.



www.cedrat.com



Future Plans

- Currently testing performance of modified breadboard subsystem to see if limiting error sources have been identified and corrected
- Prototypes of cavity assembly and electronics board are to be built and tested in FY '11 to demonstrate TRL 6.
- GRACE-FO scheduled to start in FY '11 with PDR in FY'12
 - Cavity and locking electronics development fits GRACE-FO schedule well
 - Will be challenging to bring total ranging system to flight readiness for launch in 2016.